



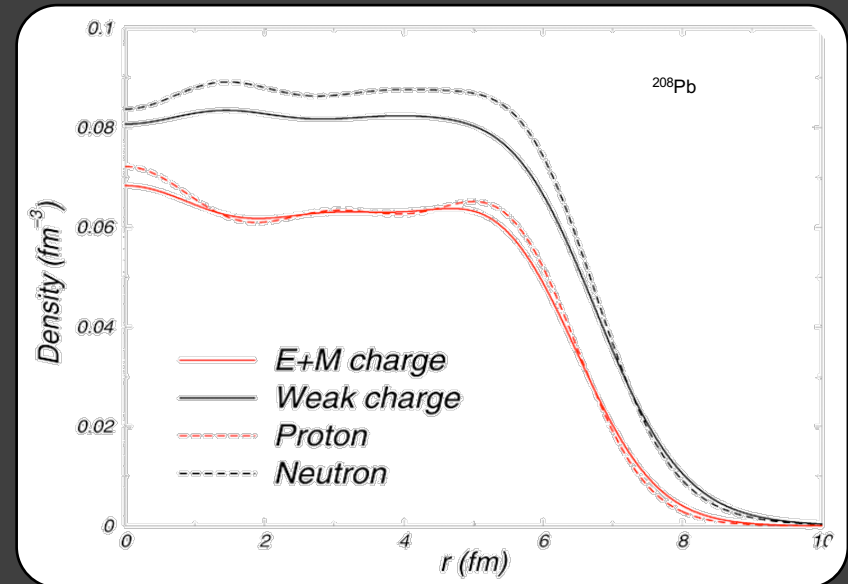
# PREX, CREX and Short-Range Correlations

Or Hen

# PREX-2 and CREX : Neutron Skin in $^{208}\text{Pb}$ and $^{48}\text{Ca}$

Nuclear theory predicts a neutron  
“skin” on heavy nuclei

Very clean interpretation makes  
this a crucial calibration for  
equation-of-state of heavy,  
neutron rich nuclei



At the measured  $Q^2$ ,  $F_w$  is primarily  
sensitive to the density dependence of  
the nuclear equation of state

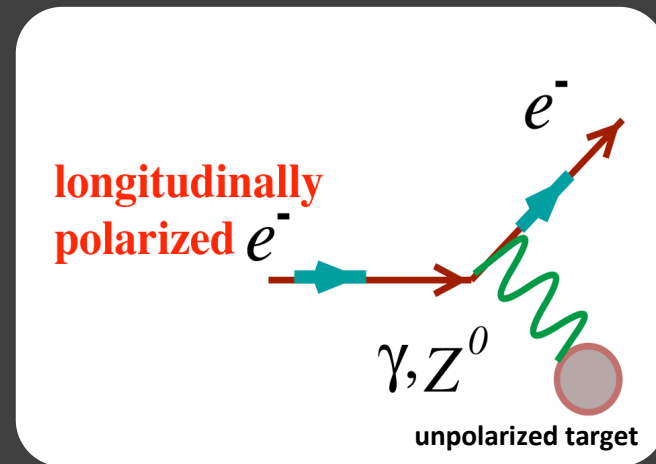
# PREX-2 and CREX : Neutron Skin in $^{208}\text{Pb}$ and $^{48}\text{Ca}$

Electron Scattering:  $\gamma$  vs  $Z^0$

$$\frac{d\sigma}{d\theta dE} \Leftrightarrow F_{\text{ch}}(Q^2)$$

$$A_{\text{PV}} \Leftrightarrow F_{\text{W}}(Q^2)$$

parity-violating  
asymmetry

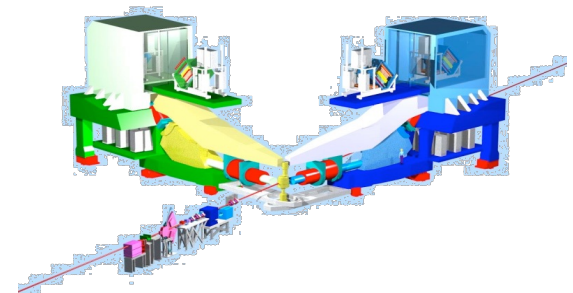
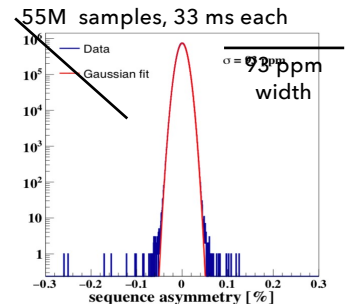
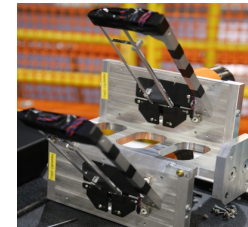


$$A_{\text{PV}} \approx \frac{G_{\text{F}} Q^2}{4\pi\alpha\sqrt{2}} \frac{Q_{\text{W}} F_{\text{W}}(Q^2)}{Z F_{\text{ch}}(Q^2)}$$

$F_{\text{W}} - F_{\text{ch}}$  : negligible model dependence  
+ well-known  $R_{\text{ch}}$  and  $\rho$  shape model  
 $\rightarrow R_{\text{W}}, R_{\text{n}}, R_{\text{n}} - R_{\text{p}}, \rho_0^{\text{W}}$

# Challenging Measurement!

- High luminosity, >4 GHz detected rate (PREX-2)
- Precise calibration & polarimetry
- Careful control of beam asymmetries, backgrounds, noise, false ...asymmetries



PRL 126, 172502 (2021)

PRL 129, 042501 (2022)

# $^{208}\text{Pb}$ and $^{48}\text{Ca}$ Results

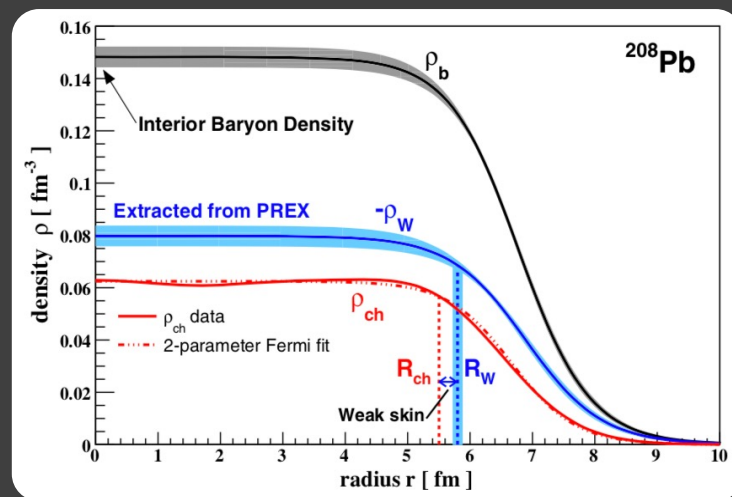
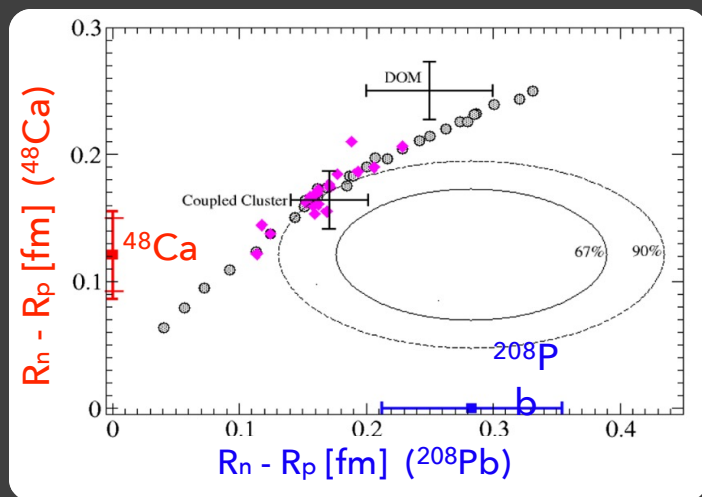
PREX-2:  $^{208}\text{Pb}$   $A_{\text{PV}} = 550 \pm 16$  (stat)  $\pm 8$  (syst) ppb

CREX:  $^{48}\text{Ca}$   $A_{\text{PV}} = 2668 \pm 106$  (stat)  $\pm 40$  (syst) ppb

Determine neutron skins with small model dependence:

PREX:  $^{208}\text{Pb}$   $R_n - R_p = \underline{0.278 \text{ fm}} \pm 0.071$  (exp)  $\pm 0.012$  (model)

CREX:  $^{48}\text{Ca}$   $R_n - R_p = \underline{0.121 \text{ fm}} \pm 0.026$  (exp)  $\pm 0.024$  (model)

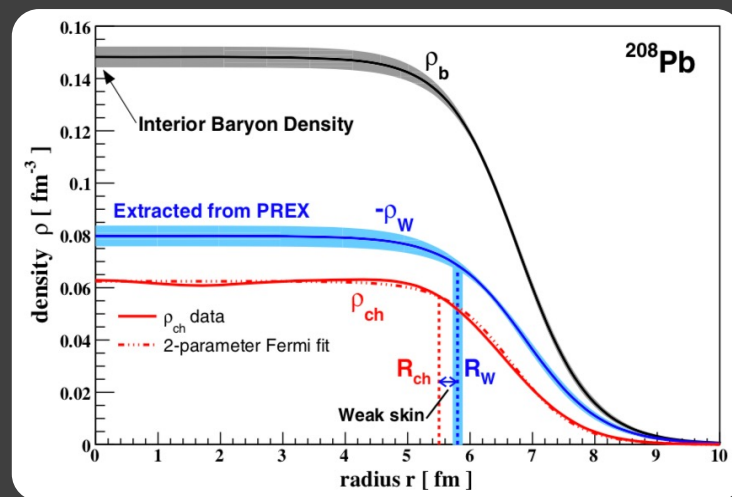
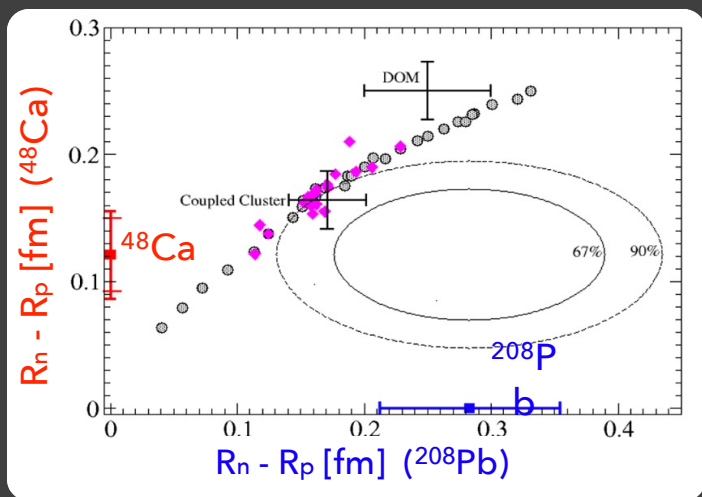


# $^{208}\text{Pb}$ and $^{48}\text{Ca}$ Results

PREX indicate a large skin, with low nuclear baryon density  $\rho_b$   
CREX indicate a smaller skin than most models

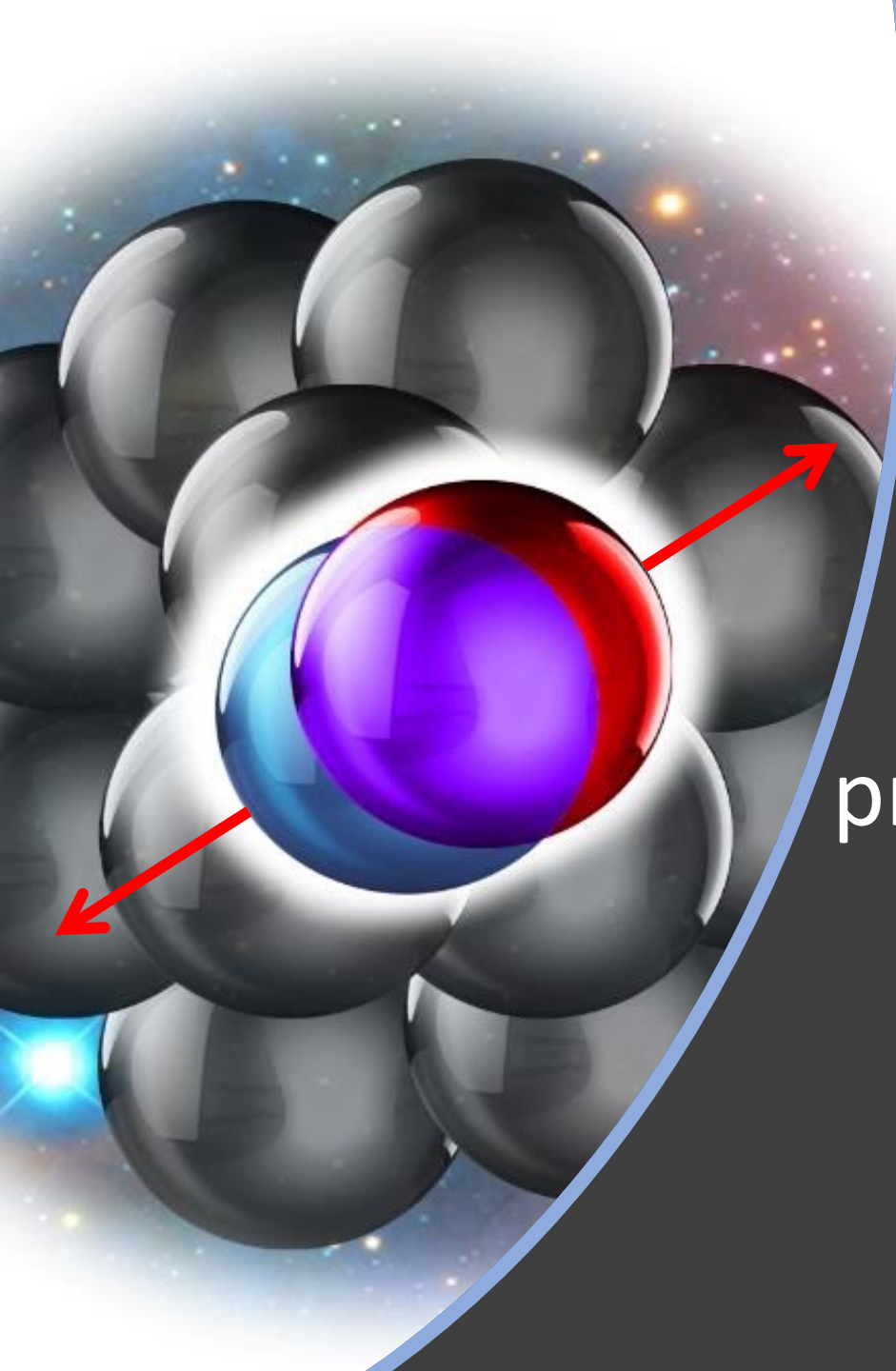
Generally, models expect both skins large or both small;  
The results presents an important empirical test

Unique, valuable contribution to the nuclear structure data set, as a  
cleanly interpretable measurement of an isovector nuclear property

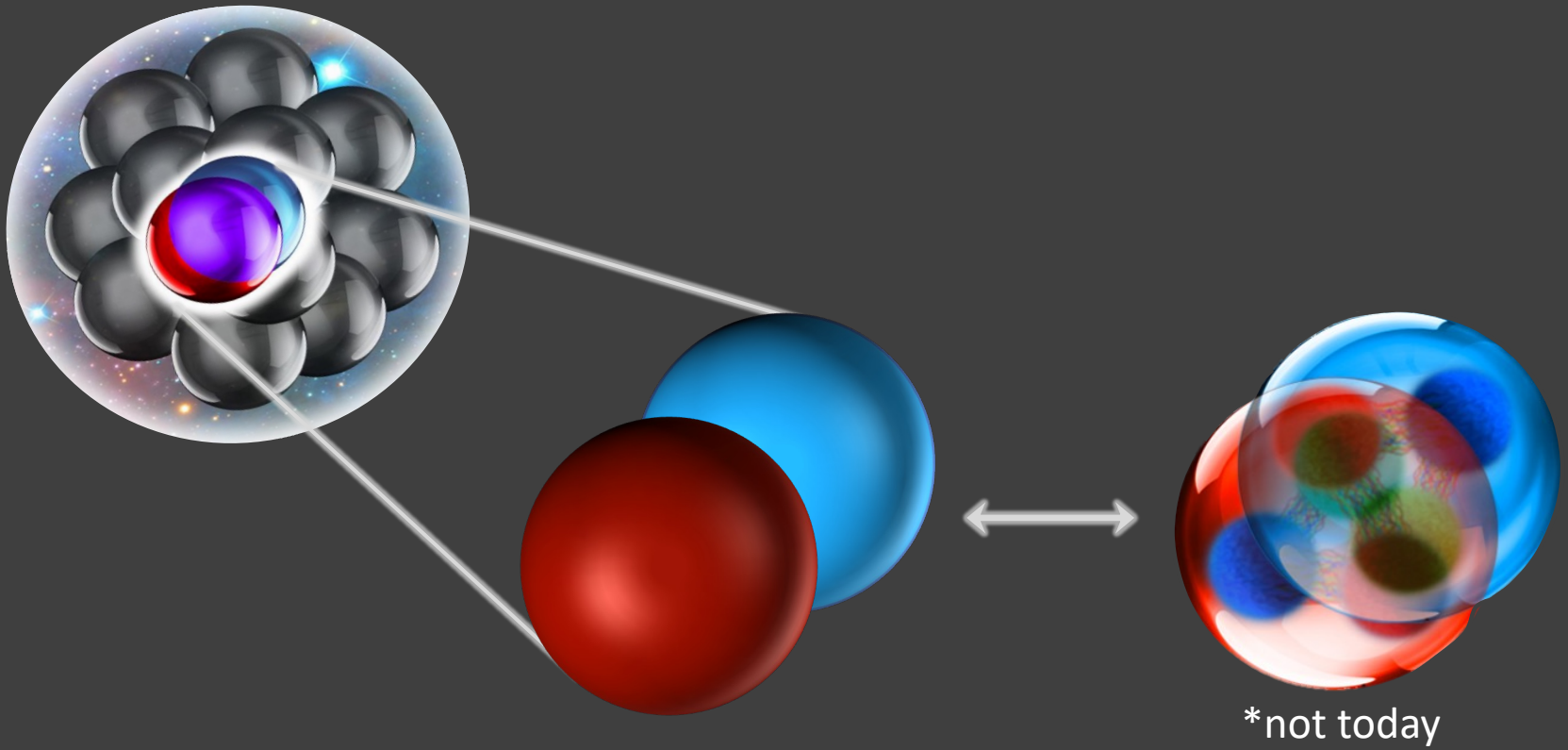


# Short-Range Correlations (SRC)

Fluctuations of close-  
proximity nucleon pairs

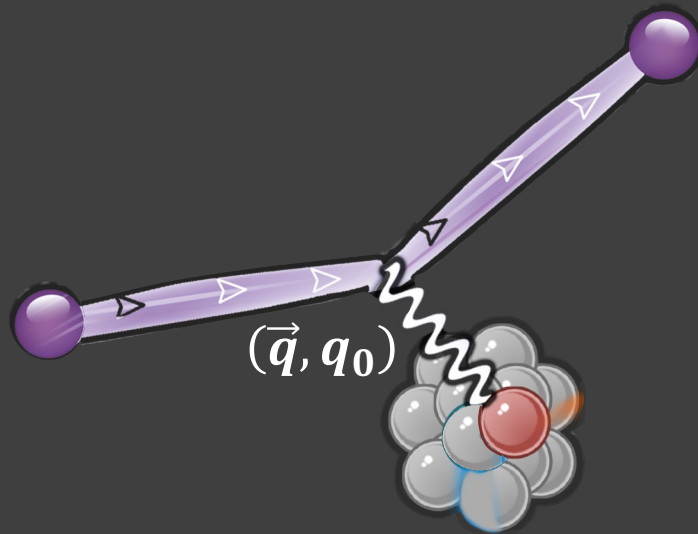


# SRCs Across Nuclear Scales





# High Res. Electromagnetic Probe



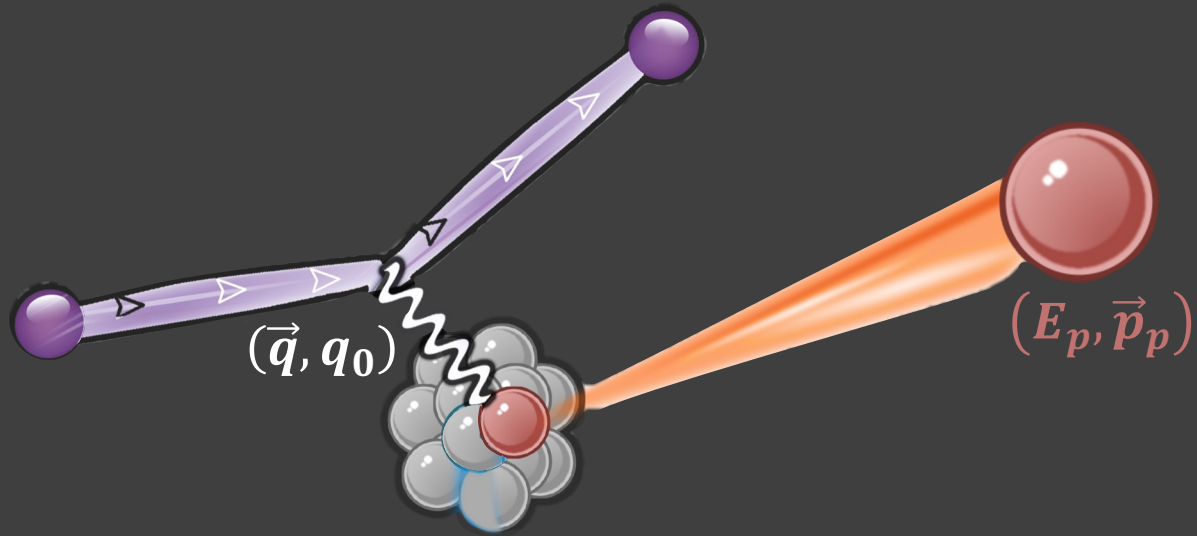
Reaction Scale:  $Q^2 = |\vec{q}|^2 - q_0^2$

Reaction scale  $\gg$  Nuclear energy scale

$$Q^2 > 1.5 - 2 \text{ GeV}^2$$

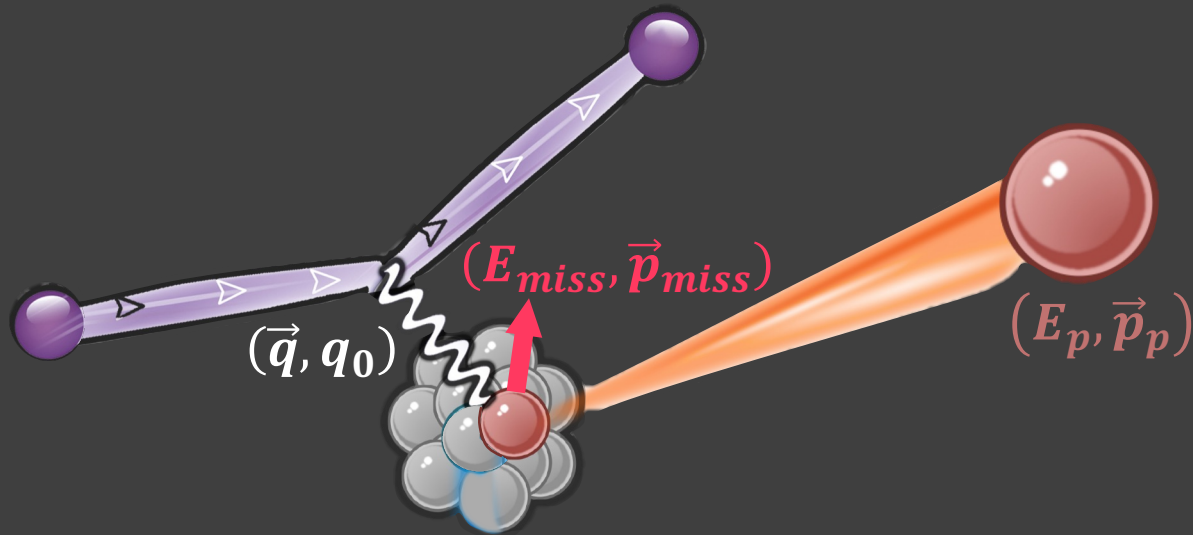
$$(q_0, \vec{q} > 1 \text{ GeV} \gg |V_{NN}|, 2 k_F)$$

# High Res. Electromagnetic Probe

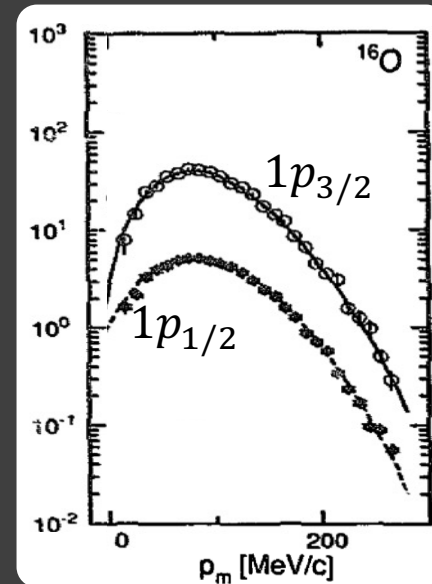


Proton carries large fraction of  $(\vec{q}, q_0)$

# High Res. Electromagnetic Probe

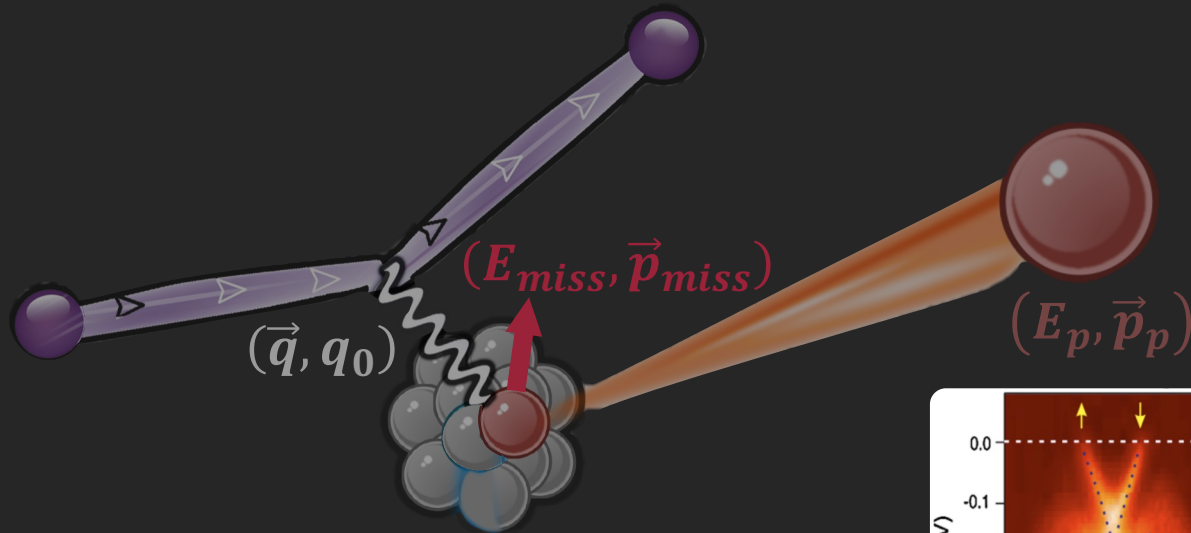


4-Momentum  $\vec{p}_{miss} = \vec{p}_p - \vec{q}$   
Conservation:  $E_{miss} = E_p - q_0$



Lapikas Nuc. Phys. A '93

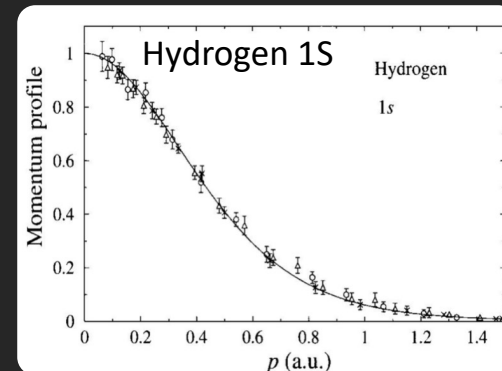
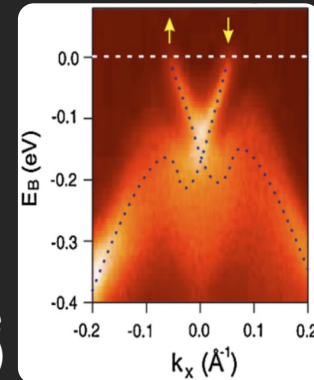
# High Res. Electromagnetic Probe



4-Momentum  $\vec{p}_{miss} = \vec{p}_p - \vec{q}$   
 Conservation:  $E_{miss} = E_p - q_0$

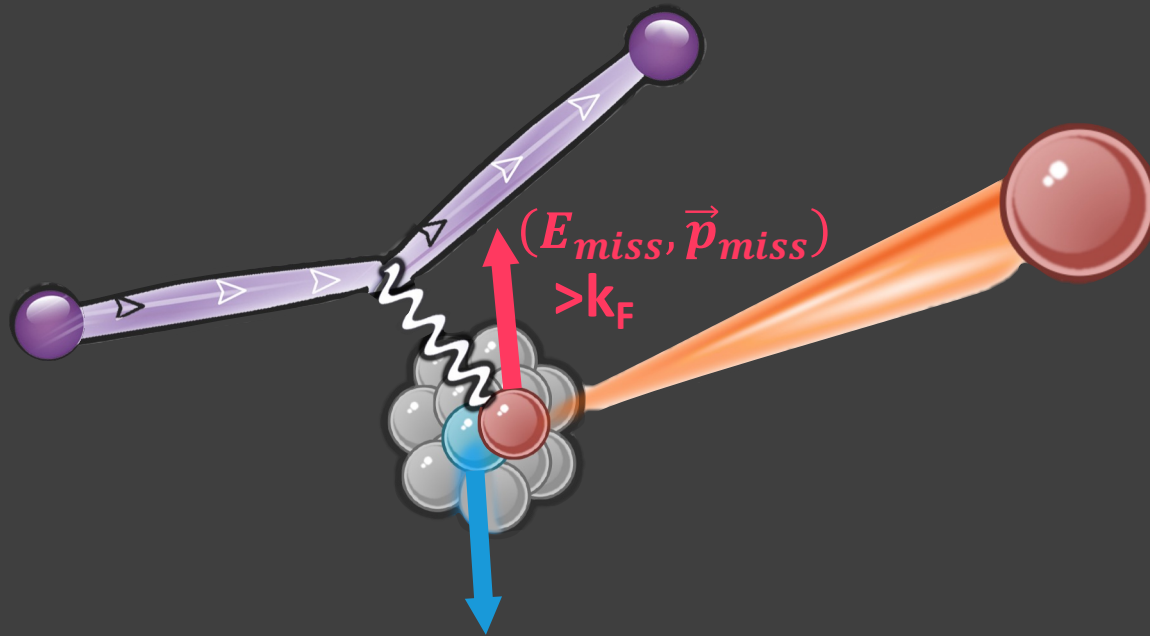
Similar to atomic / condensed matter spectral function extractions

( $\gamma, e$ ) ARPES: surface state bands (2009)

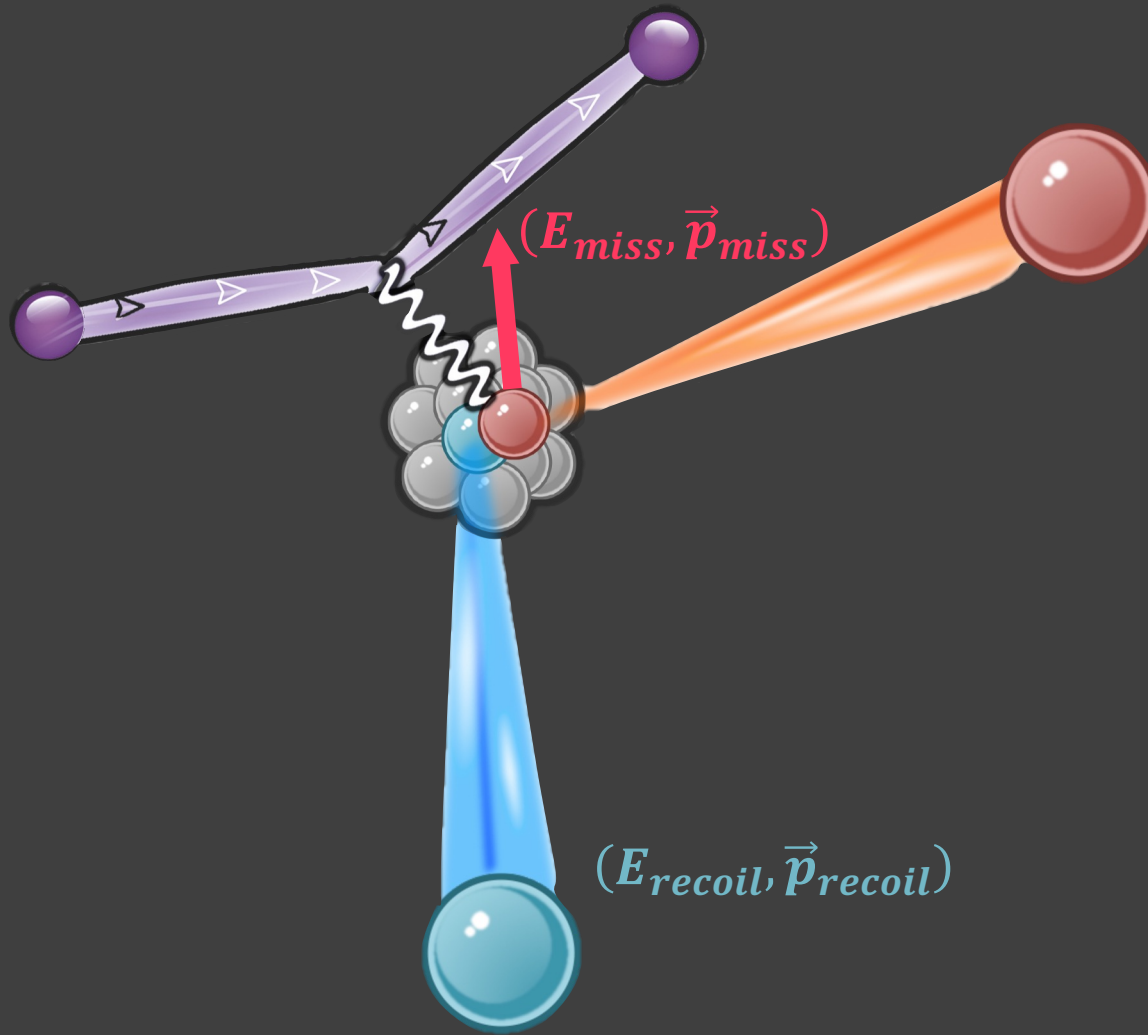


( $e, 2e$ ) reduced cross-section (1981)

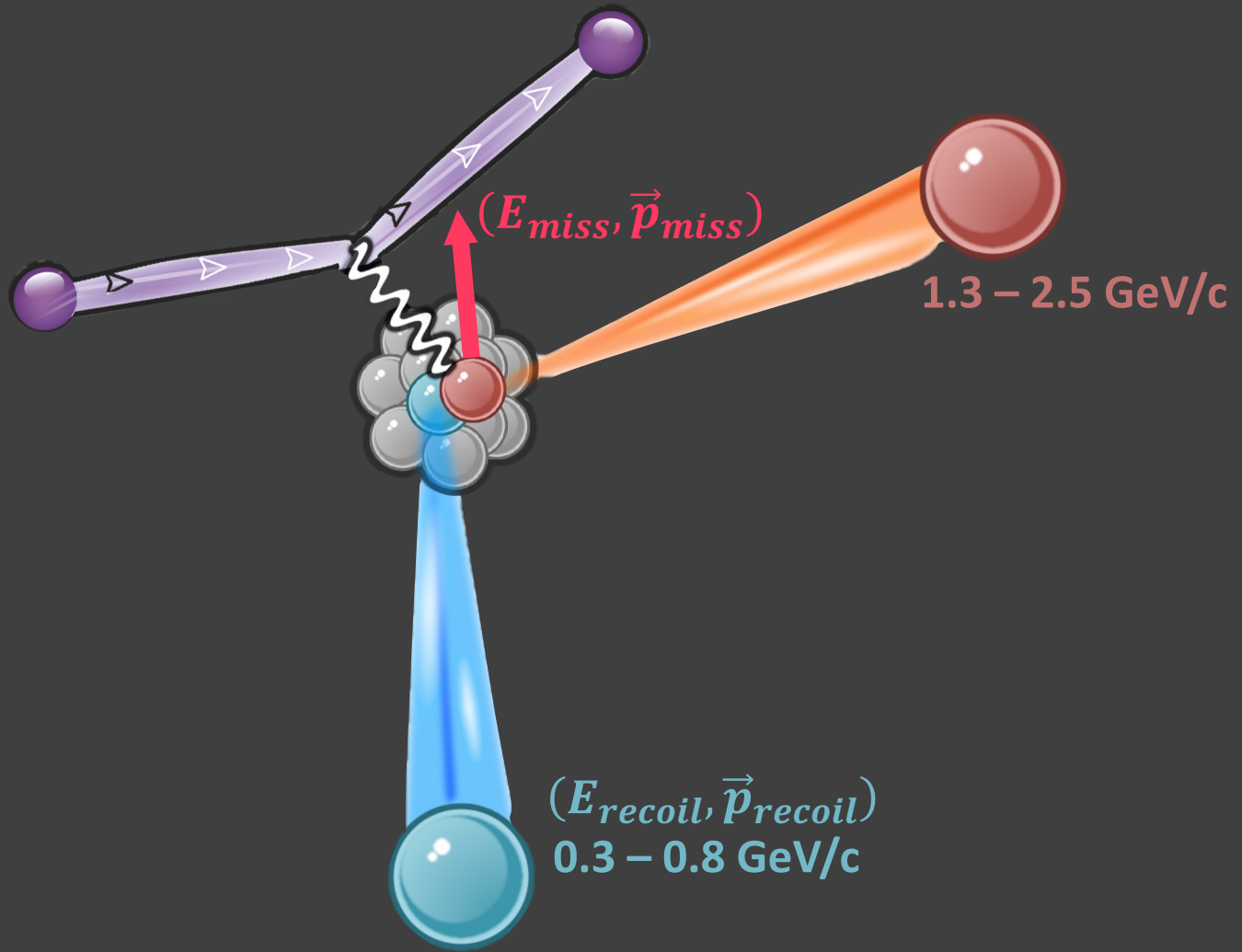
# Probing Short-Range Correlations



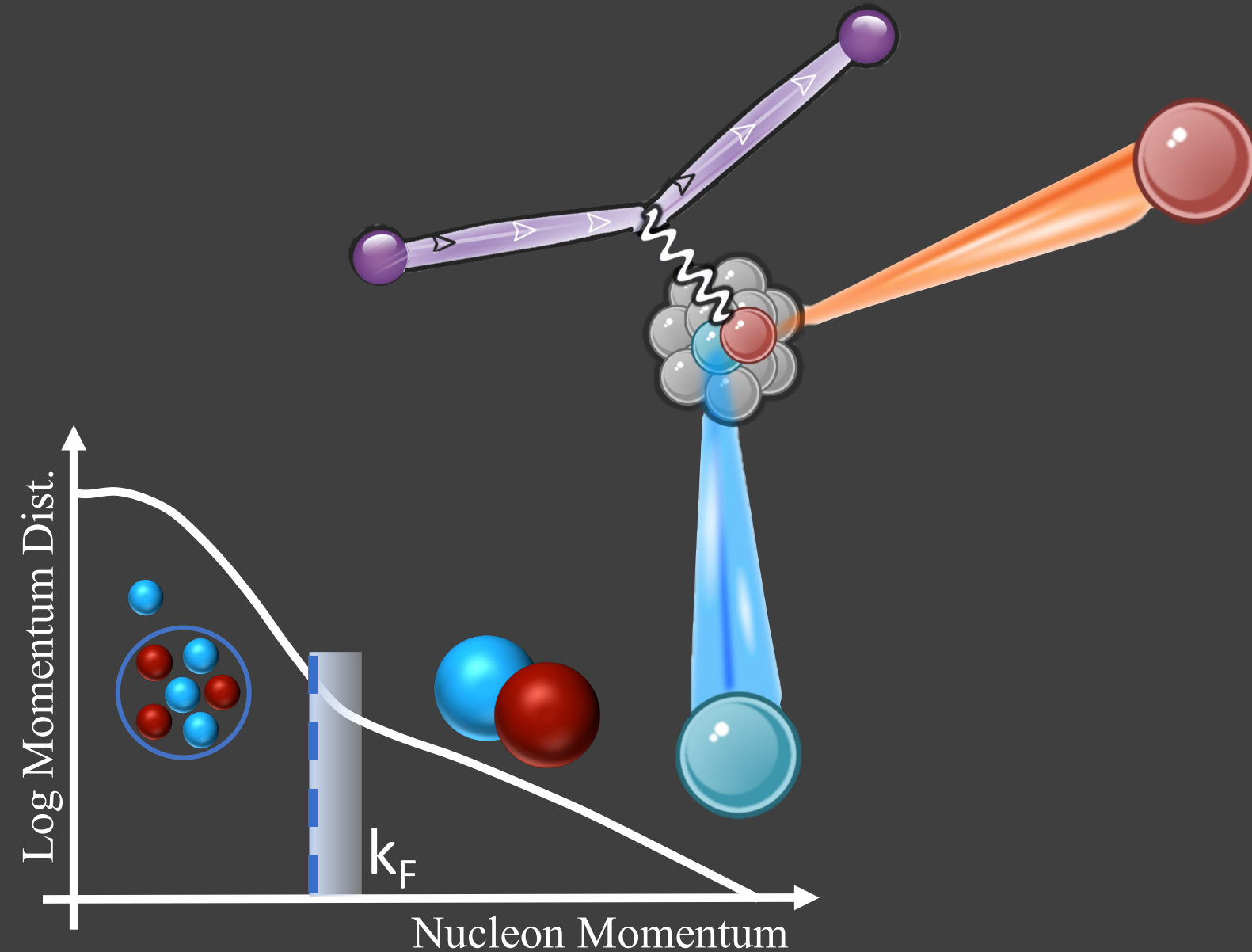
# Probing Short-Range Correlations



# Probing Short-Range Correlations



# Probing Short-Range Correlations





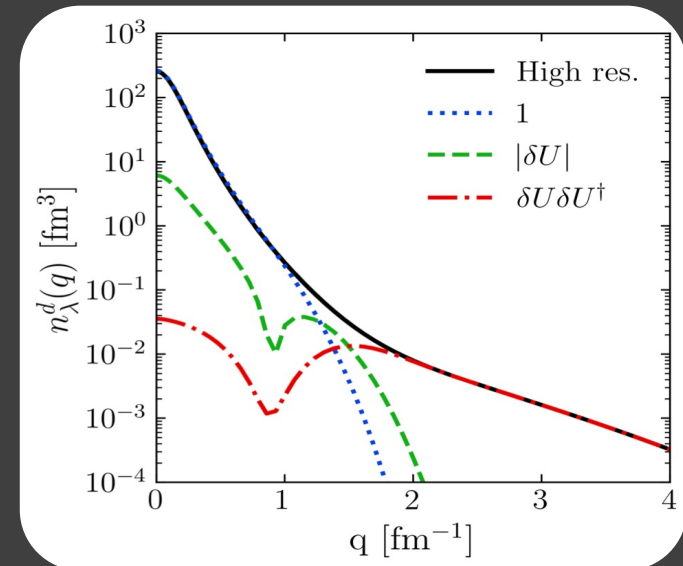
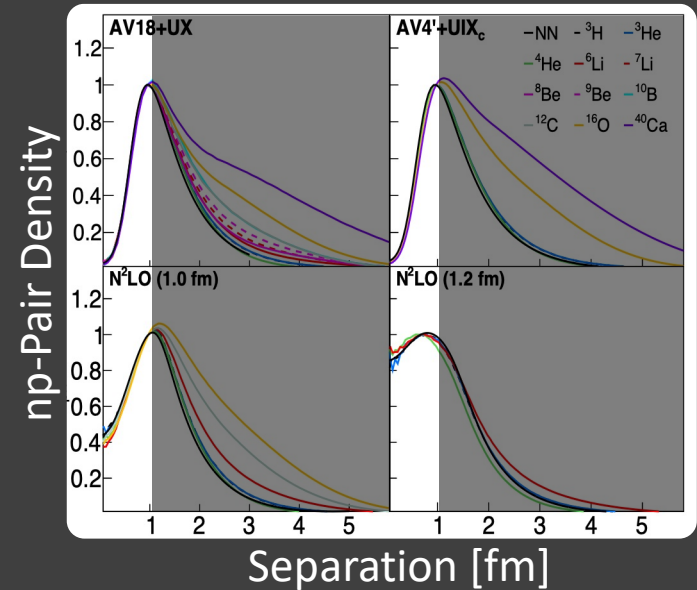
# Scale Separation and Factorization

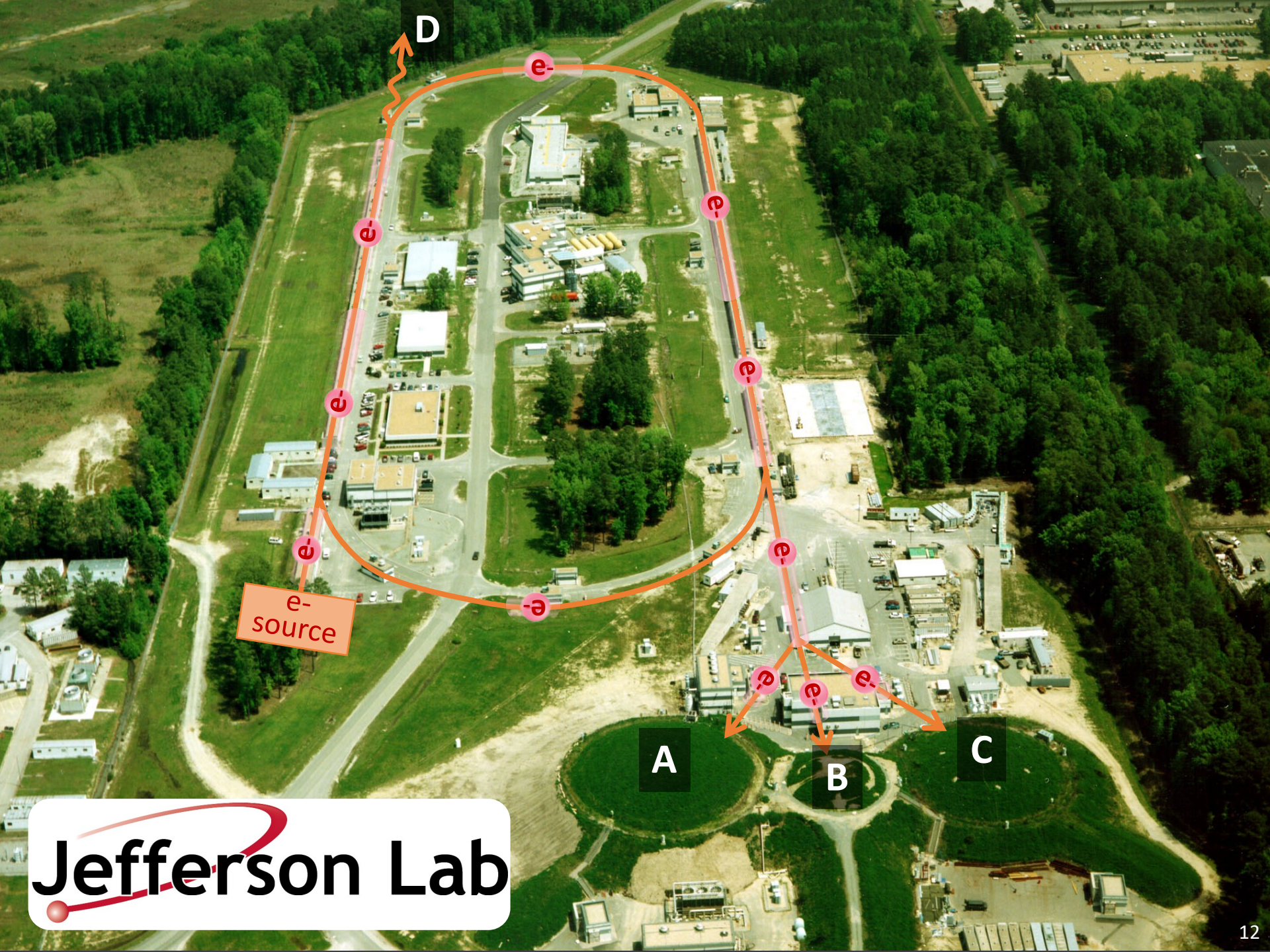
$$\rho_A^{NN,\alpha}(r) \underset{r < 1 \text{ fm}}{\cong} C_A^{NN,\alpha} \times |\psi_{NN}^\alpha(r)|^2$$

$\downarrow$  Total Dist. = Constant  $\times$  Two-body  
 (Low-Energy) (High-Energy)

$$\sigma^A \cong K \times \sigma^N \times \sum_{NN,\alpha} C_A^{NN,\alpha} |\psi_{NN}^\alpha|^2$$

Cruz-Torres et al., Nature Physics (2021)  
 Weiss et al., Phys. Lett. B 780 (2018)  
 Weiss, Bazak, Barnea, Phys. Rev. C 92 (2015)  
 Tropiano et al., Phys. Rev. C 104, 034311 (2021)  
 Lynn et al., JPG 47, 045109 (2020)  
 Chen, Detmold, Lynn, Schwenk, PRL 119 (2017)  
 Ryckebusch et al., Phys. Lett. B 792, 21 (2019)  
 ...





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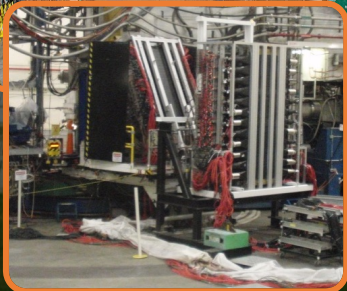
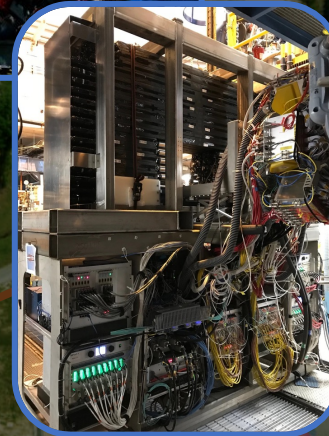
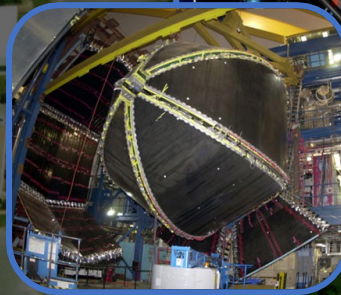
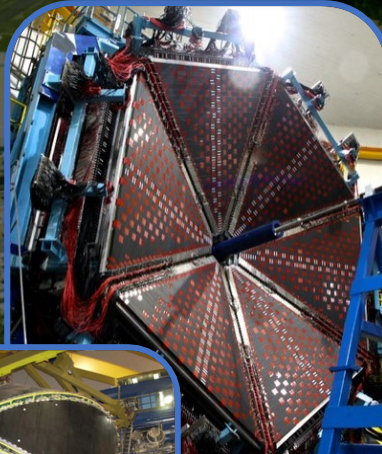
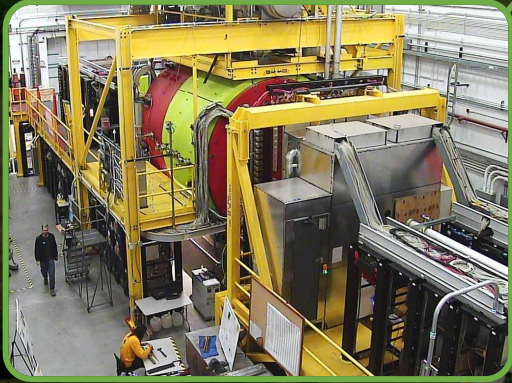
A

B

C

 Jefferson Lab

# SRC Program Span the Lab

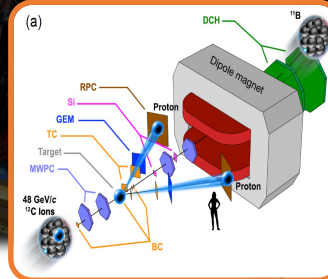


A

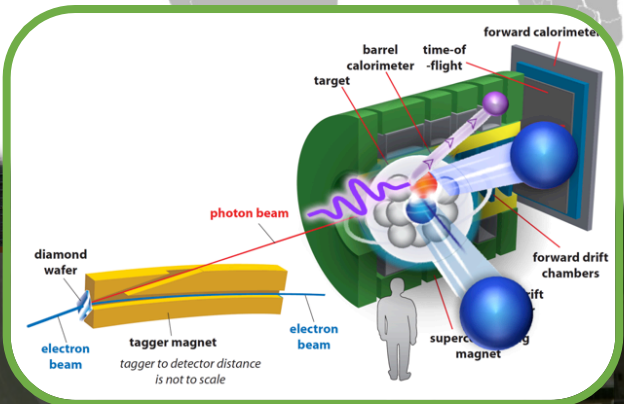
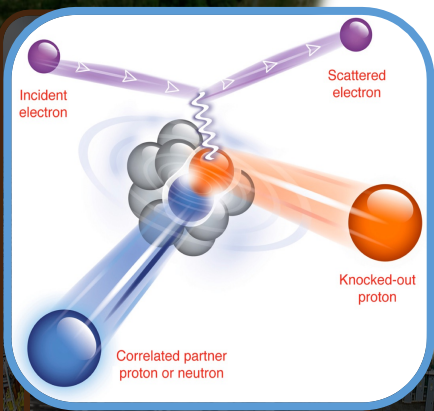
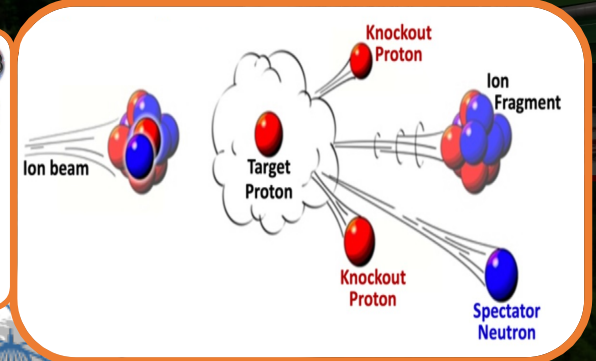
B

C

# SRC Program Span the world 🤪



$$(p, 2p \ N \ A - 2)$$



$$(e, e'NN)$$

$$(e, e'N)$$

$$(e, e')$$

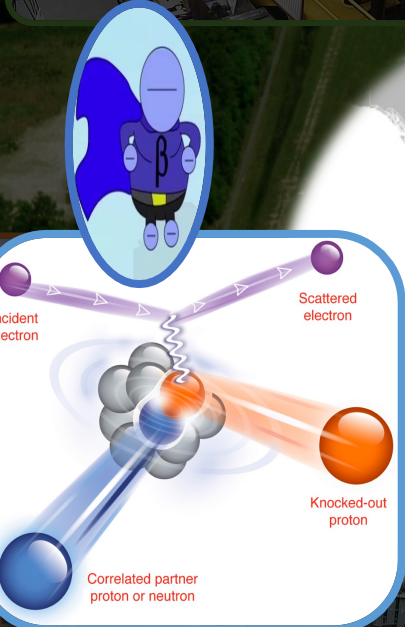
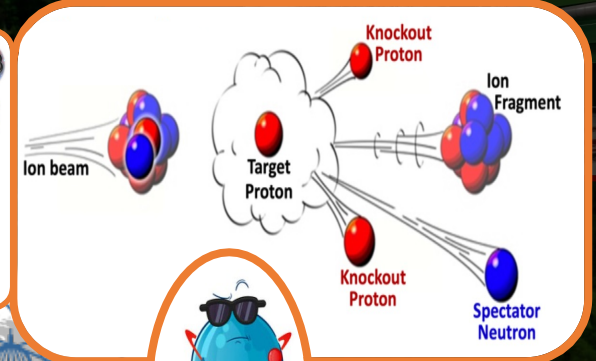
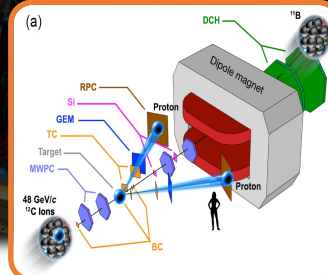
$$(\gamma, \rho NN)$$

$$(e, \pi NN)$$

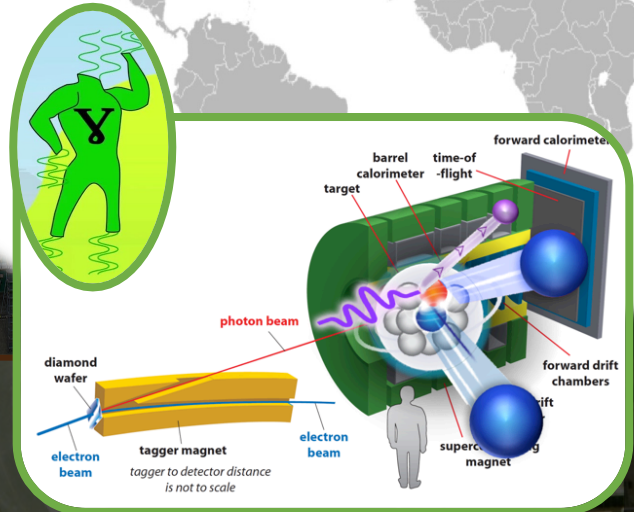
...

# Multi-probe / Reaction Approach

$(p, 2p \ N \ A - 2)$



$(e, e' NN)$   
 $(e, e' N)$   
 $(e, e')$

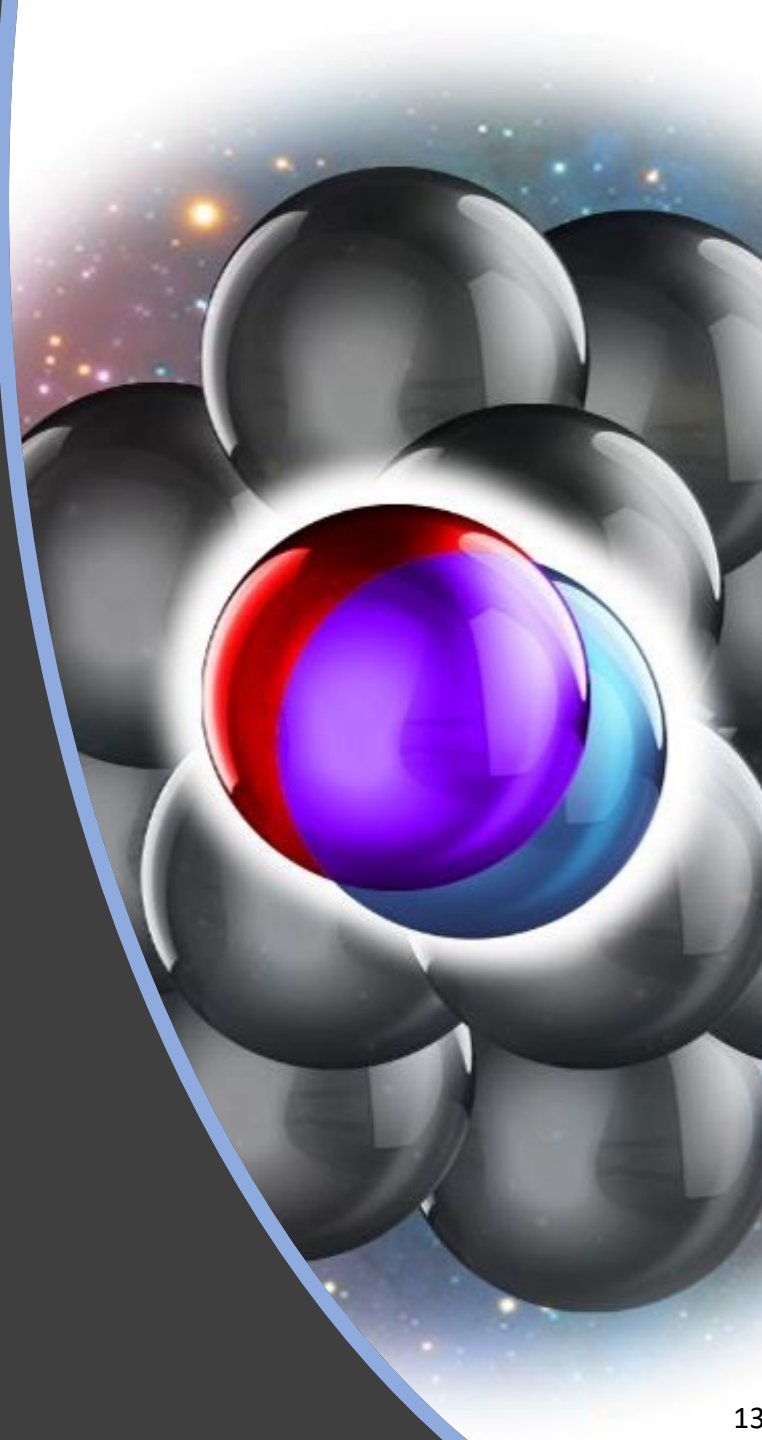
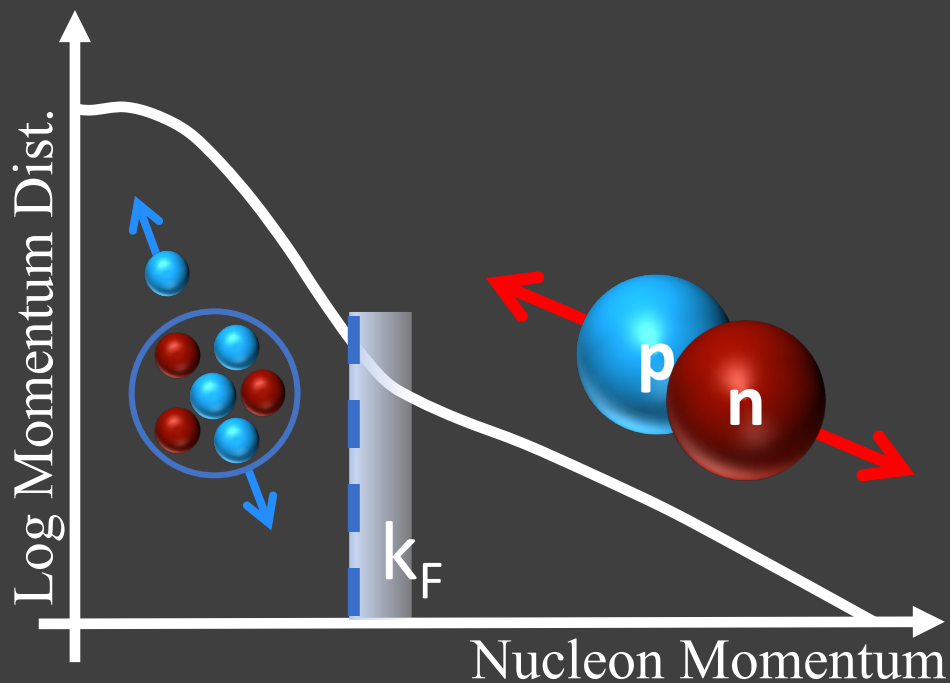


$(\gamma, \rho NN)$   
 $(e, \pi NN)$

...

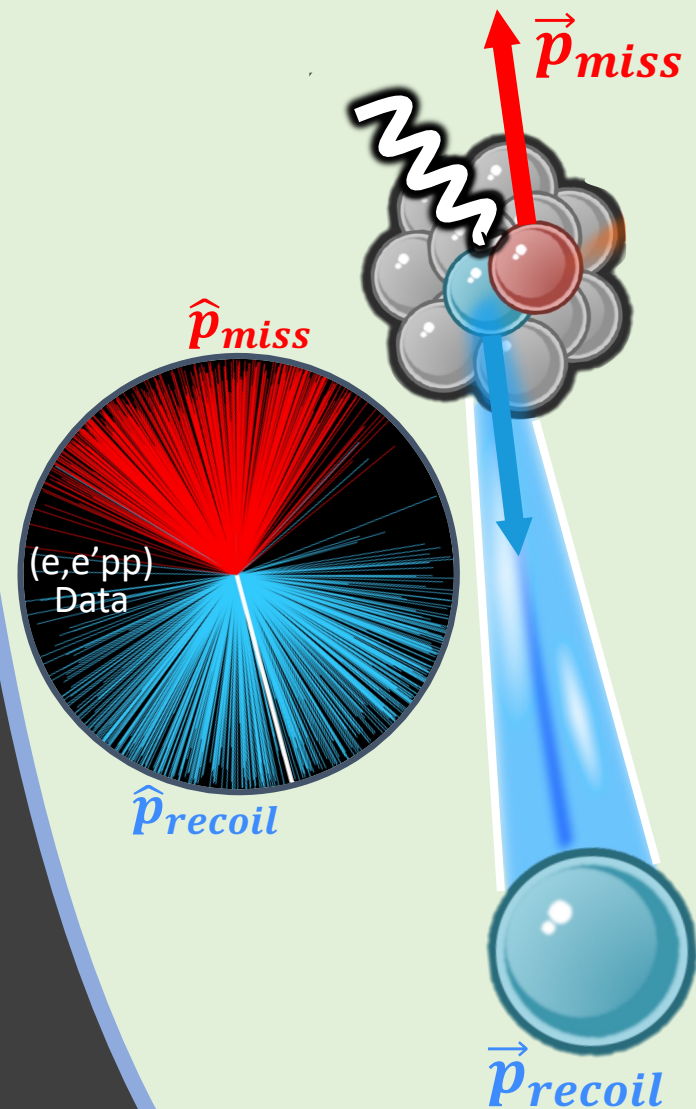
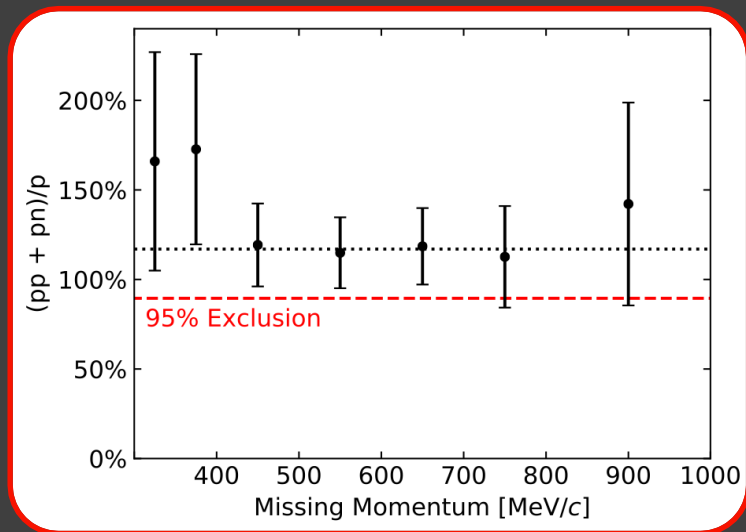
# Short-Range Correlations (SRC)

- Produce high missing-momentum ( $>k_F$ )
- Predominantly neutron-proton pairs
- Universal Deuteron-like Scaling
- Scale separated from residual system



# Short-Range Correlations (SRC)

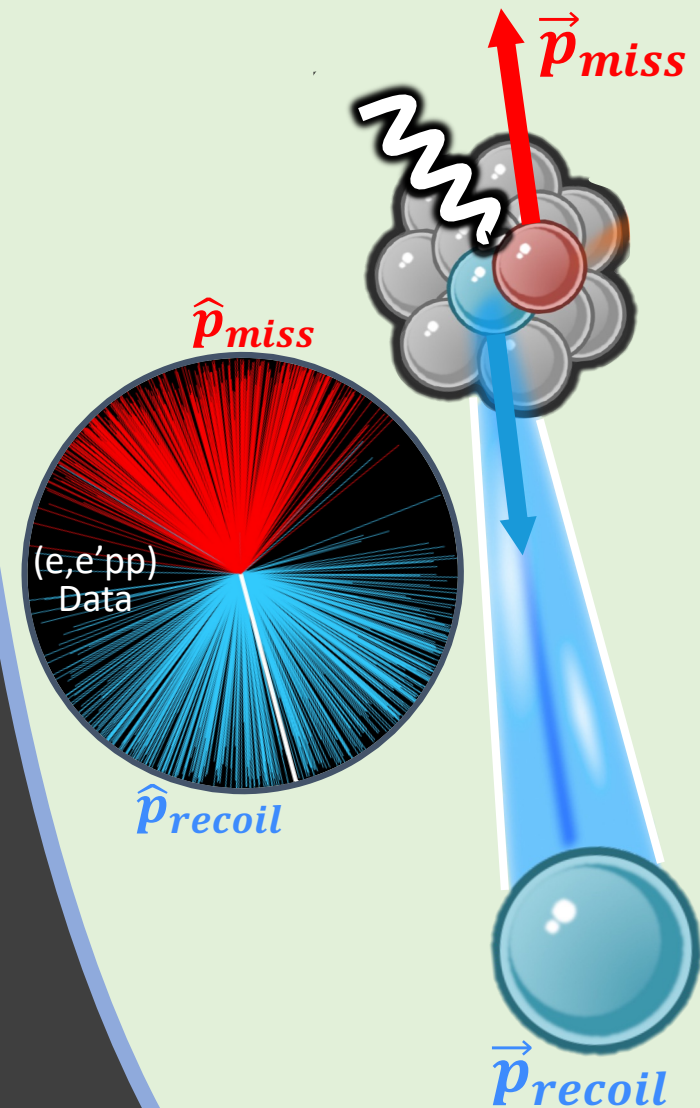
- Produce high missing-momentum ( $>k_F$ )
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Hen et al.,  
Science (2014)

# Short-Range Correlations (SRC)

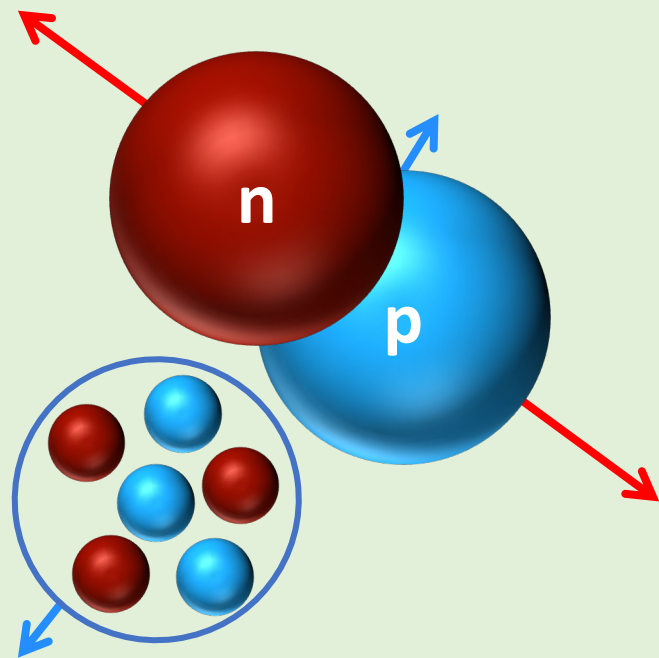
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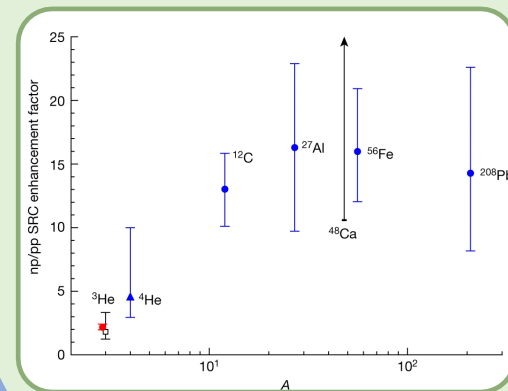
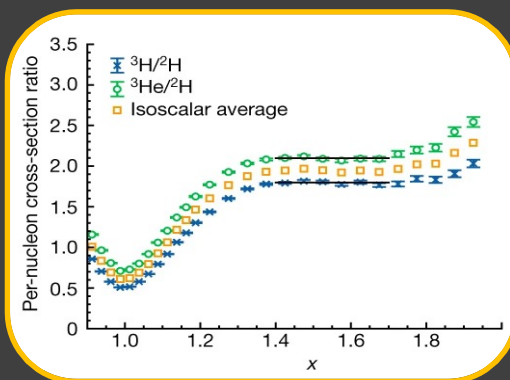
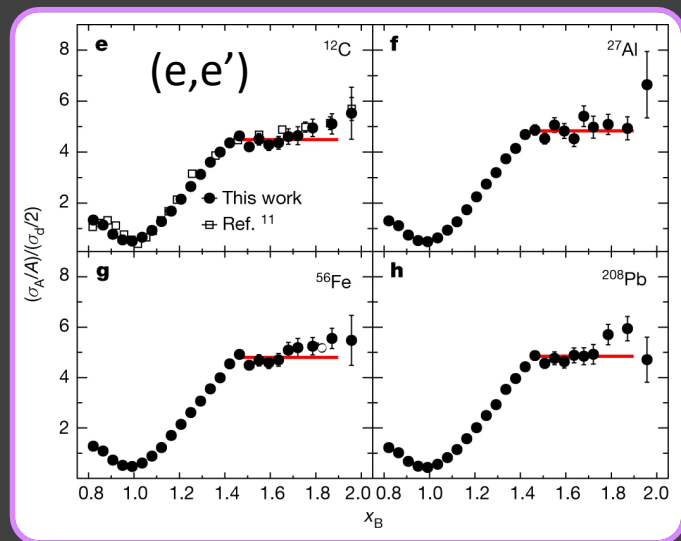


# Short-Range Correlations (SRC)

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Ratio to Deuterium

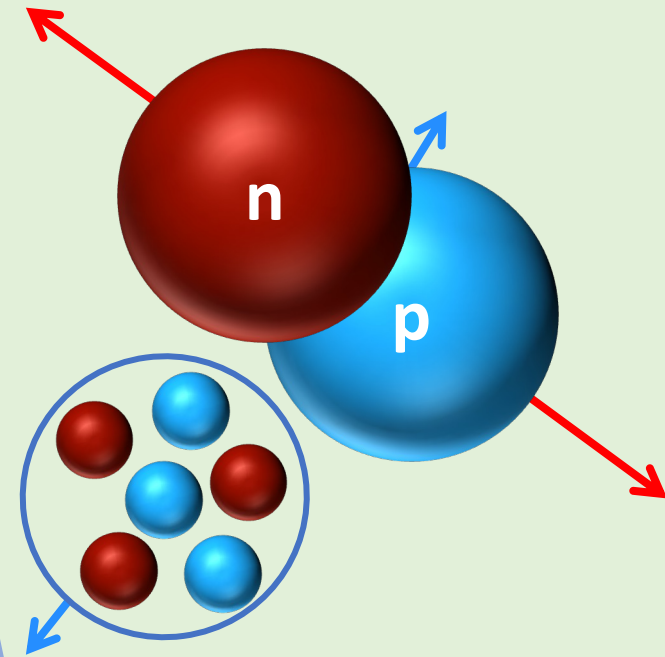


Exp: Li Nature '22, Nguyen PRC '20;  
Schmookler Nature '19; Fomin PRL '12; ...

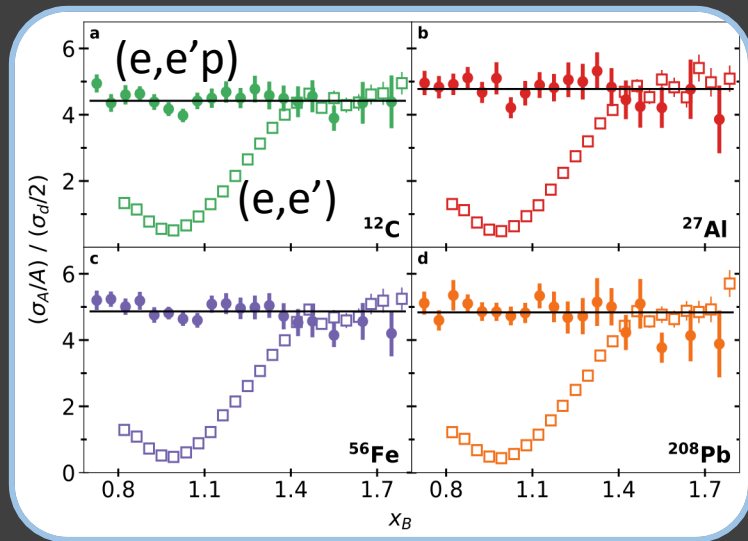
Theory example:  
Weiss et al.,  
PRC Lett '21

# Short-Range Correlations (SRC)

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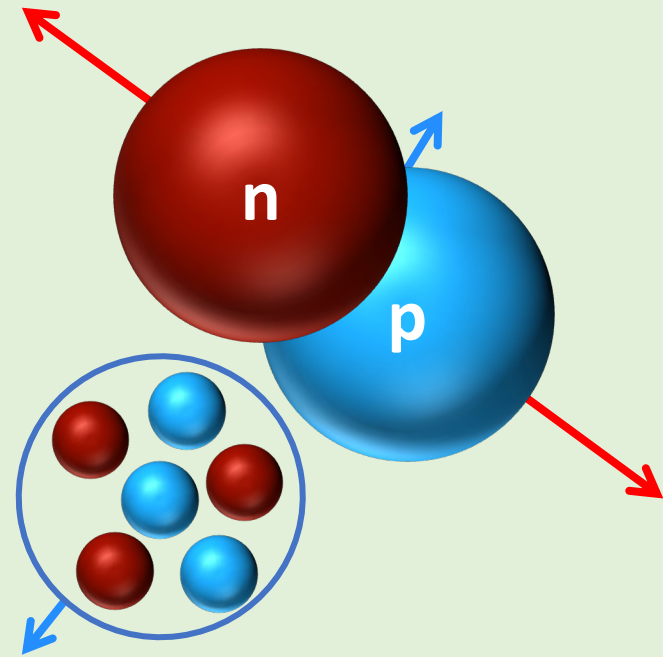
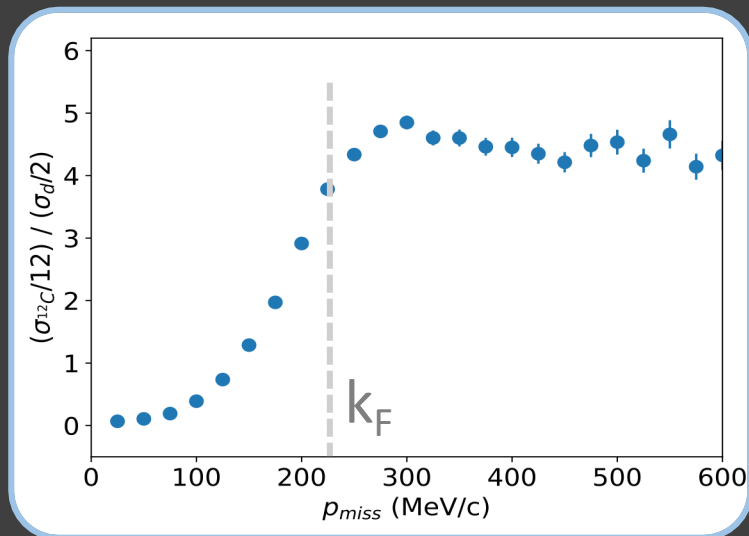
Ratio to Deuterium



# Short-Range Correlations (SRC)

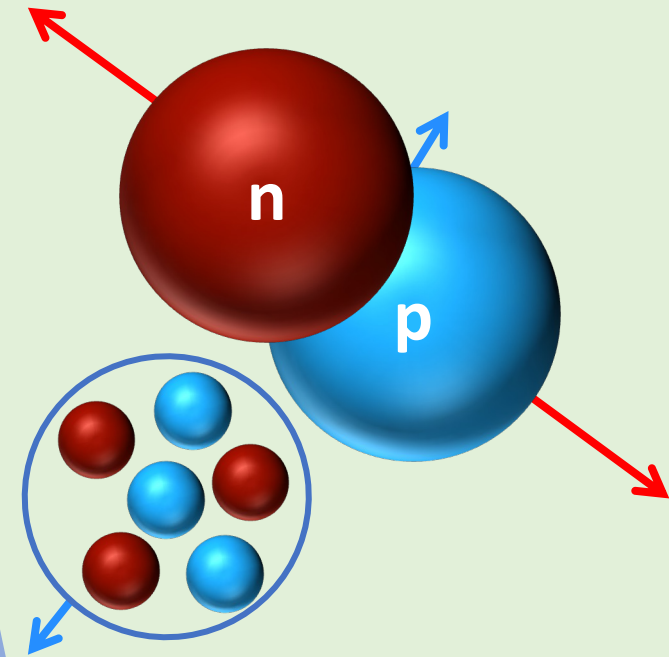
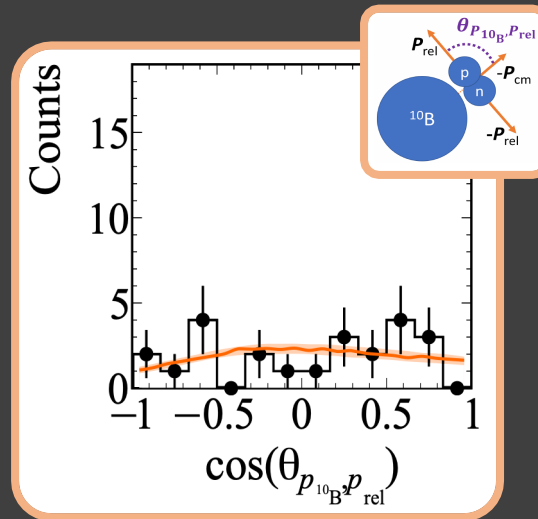
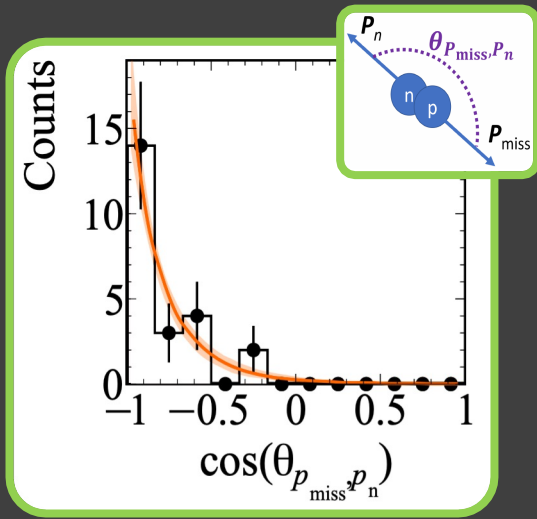
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Ratio to Deuterium



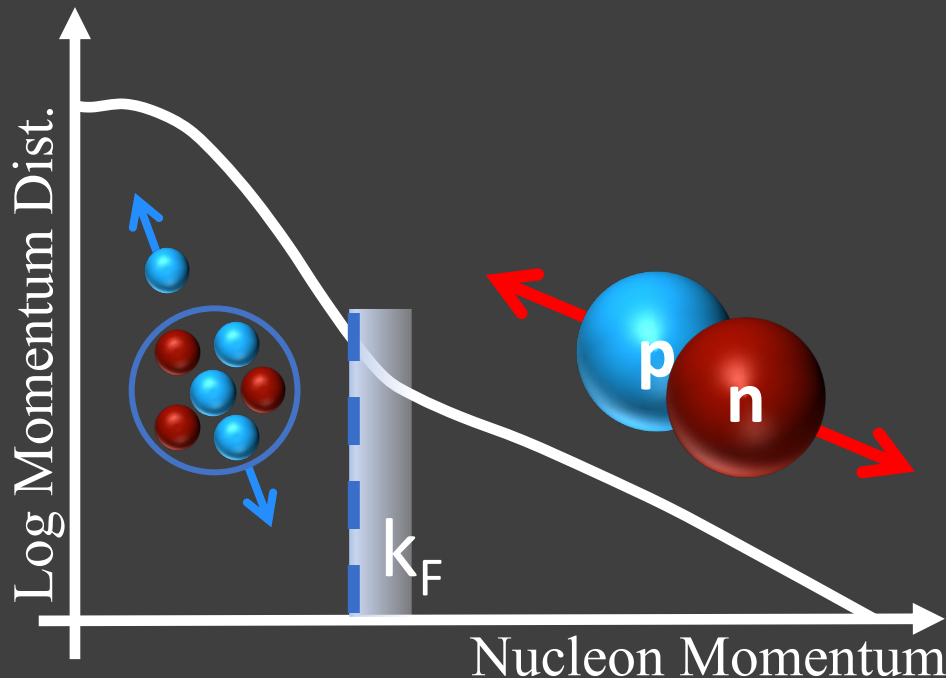
# Short-Range Correlations (SRC)

- Produce high missing-momentum ( $>k_F$ )
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# Short-Range Correlations (SRC)

- Produce high missing-momentum ( $>k_F$ )
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## Isospin Structure:

- Nature 609, 41 (2022)
- Phys. Rev. C 102, 0644004 (2020)
- Phys. Rev. Lett. 122, 172502 (2019)
- Nature 560, 617 (2018)
- Science 346, 614 (2014)
- Phys. Rev. Lett. 113, 022501 (2014)

## C.M. Motion:

- Phys. Rev. Lett. 121, 092501 (2018)

## Hard-Reaction Dynamics:

- Nature Physics 17, 693 (2021)
- Phys. Lett. B 797, 134792 (2019)
- Phys. Lett. B 722, 63 (2013)

## Nuclei / Nuclear Matter Properties:

- Phys. Lett. B 800, 135110 (2020)
- Phys. Lett. B 793, 360 (2019)
- Phys. Lett. B 785, 304 (2018)
- Phys. Rev. C 91, 025803 (2015)

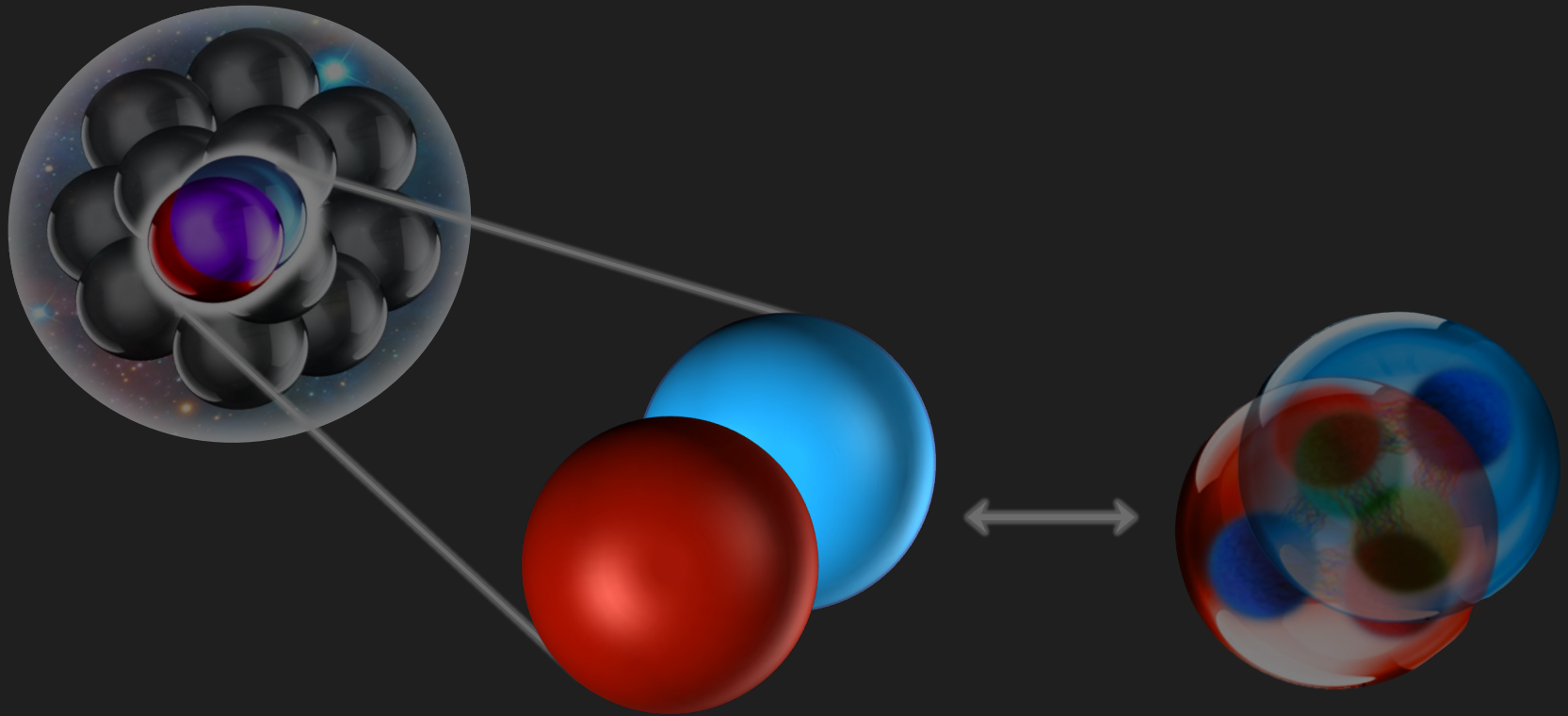
## Effective Theory:

- Nature Physics 17, 306 (2021)
- Phys. Lett. B 805, 135429 (2020)
- Phys. Lett. B 791, 242 (2019)
- Phys. Rev. C 104, 034311 (2021)
- Phys. Lett. B 792, 21 (2019)

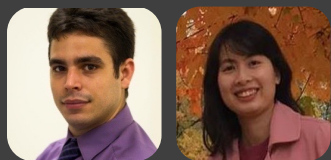
## Quantum Numbers, Mass, Asymmetry Dependence:

- Phys. Rev. C 103, L031301 (2021)
- Phys. Lett. B 780, 211 (2018)
- PRC 92, 024604 (2015)
- PRC 92, 045205 (2015)

# High-resolution Interactions

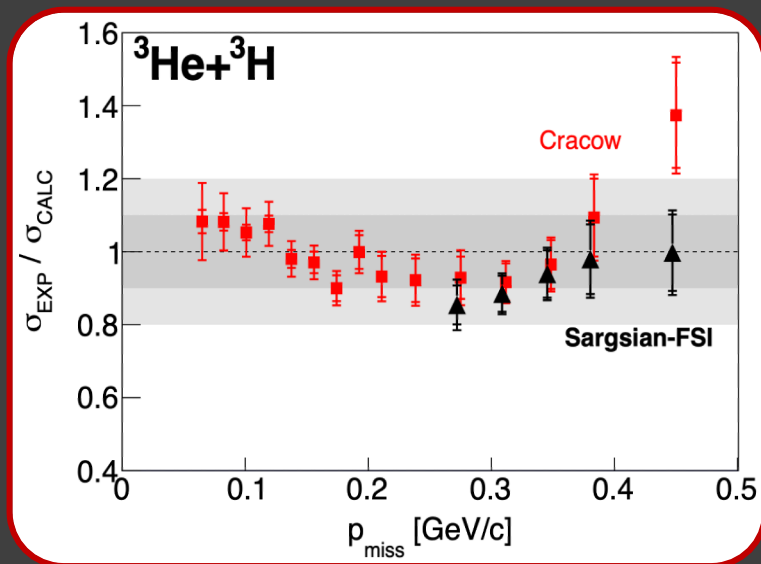


# (e,e'p) Few-Body Cross-Sections

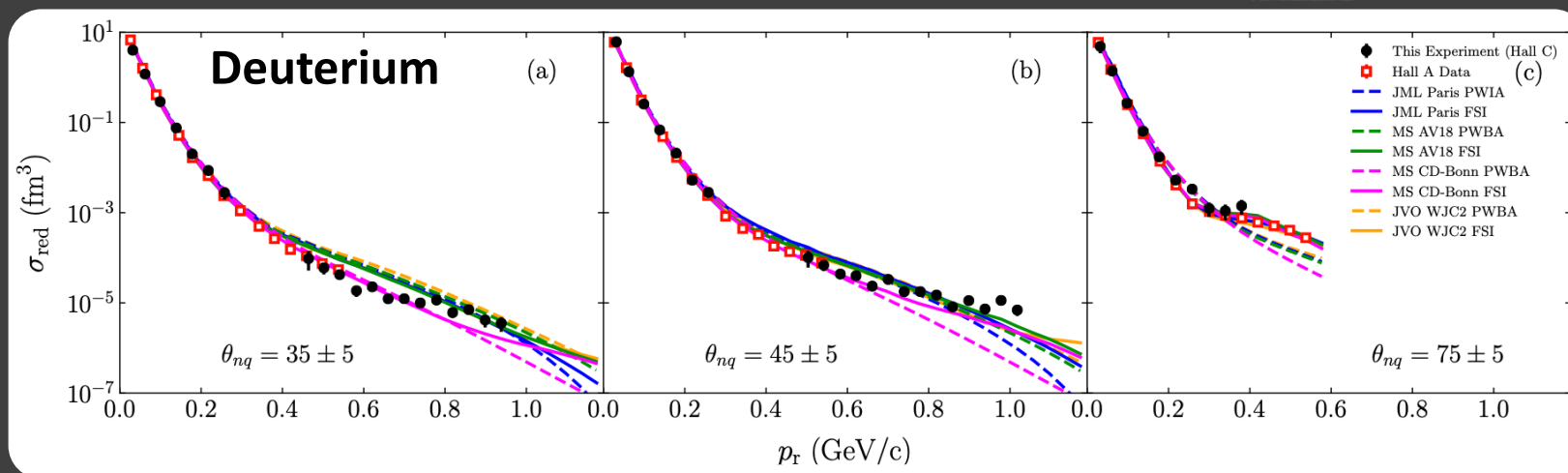


Editors' Suggestion

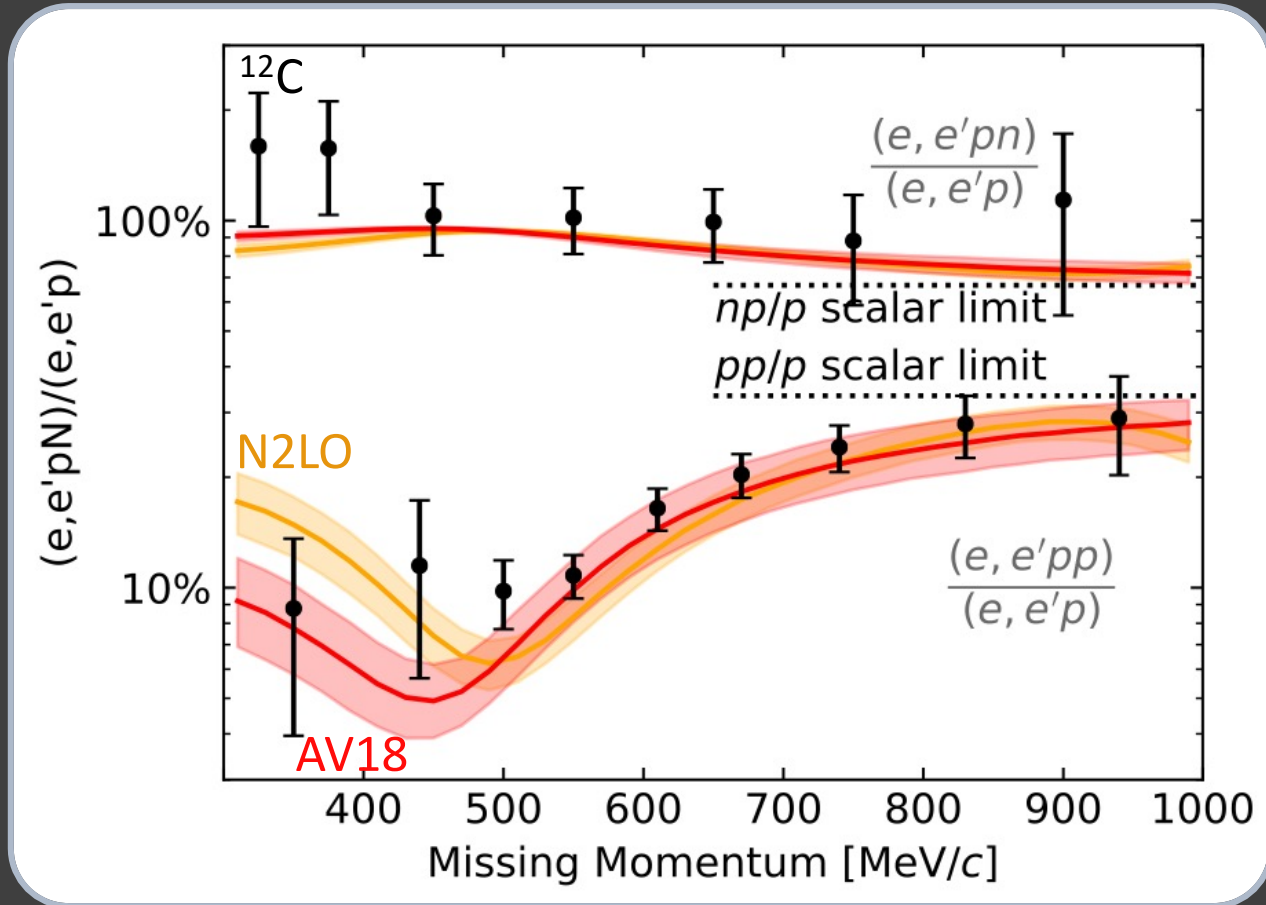
Cruz Torres and Nguyen et al., Phys. Rev. Lett (2020)



Yero et al., Phys. Rev. Lett (2020)



# pp-SRC Fraction Increase

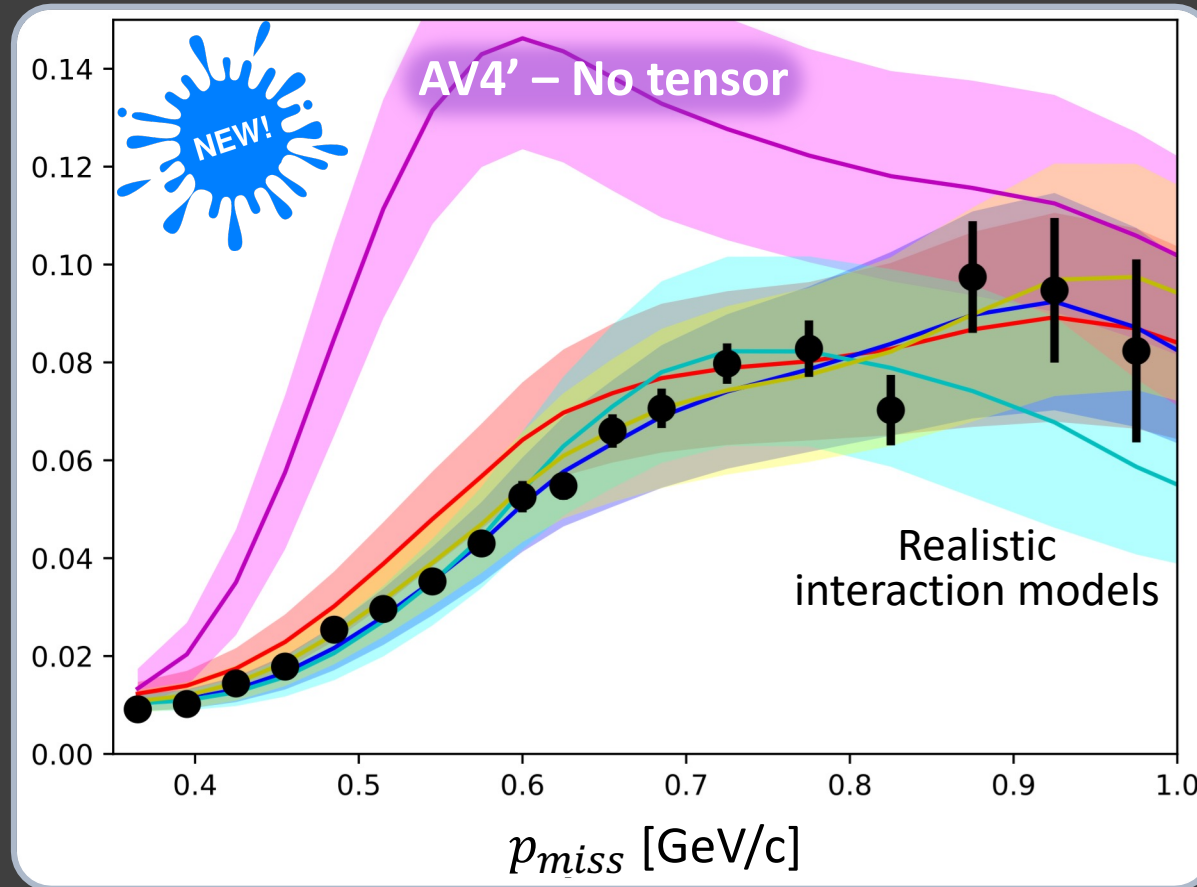


Schmidt and Pybus et al., Nature (2020)  
Pybus et al., PLB (2020);  
Korover and Pybus et al., PLB (2021)

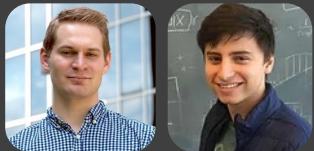


# New: High-Precision Data

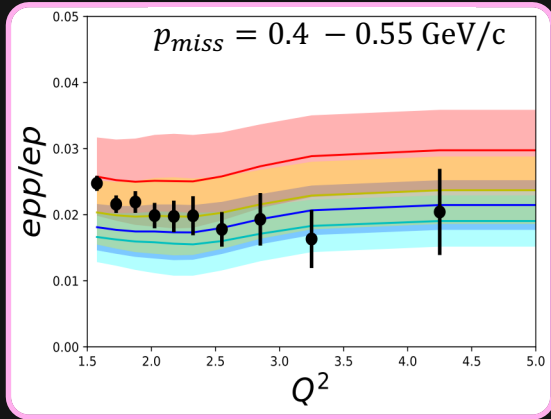
$$\frac{{}^4\text{He}(e, e'pp)}{{}^4\text{He}(e, e'p)}$$



\*40% of collected data



# Scale Independence



$(e, e'pN)$

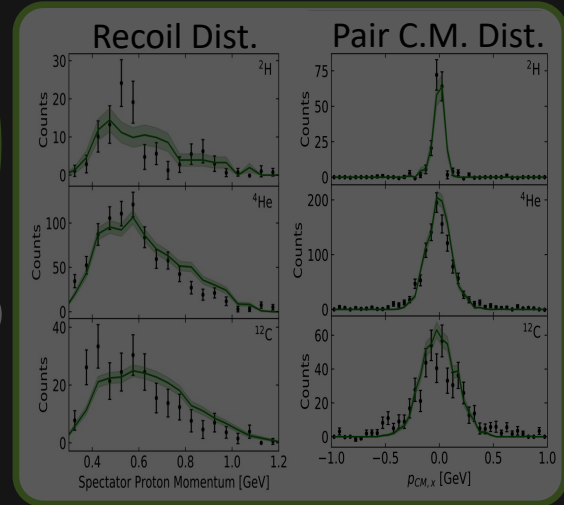
Electromagnetic

Hadronic

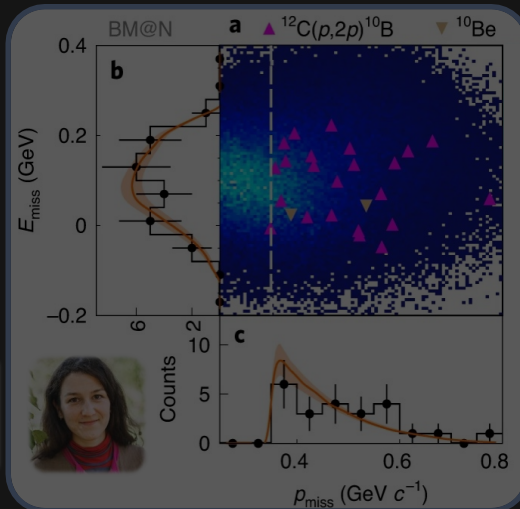
Photonuclear



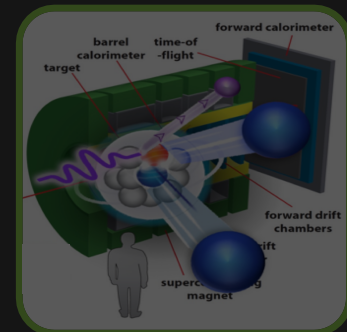
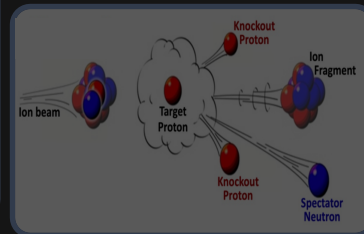
$(\gamma, \rho^- pp)$



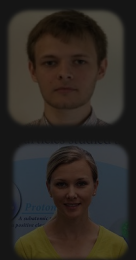
\*40% of collected data



$(p, 2p A-2)N$

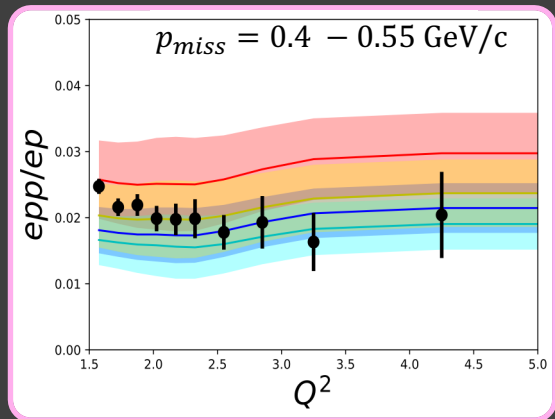


<20% of collected data



Patsyuk and Kahlbow et al., Nature Physics (2021)

# Scale and Probe Independence



$(e, e'pN)$

$$\sigma_{probe}^A = K \times \sigma_{probe}^N \times \sum_{NN,\alpha} C_{NN,\alpha}^A |\psi_{NN,\alpha}|^2$$

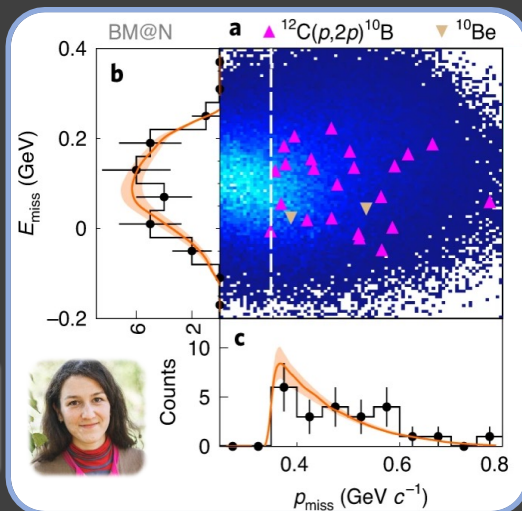
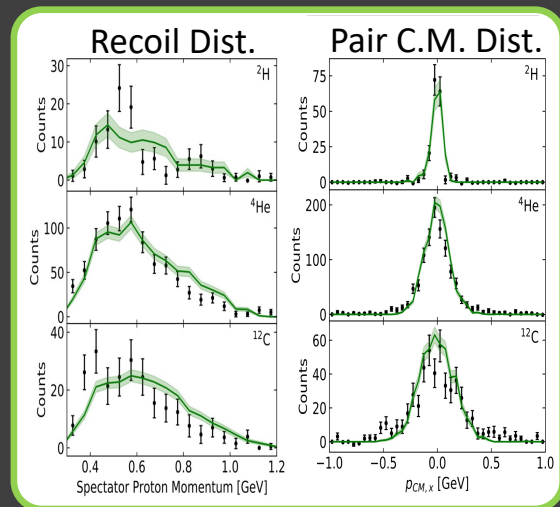
Electromagnetic

Hadronic

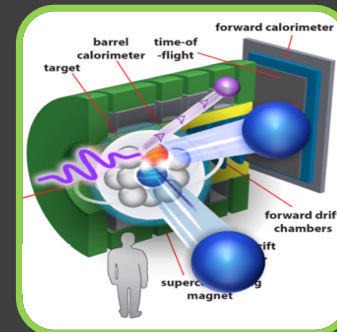
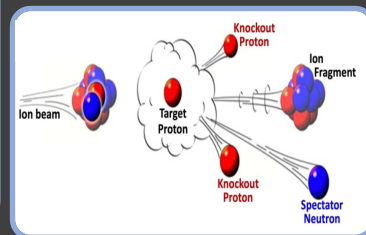
Photonuclear



$(\gamma, \rho^- pp)$



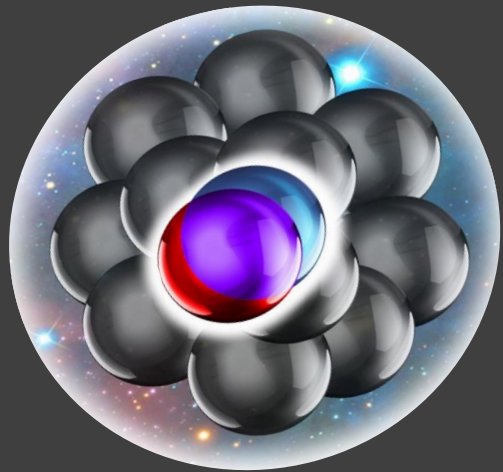
$(p, 2p A-2)N$



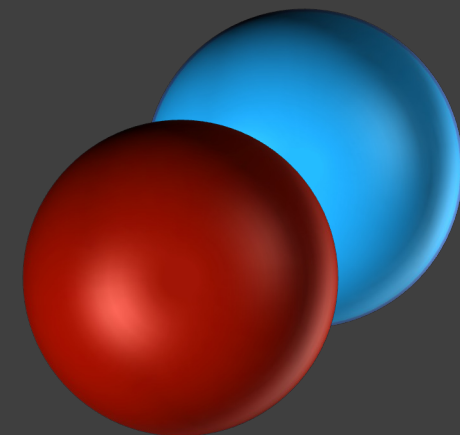
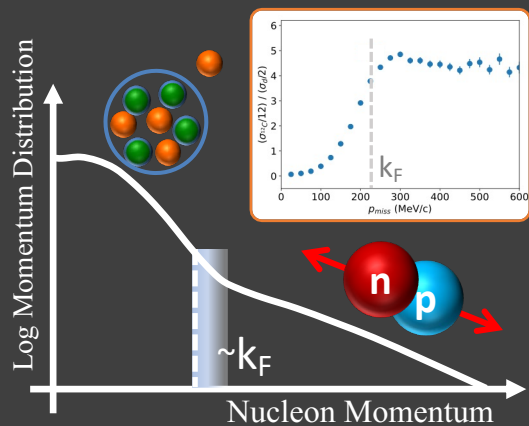
<20% of collected data



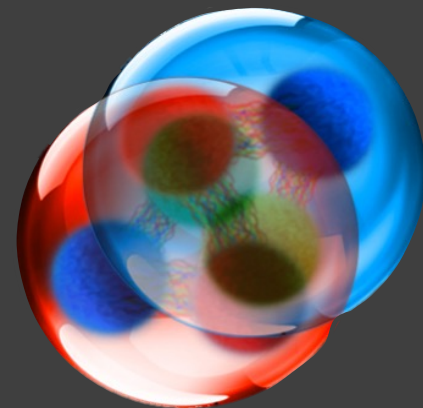
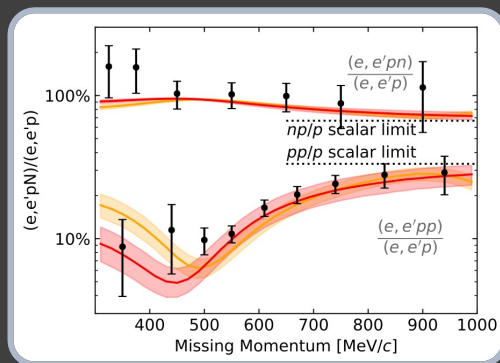
Patsyuk and Kahlbow et al., Nature Physics (2021)



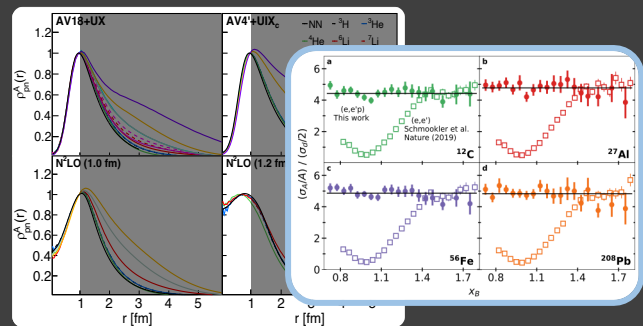
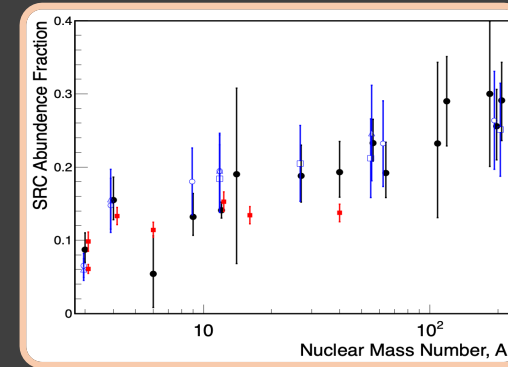
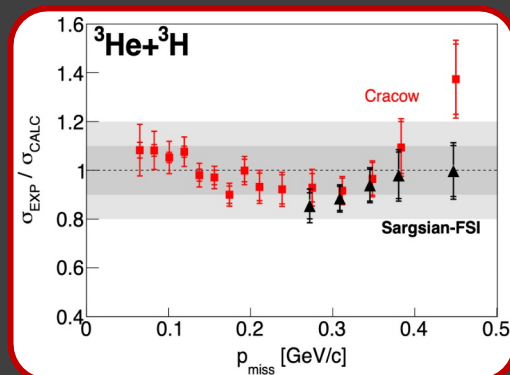
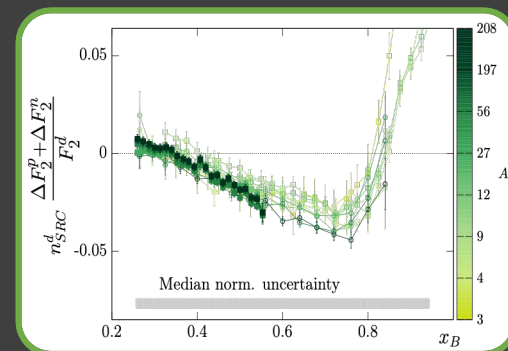
## High-resolution nuclear structure

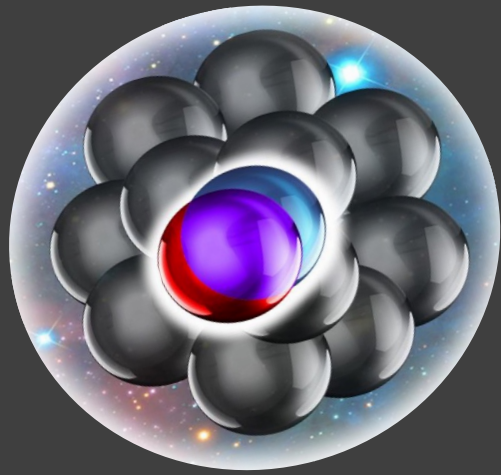


## Nucleon-nucleon interaction

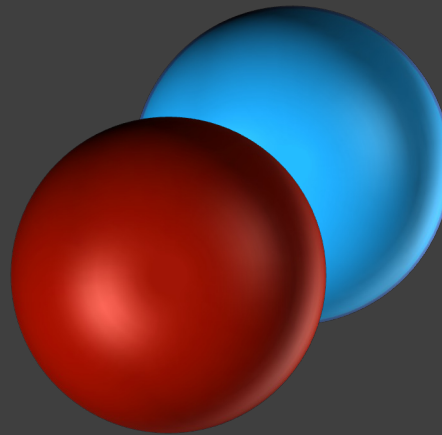


## Quark-gluon structure of nuclei

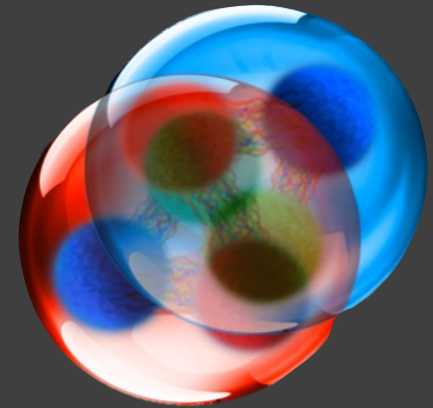




High-resolution nuclear structure

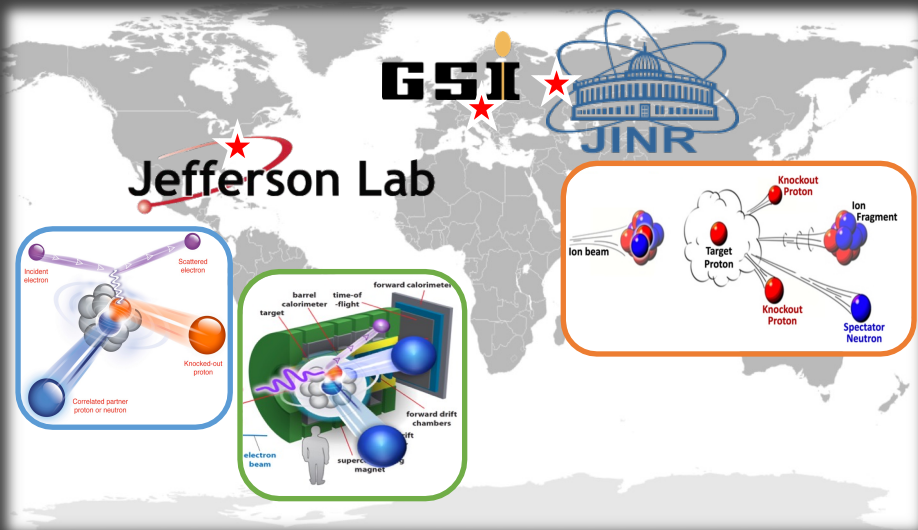


Nucleon-nucleon interaction



Quark-gluon structure of nuclei

## Growing International Effort + Forthcoming Electron-Ion Collider



# What's Coming?

JLab – electron:

- $d(e,e'p)$ : Precision data at high missing-momentum
- High-statistics  $(e,e')$ ,  $(e,e'N)$  and  $(e,e'Np)$  on  ${}^3\text{He}$ - ${}^3\text{H}$  mirror nuclei
- $(e,e')$  and  $(e,e'p)$ : stable light- and medium-mass asymmetric nuclei
- High-statistic  $(e,e'NN)$  spanning light ( ${}^4\text{He}$ ) to heavy ( ${}^{120}\text{Sn}$ ) nuclei

JLab – photon:

- Wide array of photonuclear reactions with  $d$ ,  ${}^4\text{He}$ , and  ${}^{12}\text{C}$

GSI:

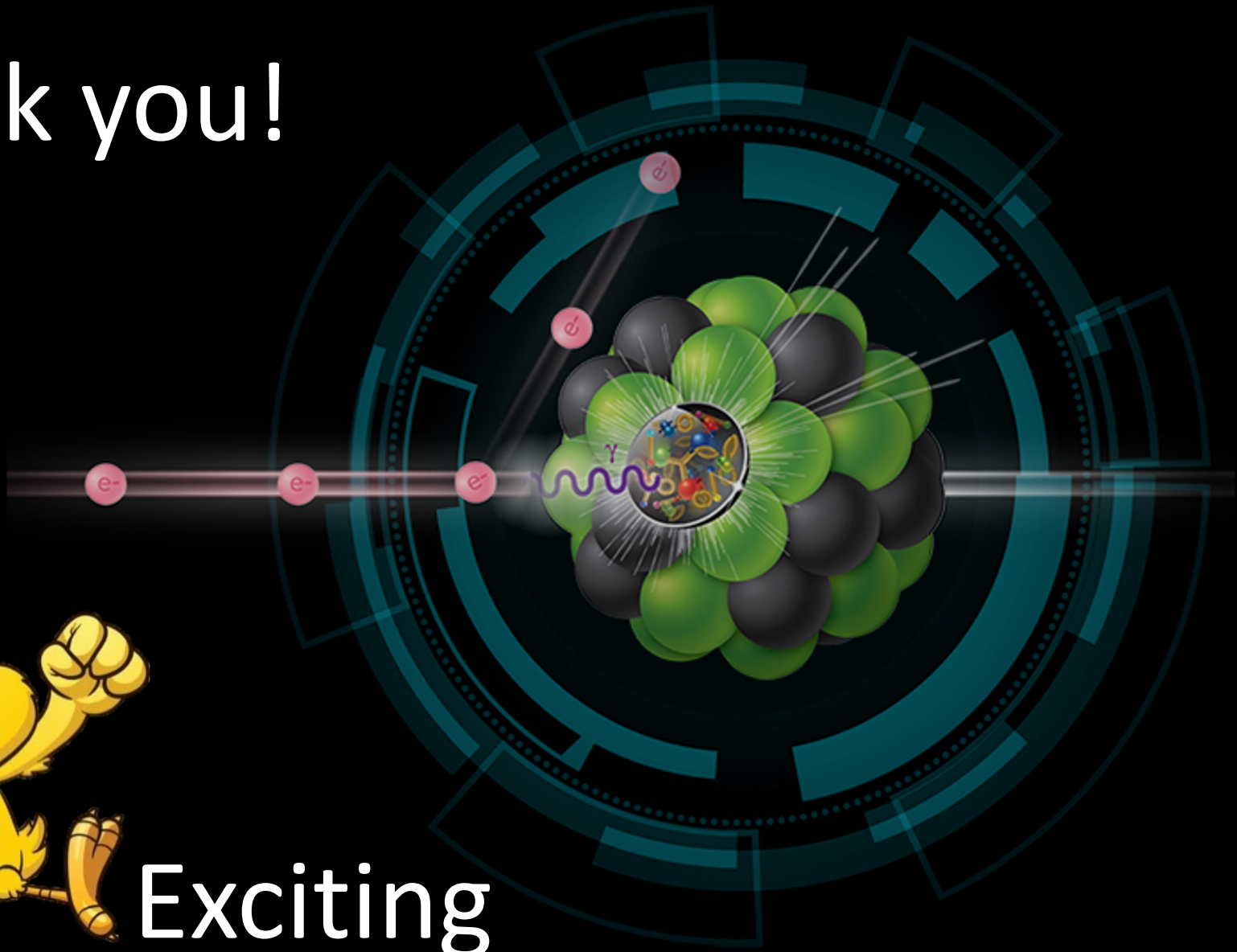
- $(p,2p A-2 N)$ :  $\sim 1.5$  GeV/u Carbon 12 & 16
- $(p,d)$  SRC study

JINR:

- $(p,2p A-2 N)$ : high-stat. + PID with  $\sim 3$  GeV/u Carbon 12

**Your ideas here!**

# Thank you!



# Exciting Times Ahead!

# Backup

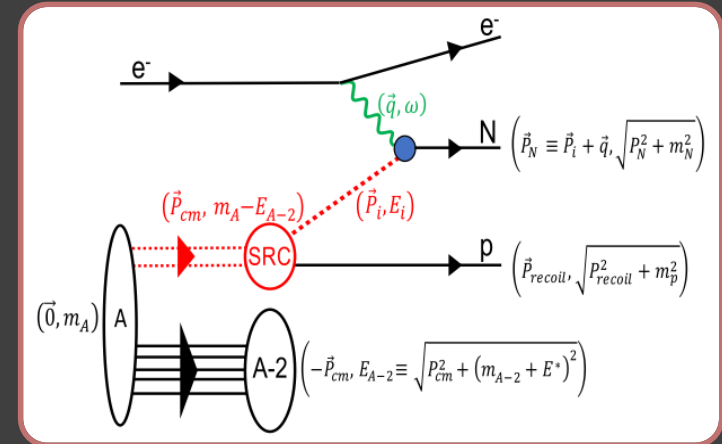


# Using Hard Reactions

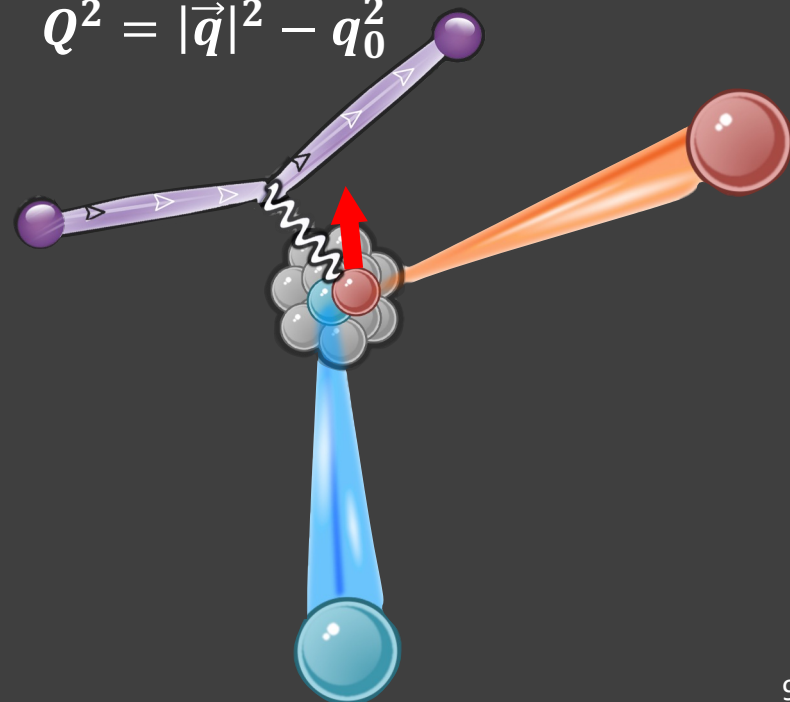
Probe energy  $\gg$  Nuclear energy scale

$$q_0, \vec{q} > 1\text{GeV} \gg |V_{NN}|, 2k_F$$

- Provides resolving power



$$Q^2 = |\vec{q}|^2 - q_0^2$$



Ciofi degli Atti, Physics Reports (2015)

Sargsian+, Phys. Rev. C (2005)

Frankfurt, Sargsian, and Strikman, Phys. Rev. C (1997)

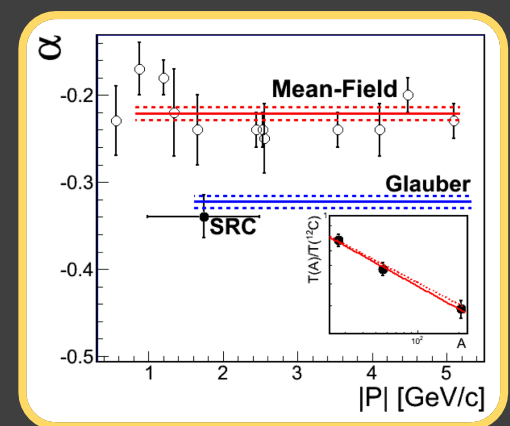
Frankfurt and Strikman, Physics Reports (1981)

# Using Hard Reactions

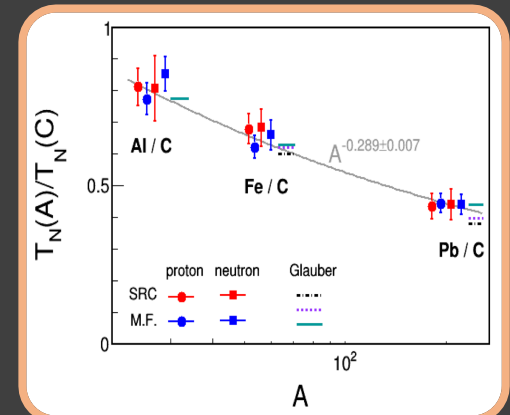
Probe energy  $\gg$  Nuclear energy scale

$$q_0, \vec{q} > 1\text{GeV} \gg |V_{NN}|, 2 k_F$$

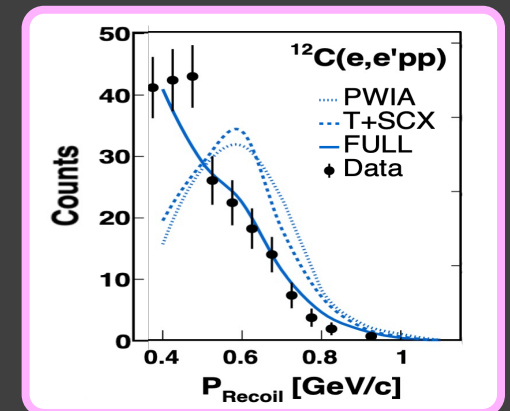
- Provides resolving power
- Use verifiable Glauber calculations for Final State Interaction (FSI) assessment and correction



Hen+ Phys. Lett. B (2013)



Duer+ Phys. Lett. B (2019)



Wright+ arXiv (2022)

Ciofi degli Atti, Physics Reports (2015)

Sargsian+, Phys. Rev. C (2005)

Frankfurt, Sargsian, and Strikman, Phys. Rev. C (1997)

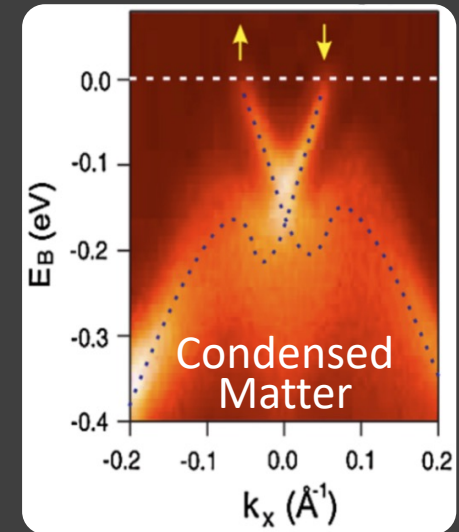
Frankfurt and Strikman, Physics Reports (1981)

# Using Hard Reactions

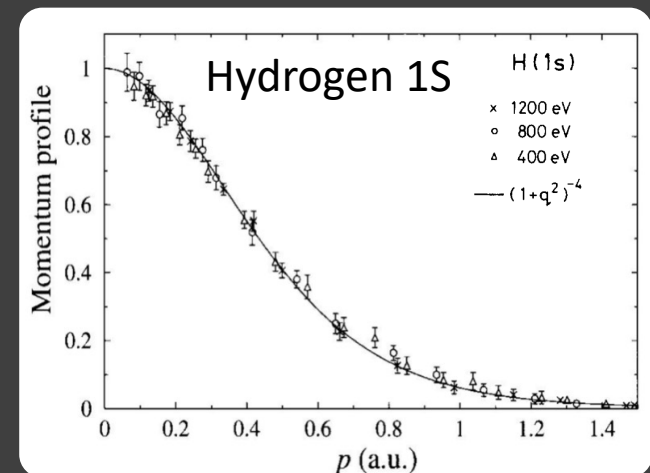
Probe energy  $\gg$  Nuclear energy scale

$$q_0, \vec{q} > 1\text{GeV} \gg |V_{NN}|, 2k_F$$

- Provides resolving power
- Use verifiable Glauber calculations for Final State Interaction (FSI) assessment and correction
- Similar to atomic / condensed matter spectral function extractions



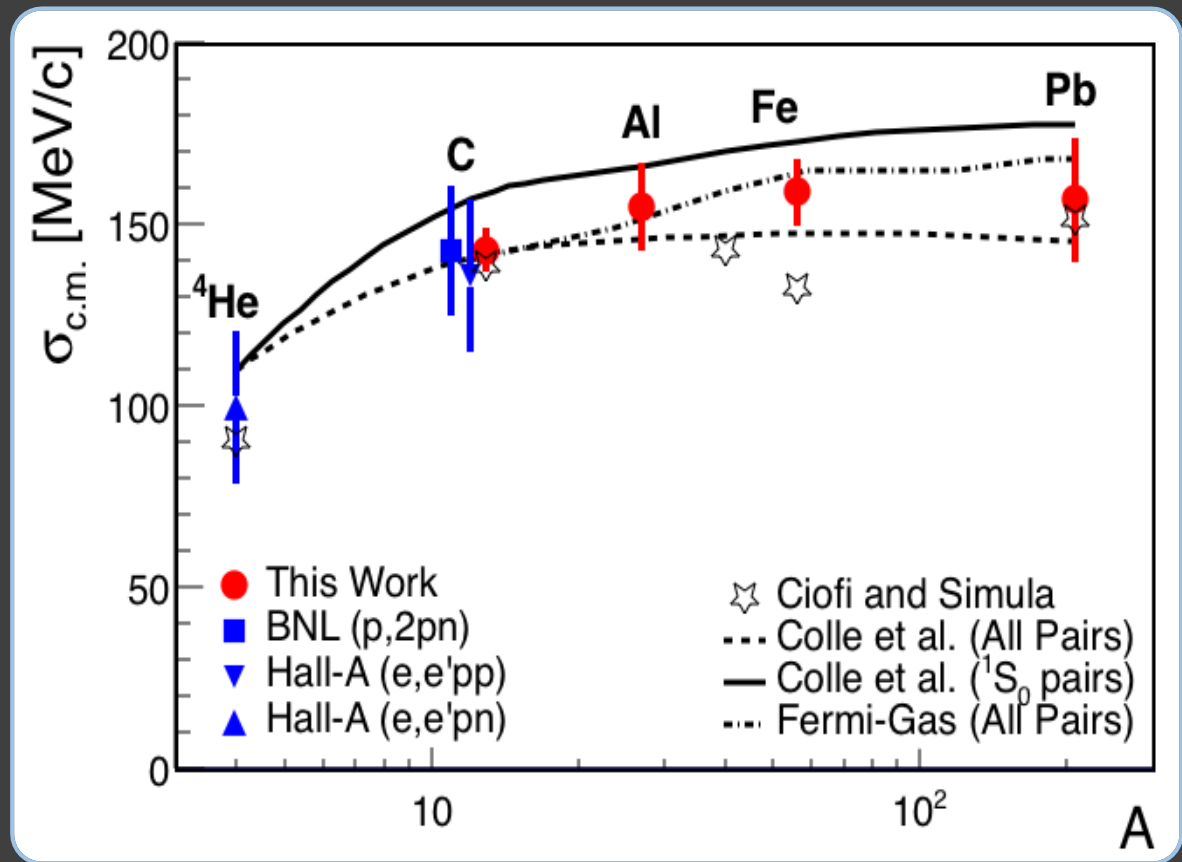
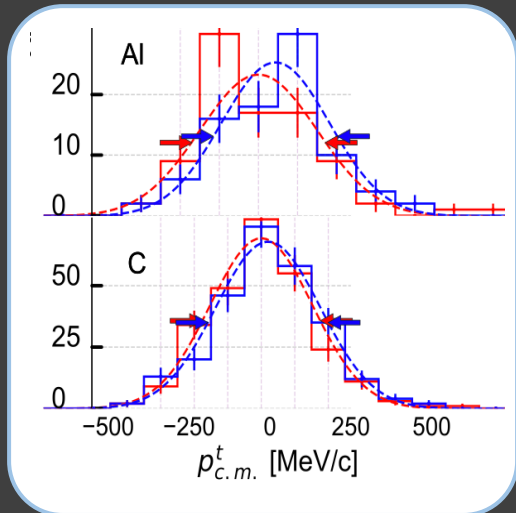
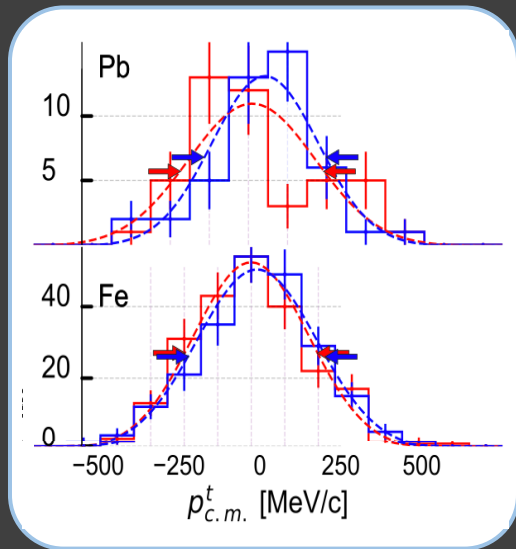
( $\gamma, e$ ) ARPES: surface state bands in  $\text{Bi}_{2-x}\text{Mn}_x\text{Te}_3$  (2009)



( $e, 2e$ ) reduced cross-section (1981)

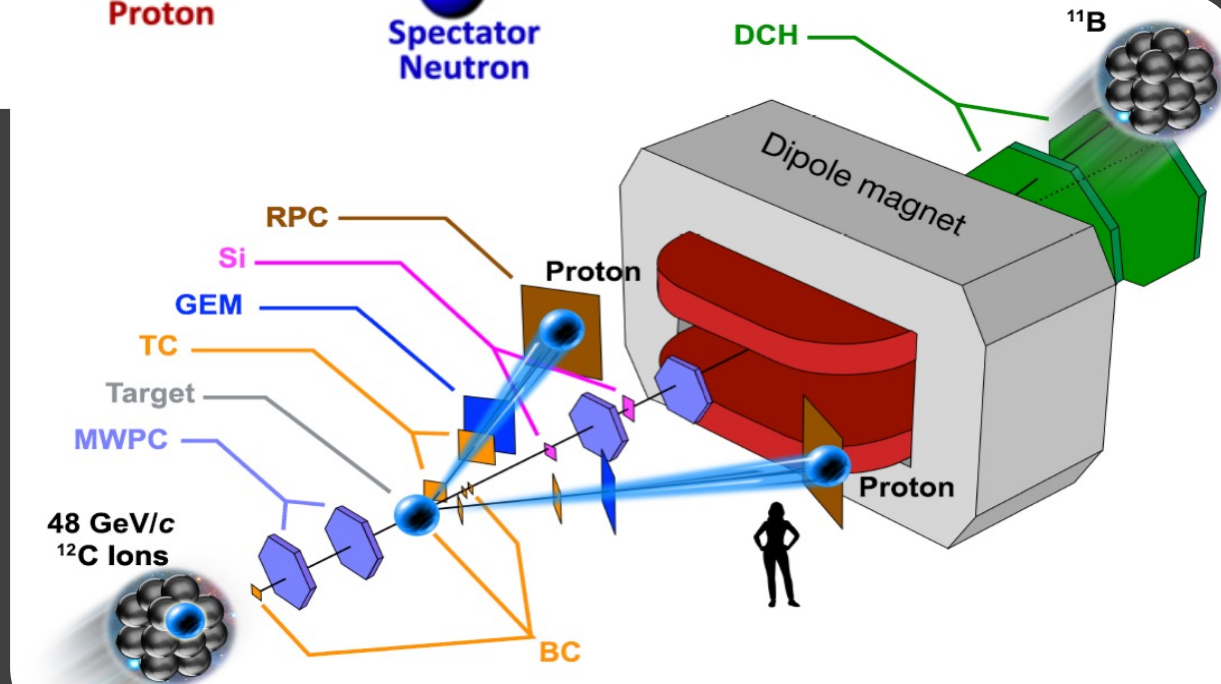
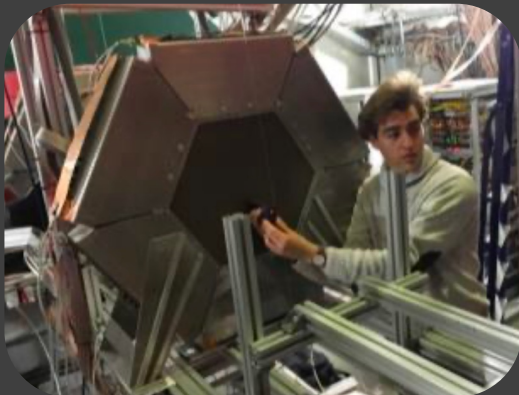
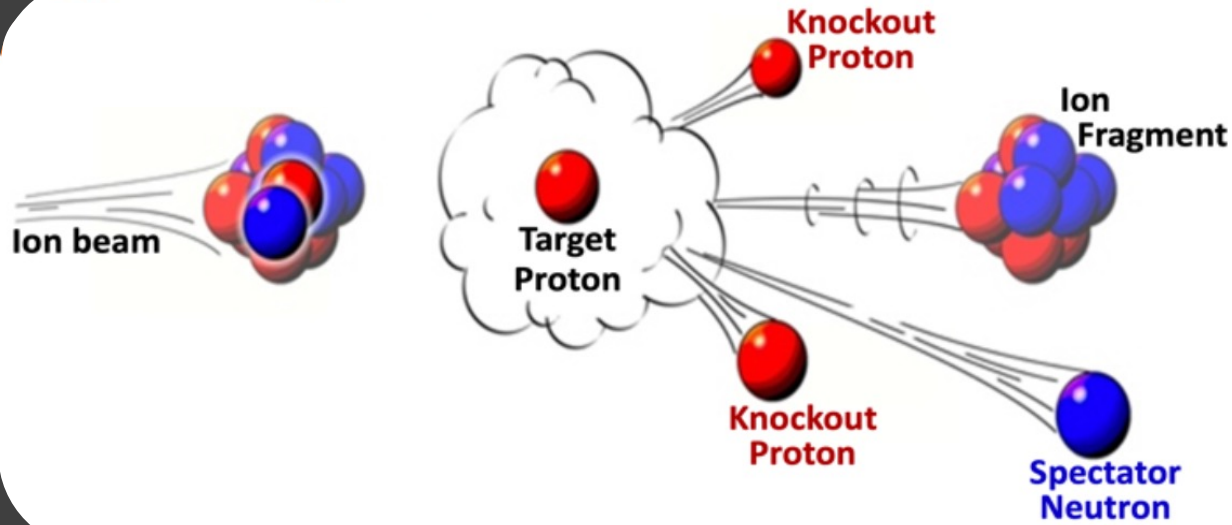
# Strongly correlated? Yes!

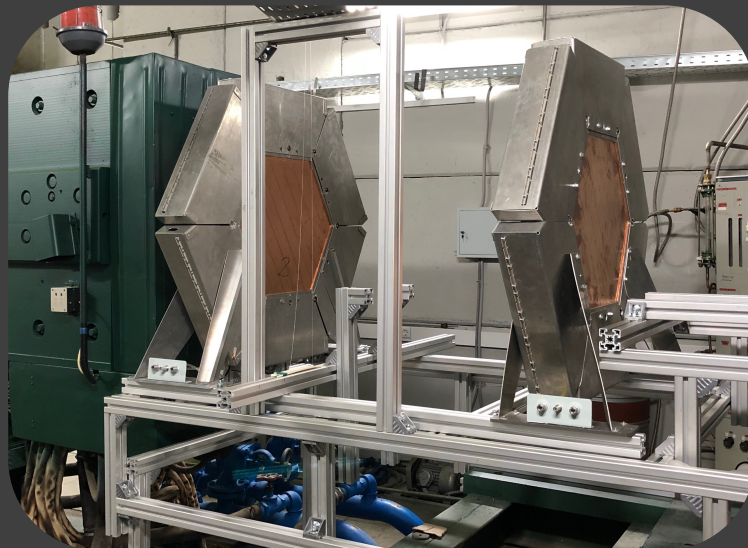
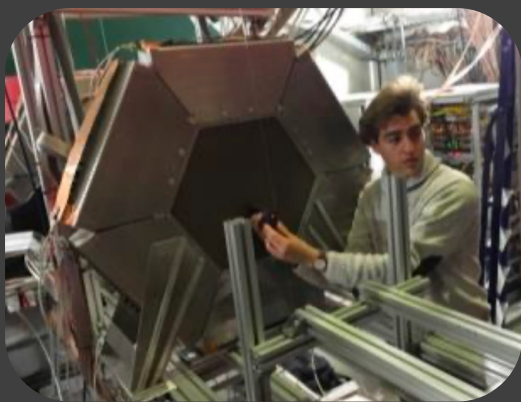
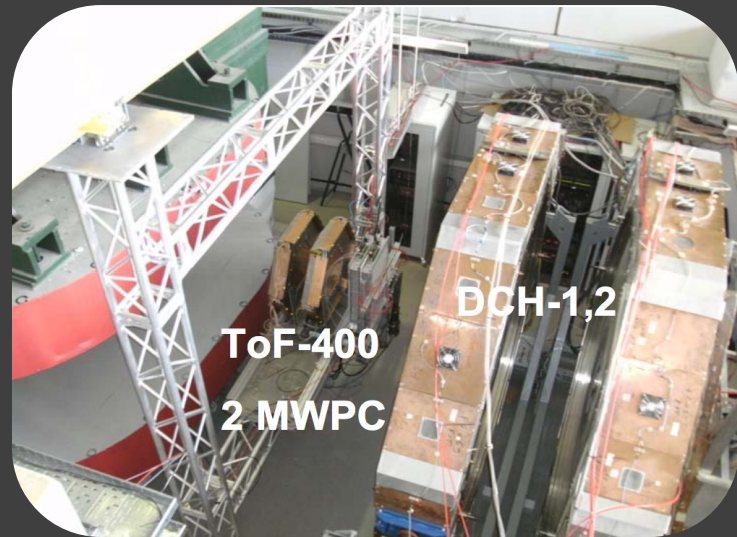
## low pair c.m. motion



Cohen et al.,  
PRL (2018)

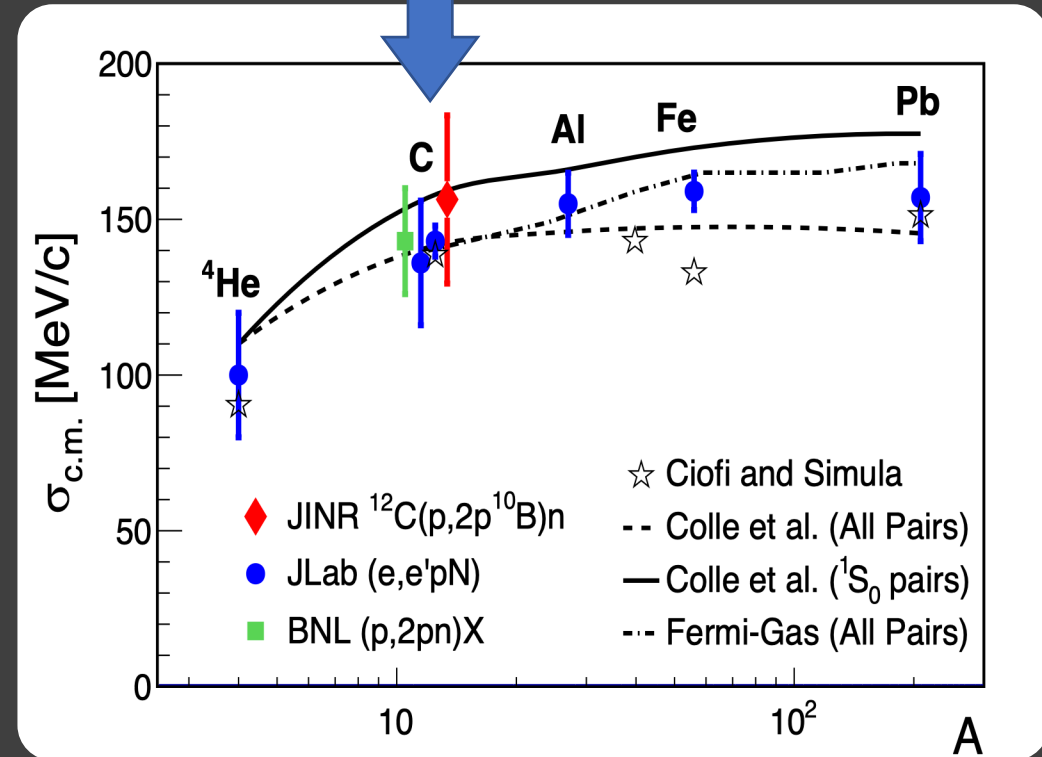
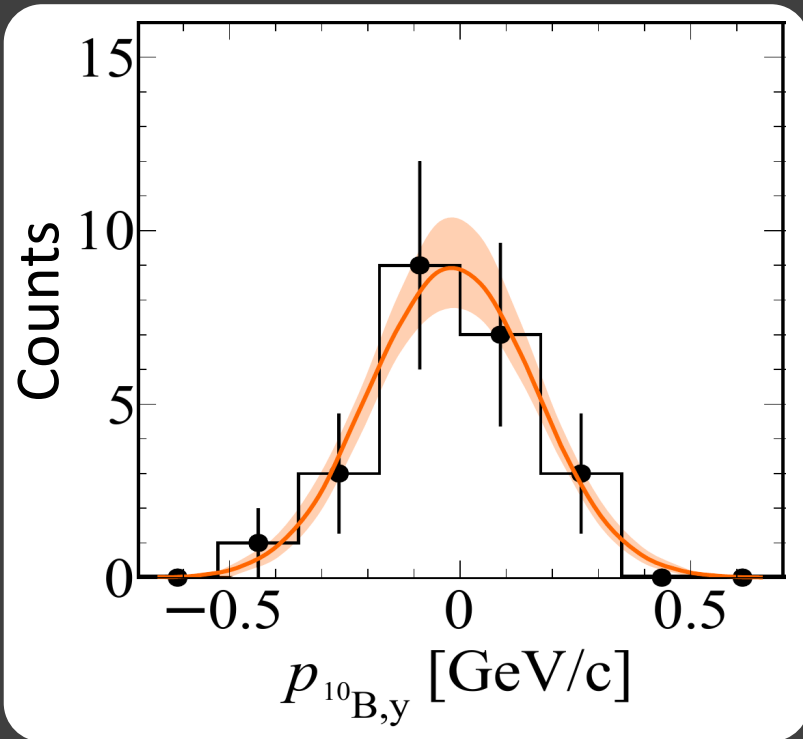
# Inverse Kinematics Measurement





# Pair Motion from $^{10}\text{B}$

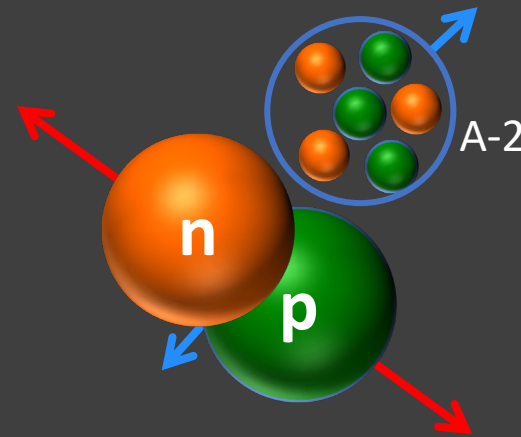
(Agree \w JLab data)



Patsyuk and Kahlbow et al.,  
 Nature Physics 17, 693 (2021)

Experimental Signature: no correlation  
between relative & c.m. motions

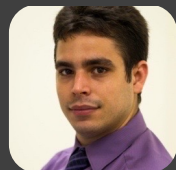
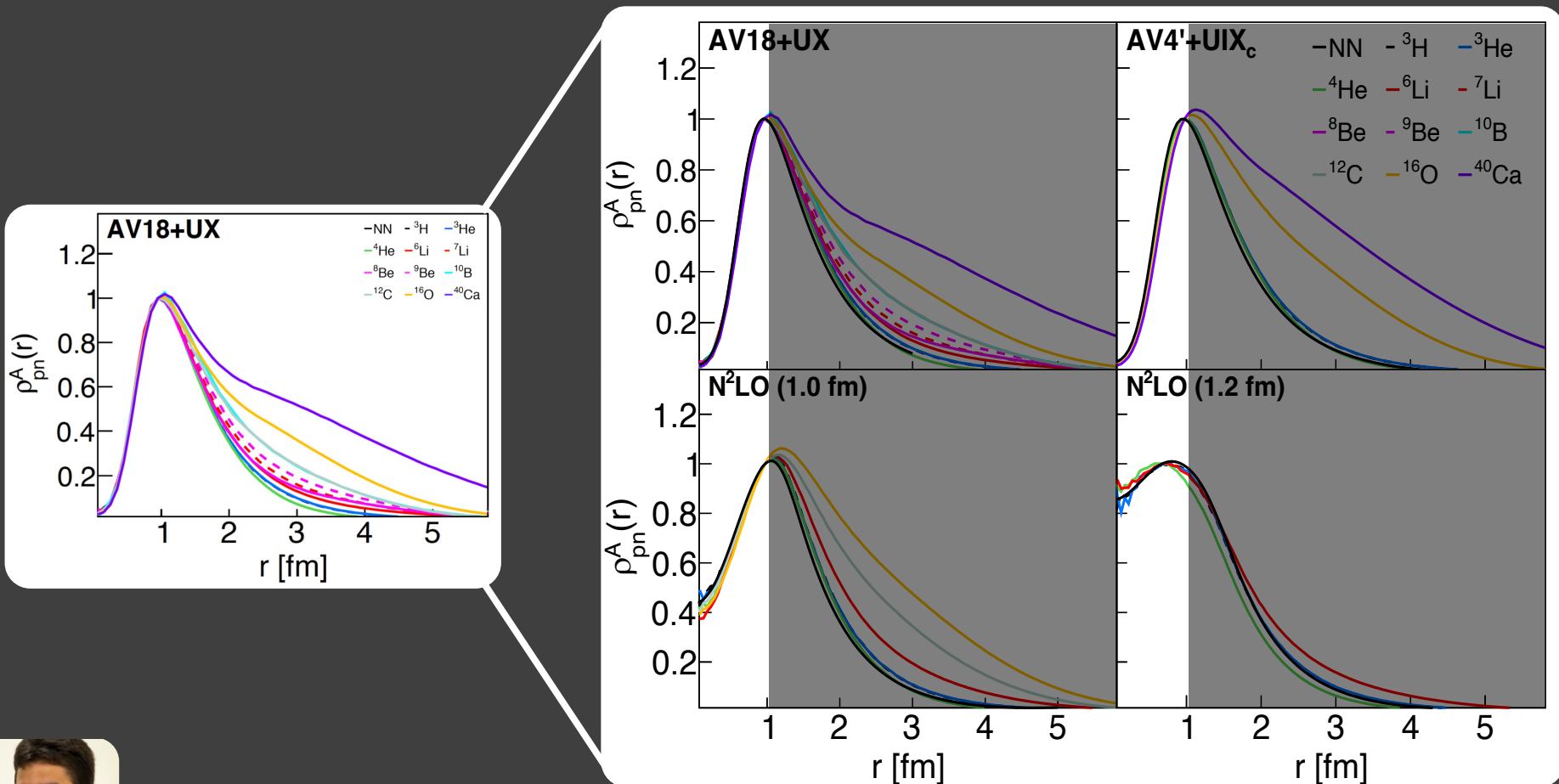
$$f(p_{rel}, p_{c.m.}, \theta_{rel,c.m.}) \approx C(p_{c.m.}) \times \varphi(p_{rel})$$





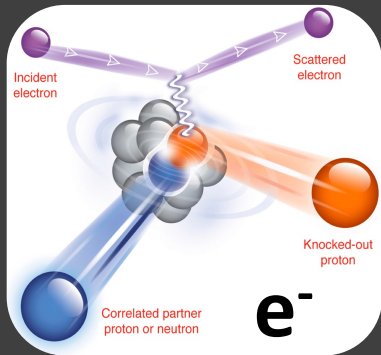
# How model dependent is all of this? i.e., changing scale / scheme changes conclusions?

Short-distance factorization is model independent!

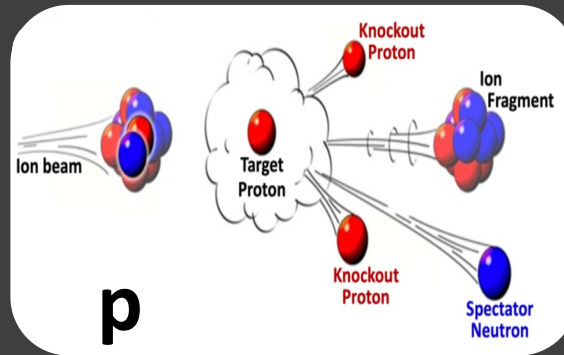


Cruz Torres et al.,  
Nature Physics (2021)

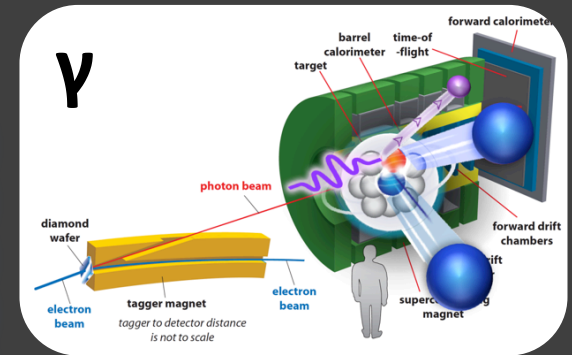
# Results Independent of probe type



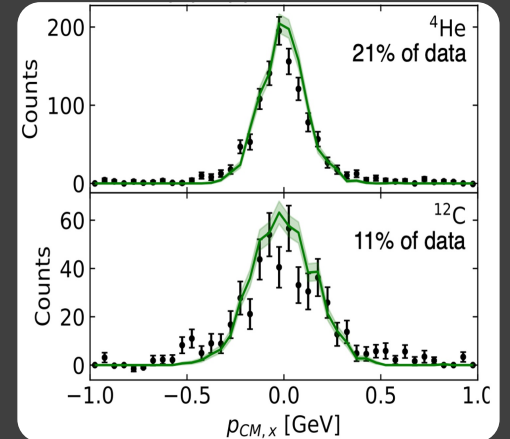
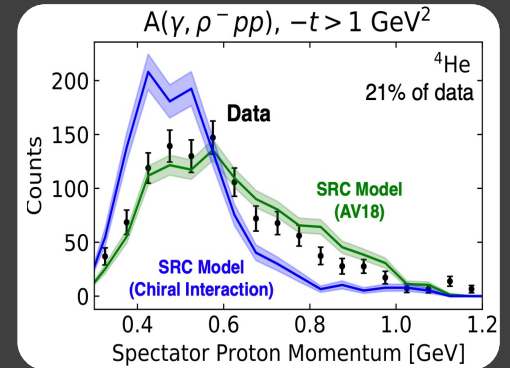
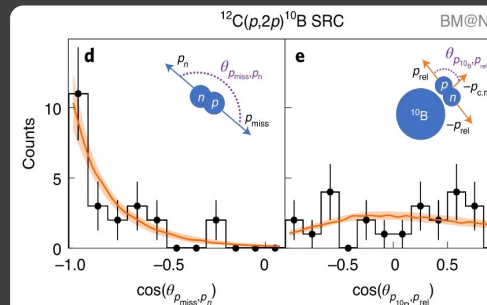
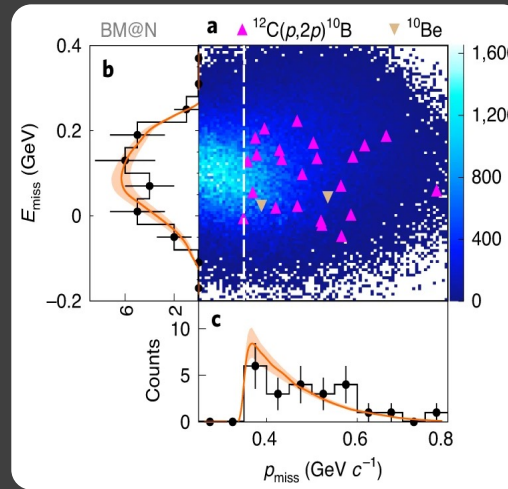
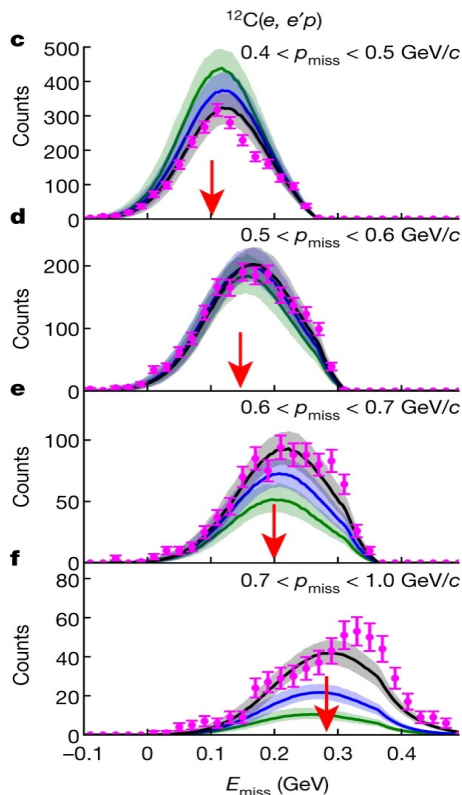
Nature '20



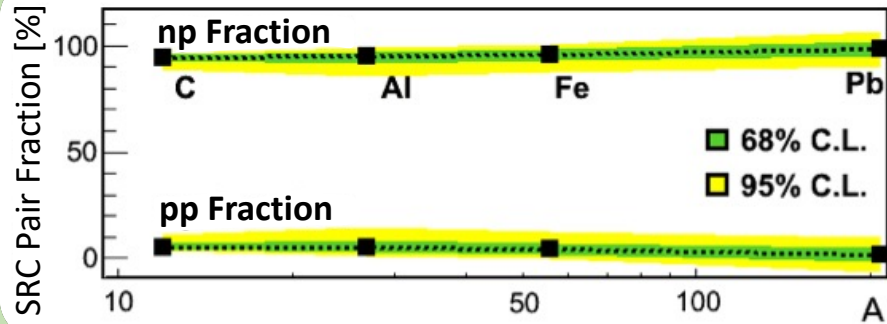
Nature Physics '21



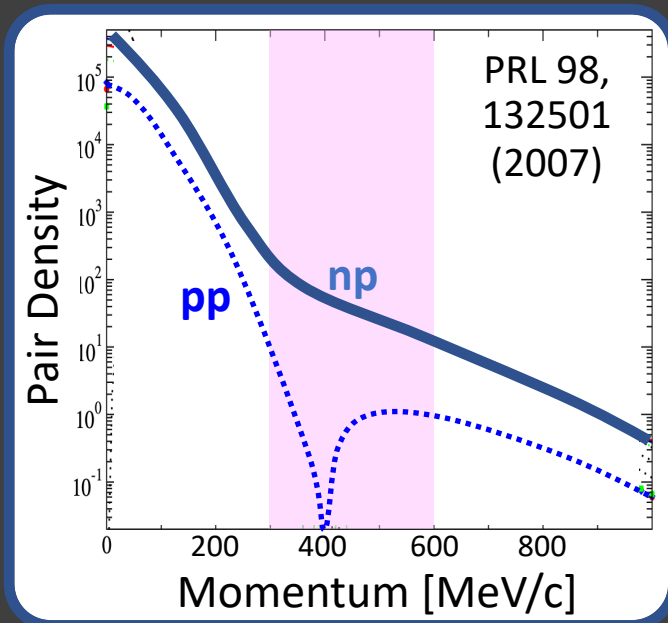
Brand new data!



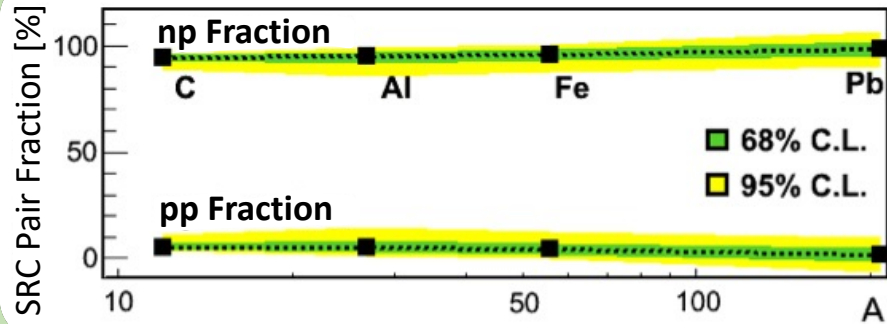
# Short-Ranged Interactions



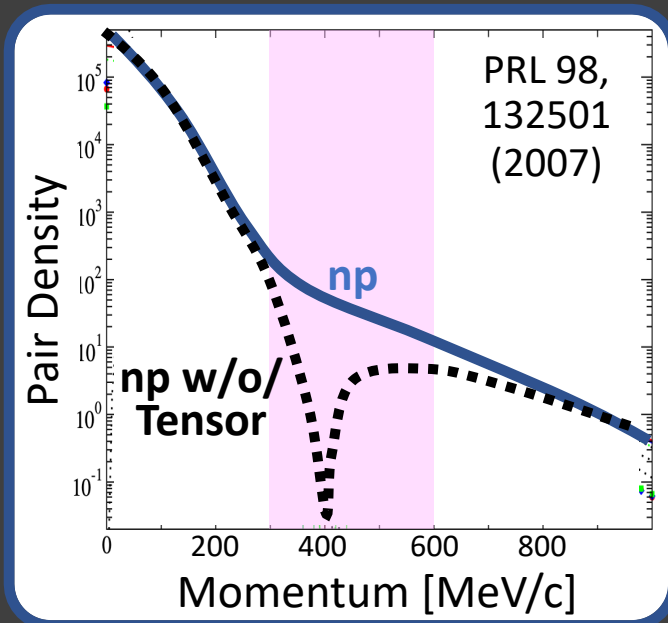
np pairs = Tensor force dominance (spin-dependent)



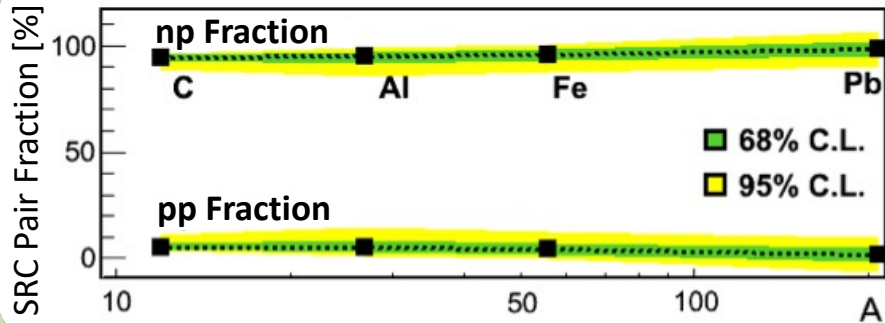
# Short-Ranged Interactions



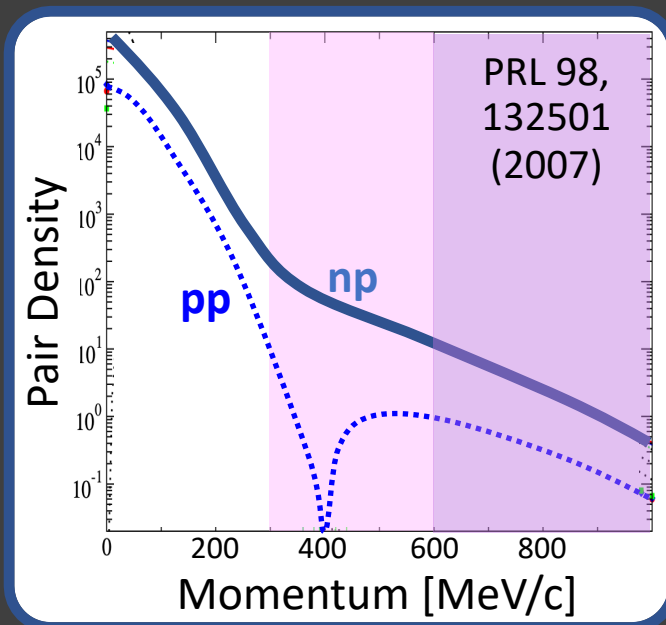
np pairs = Tensor force dominance (spin-dependent)



# Short-Ranged Interactions

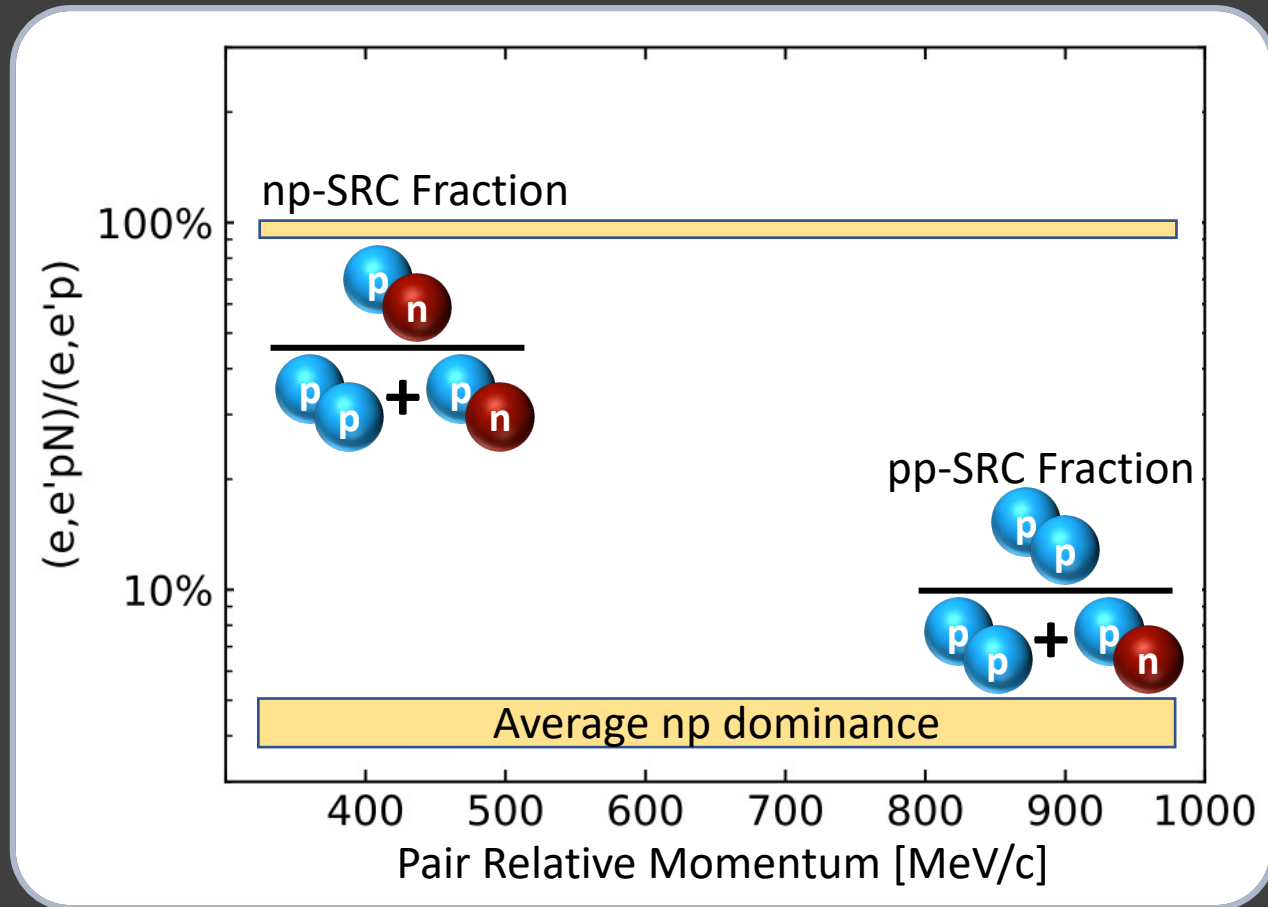


np pairs = Tensor force dominance (spin-dependent)



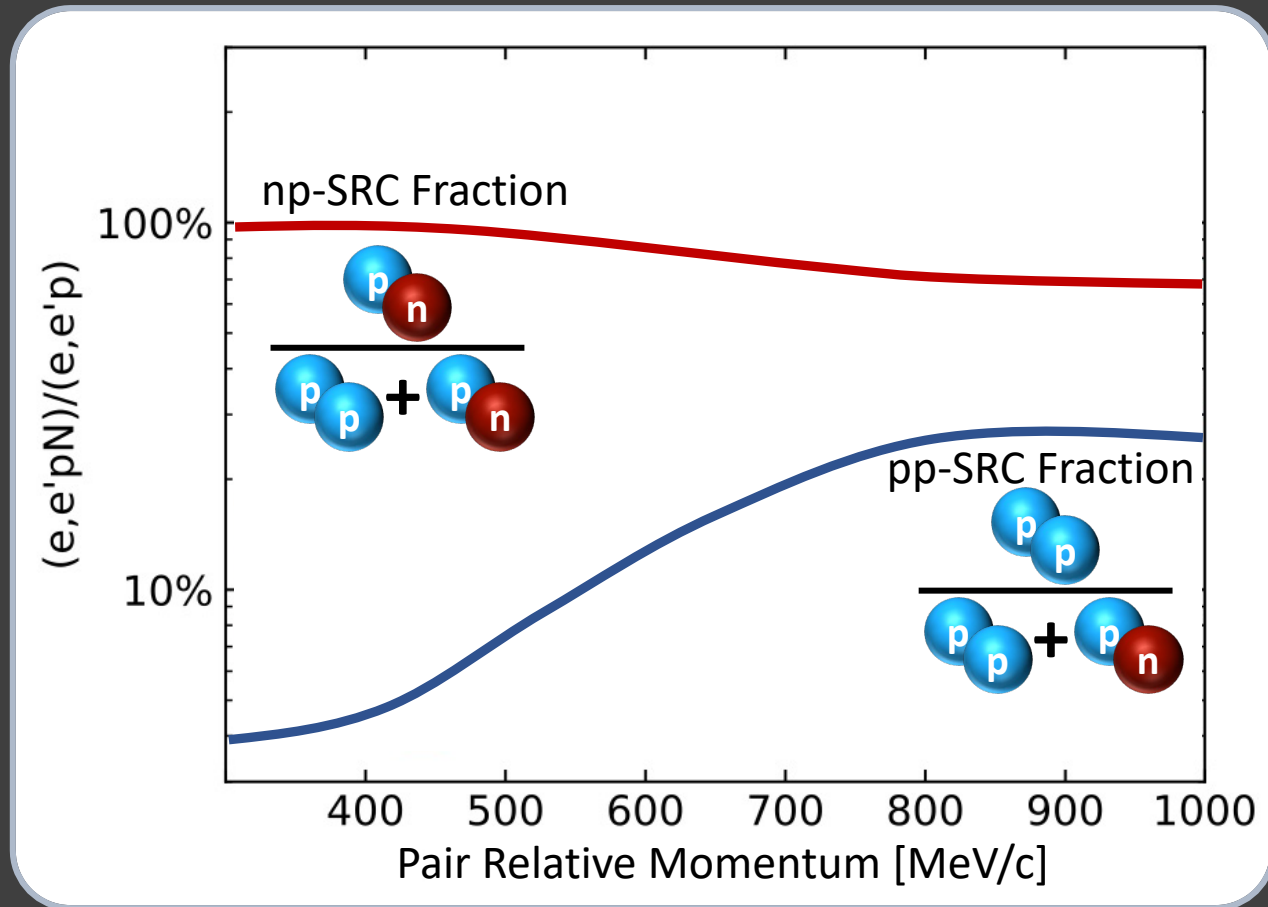
Repulsive core transition:  
Scalar (spin-independent) core produces more pp pairs

# Probing the Repulsive Core



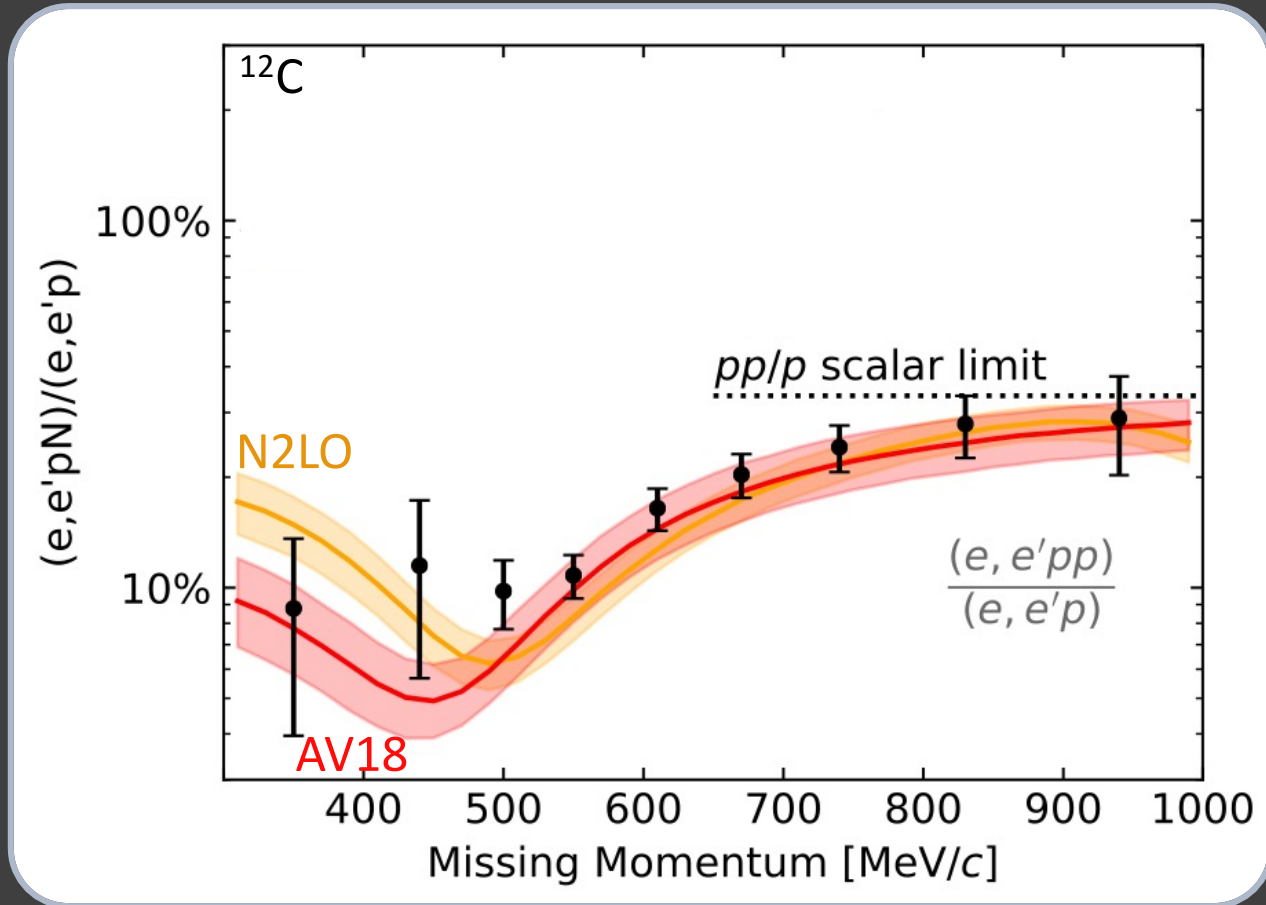
Schmidt and Pybus et al., Nature (2020)  
Pybus et al., PLB (2020);  
Korover and Pybus et al., PLB (2021)

# Probing the Repulsive Core



Schmidt and Pybus et al., Nature (2020)  
Pybus et al., PLB (2020);  
Korover and Pybus et al., PLB (2021)

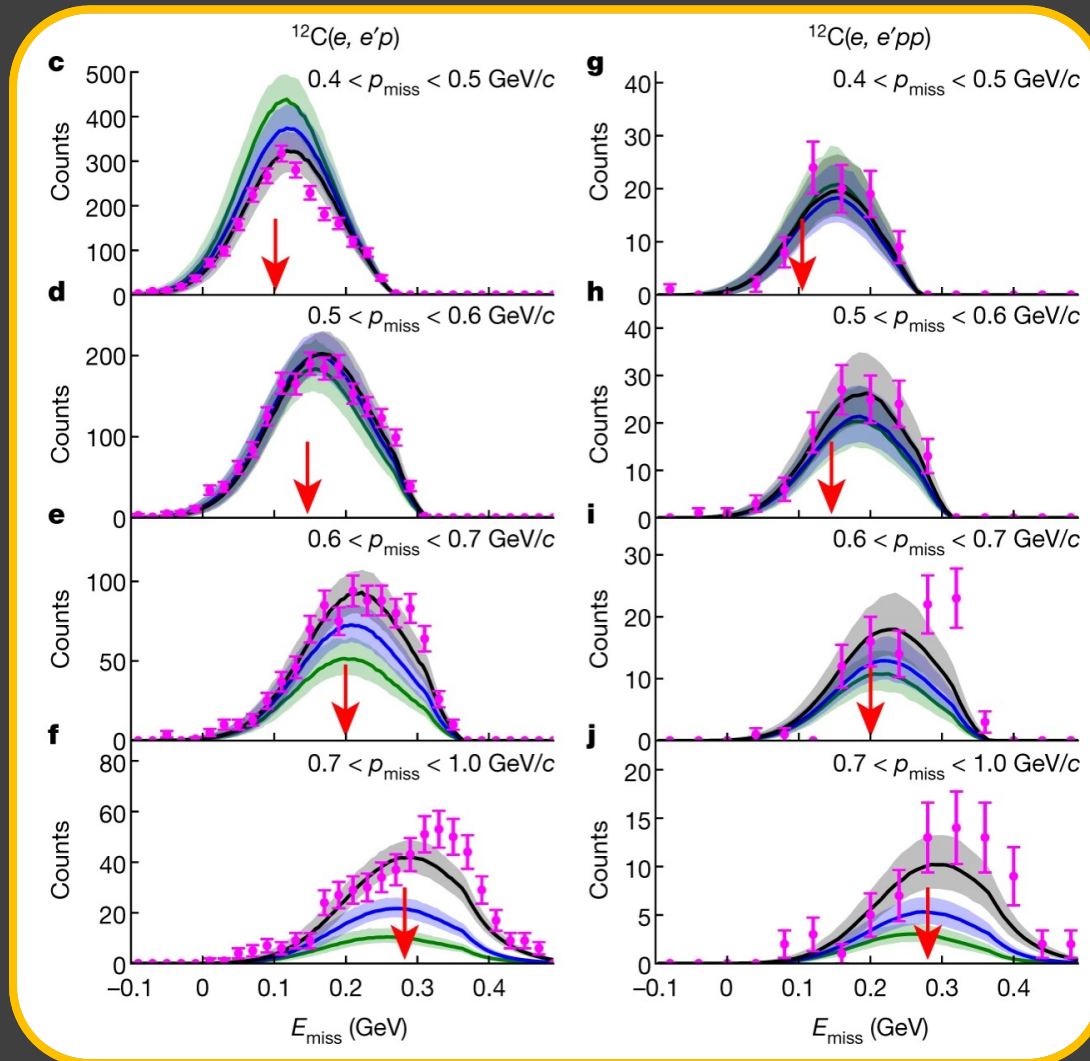
# Probing the Repulsive Core



Schmidt and Pybus et al., Nature (2020)  
Pybus et al., PLB (2020);  
Korover and Pybus et al., PLB (2021)



# First Spectral Function Mapping



Momentum

400 – 500 MeV/c

500 – 600 MeV/c

600 – 700 MeV/c

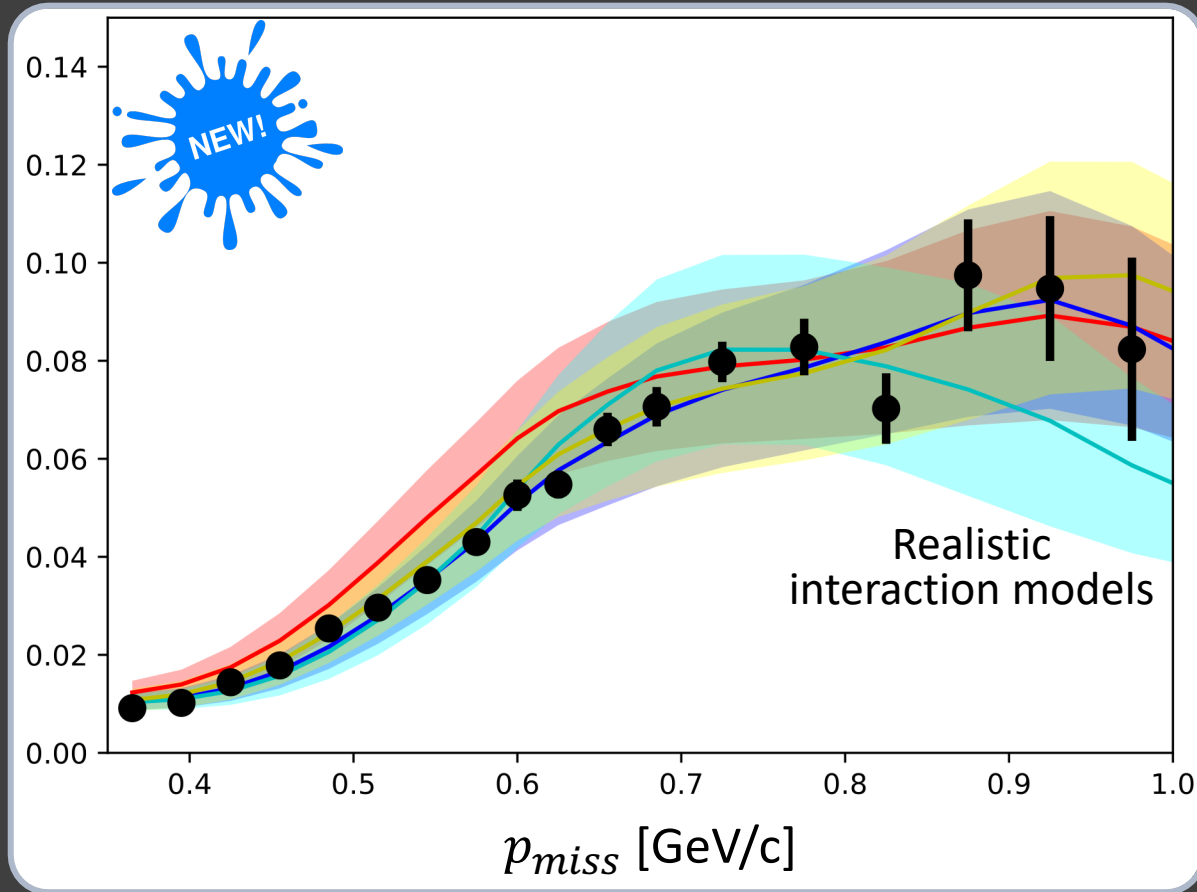
700 – 1000 MeV/c

Energy

Schmidt and Pybus  
et al., Nature (2020)

# New: High-Precision Data

$$\frac{{}^4\text{He}(e, e'pp)}{{}^4\text{He}(e, e'p)}$$



Nature 2020: ~500 pp-SRC events

New data: x20 higher stat (4,000 events used above)!

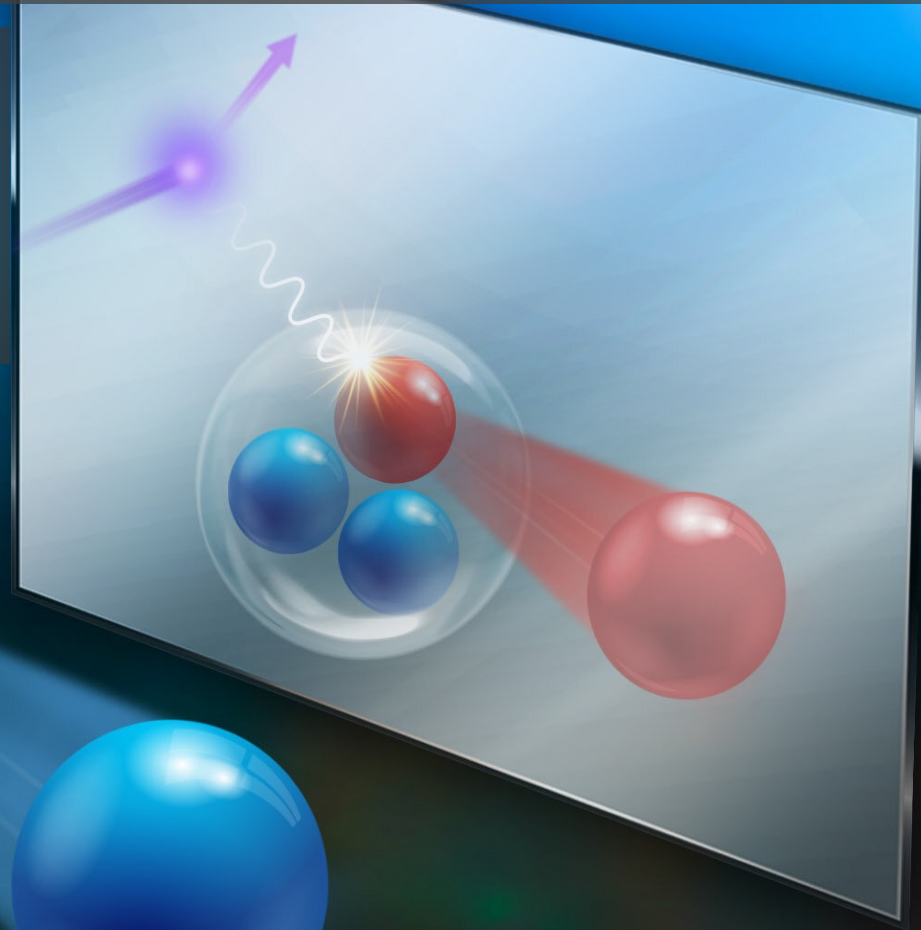
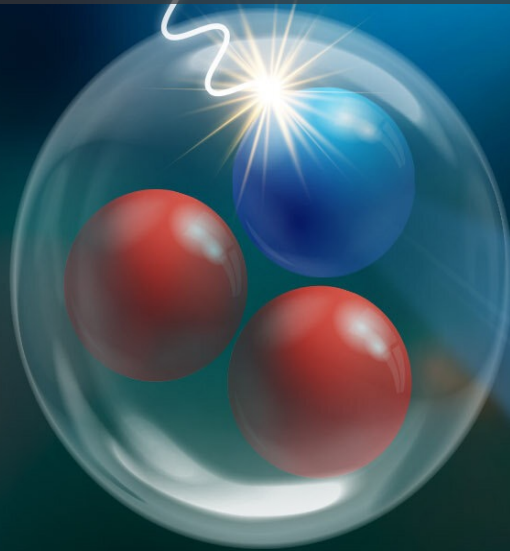


# Test With Tritium & Helium-3

Exactly calculatable 😊

Mirror nuclei:

Tritium  $p =$  Helium-3  $n$

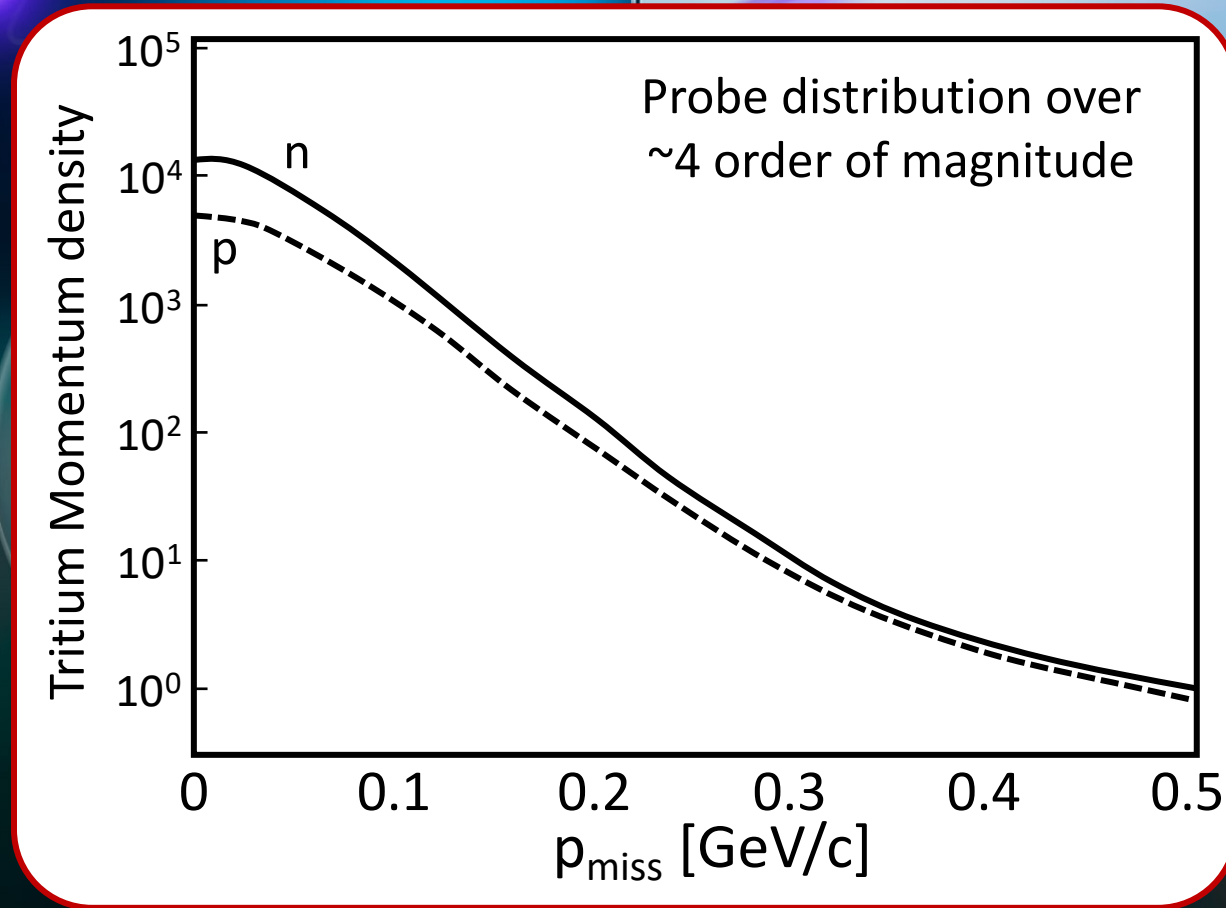


## Editors' Suggestion

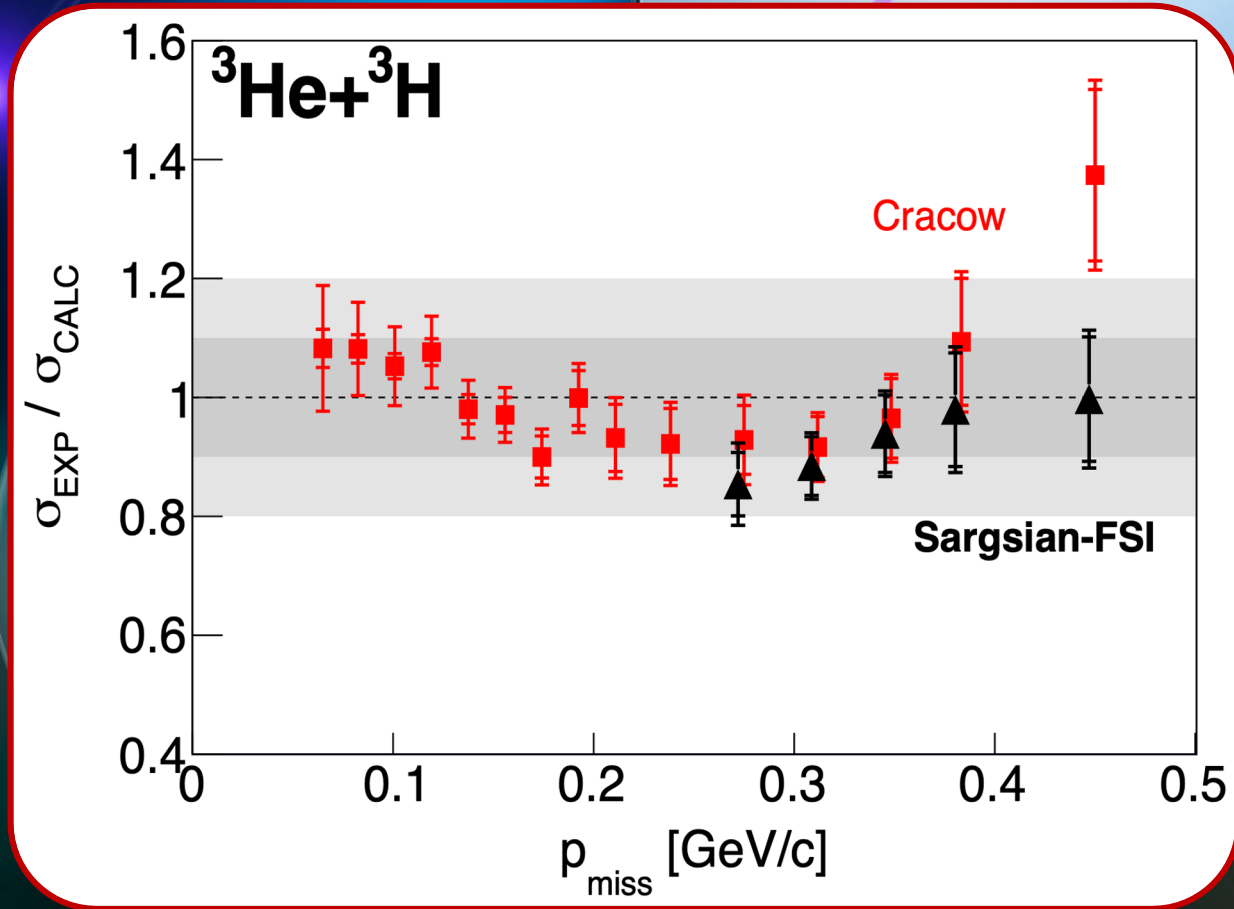
Cruz Torres and Nguyen  
et al., Phys. Rev. Lett (2020)

# Challenges: Tritium Radioactivity

## Low High-p States Density



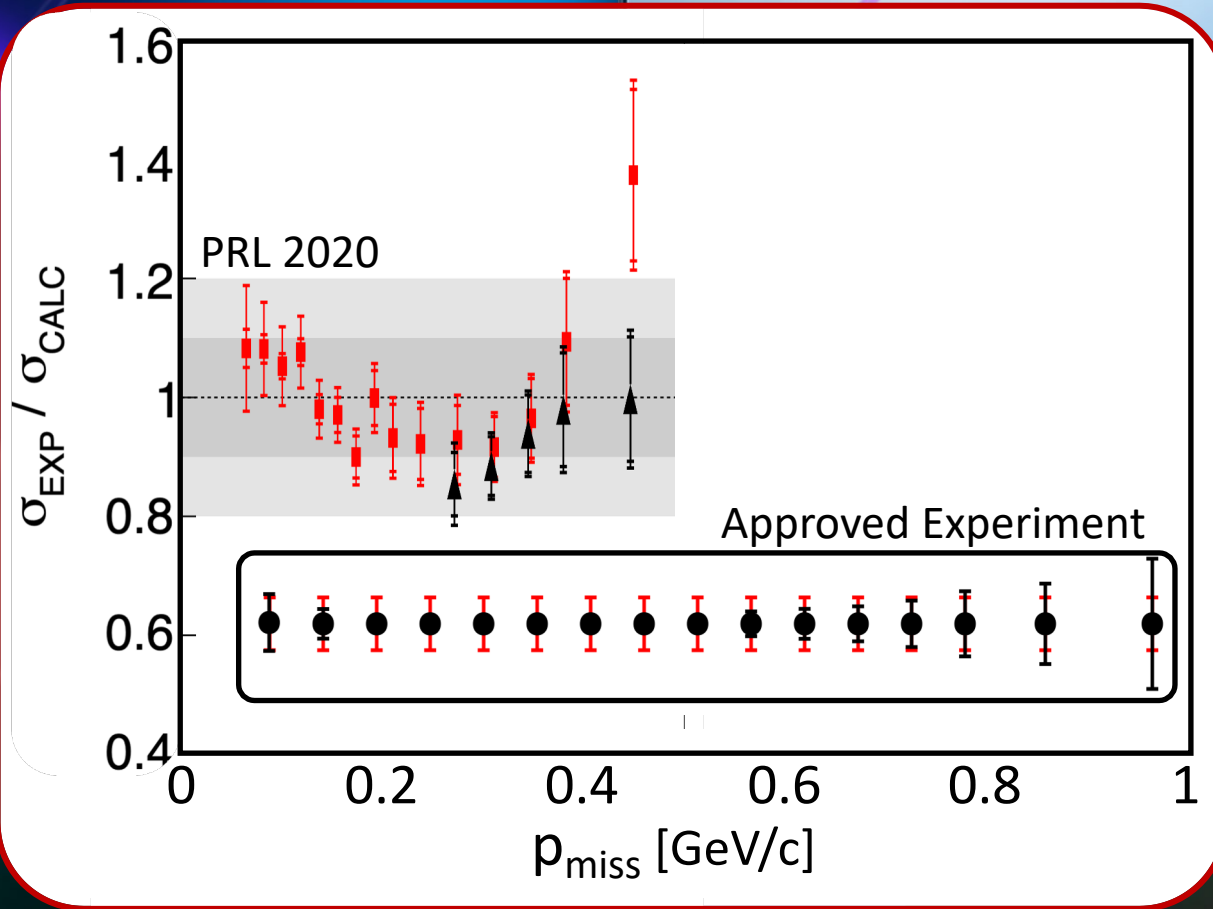
# ✓ Agreement Over 4 Orders of Magnitude



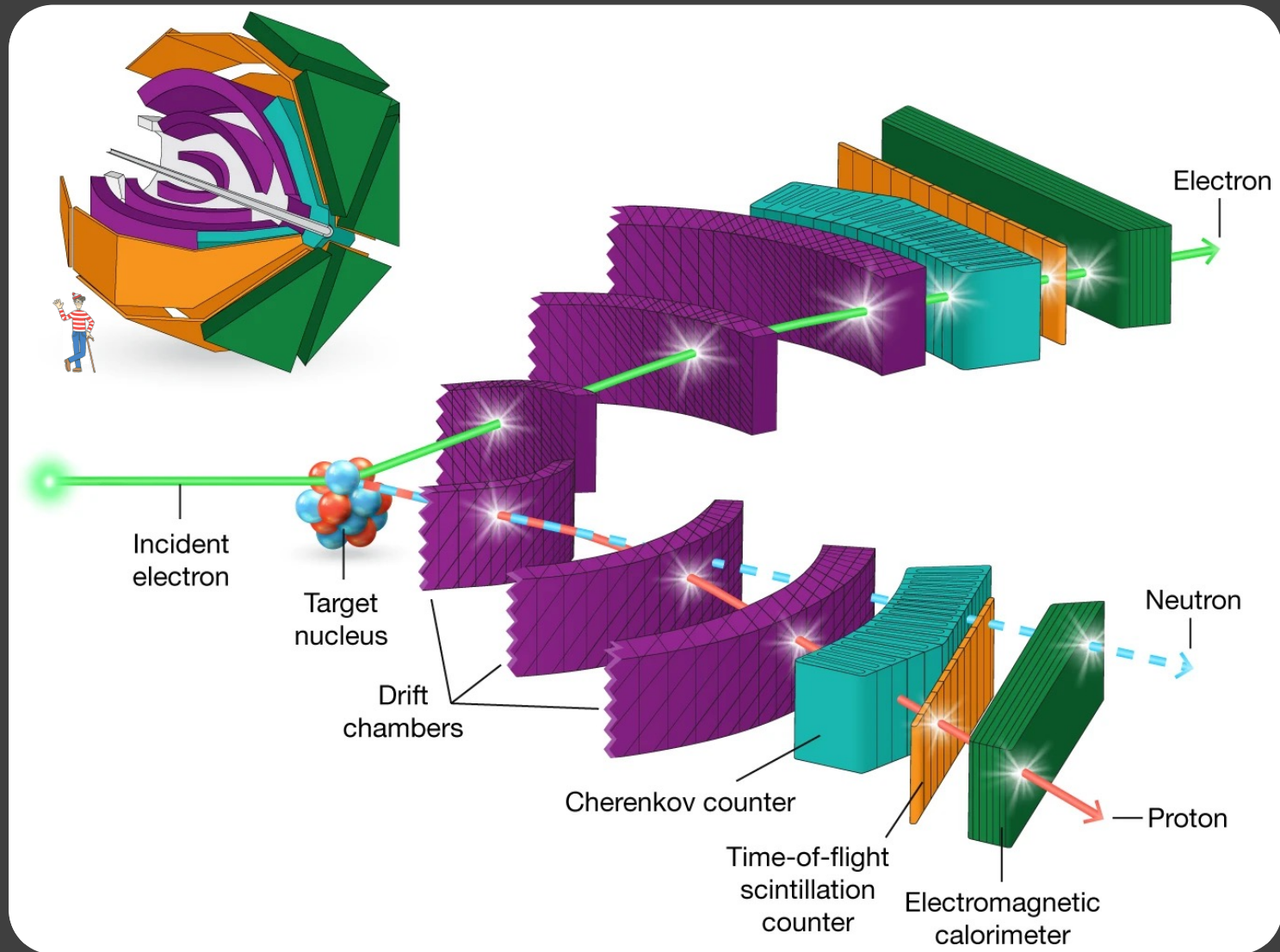
Editors' Suggestion

Cruz Torres and Nguyen  
et al., Phys. Rev. Lett (2020)

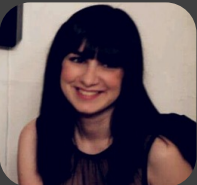
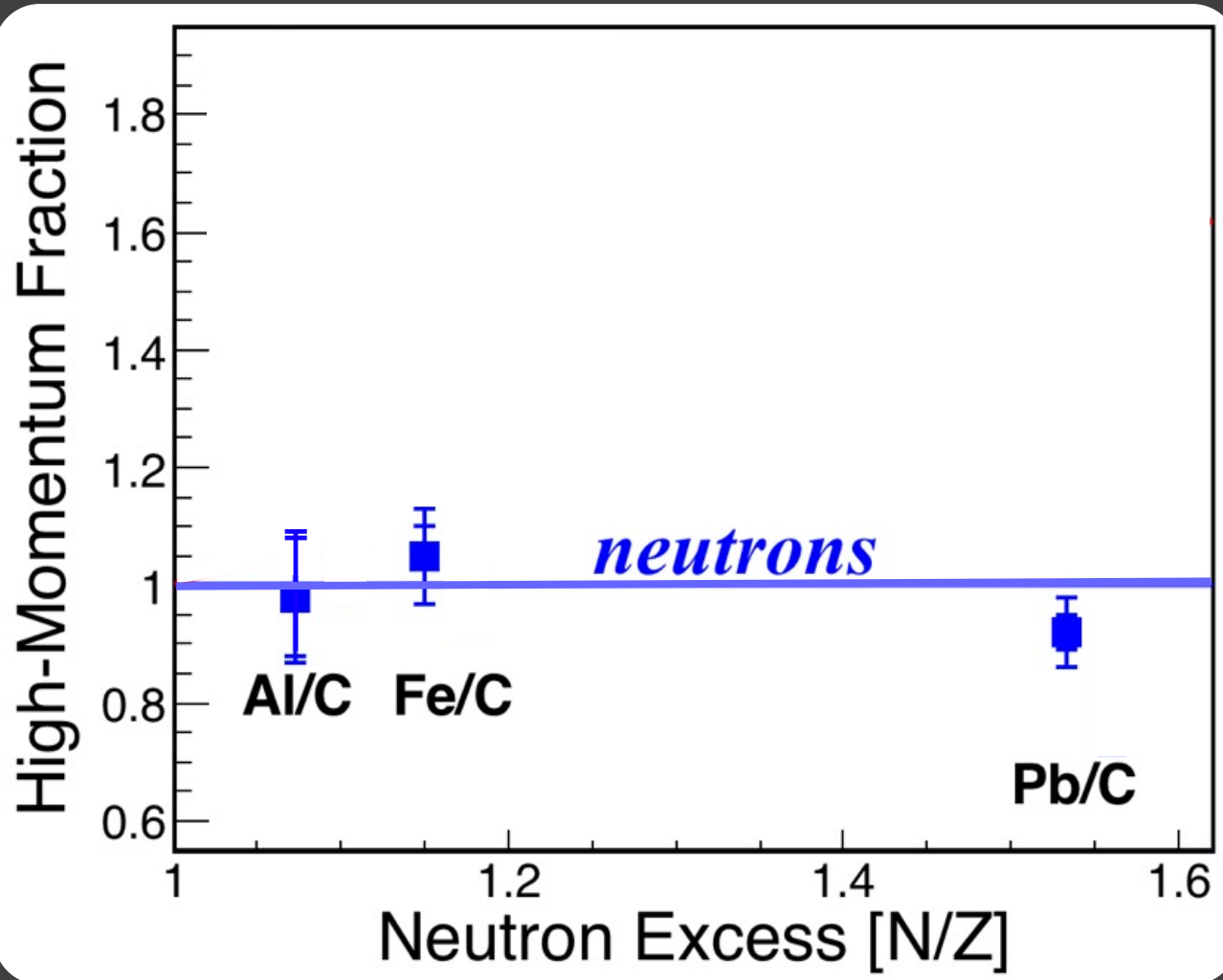
# Next Step: Double the Reach



# Comparing Proton & Neutron Dynamics

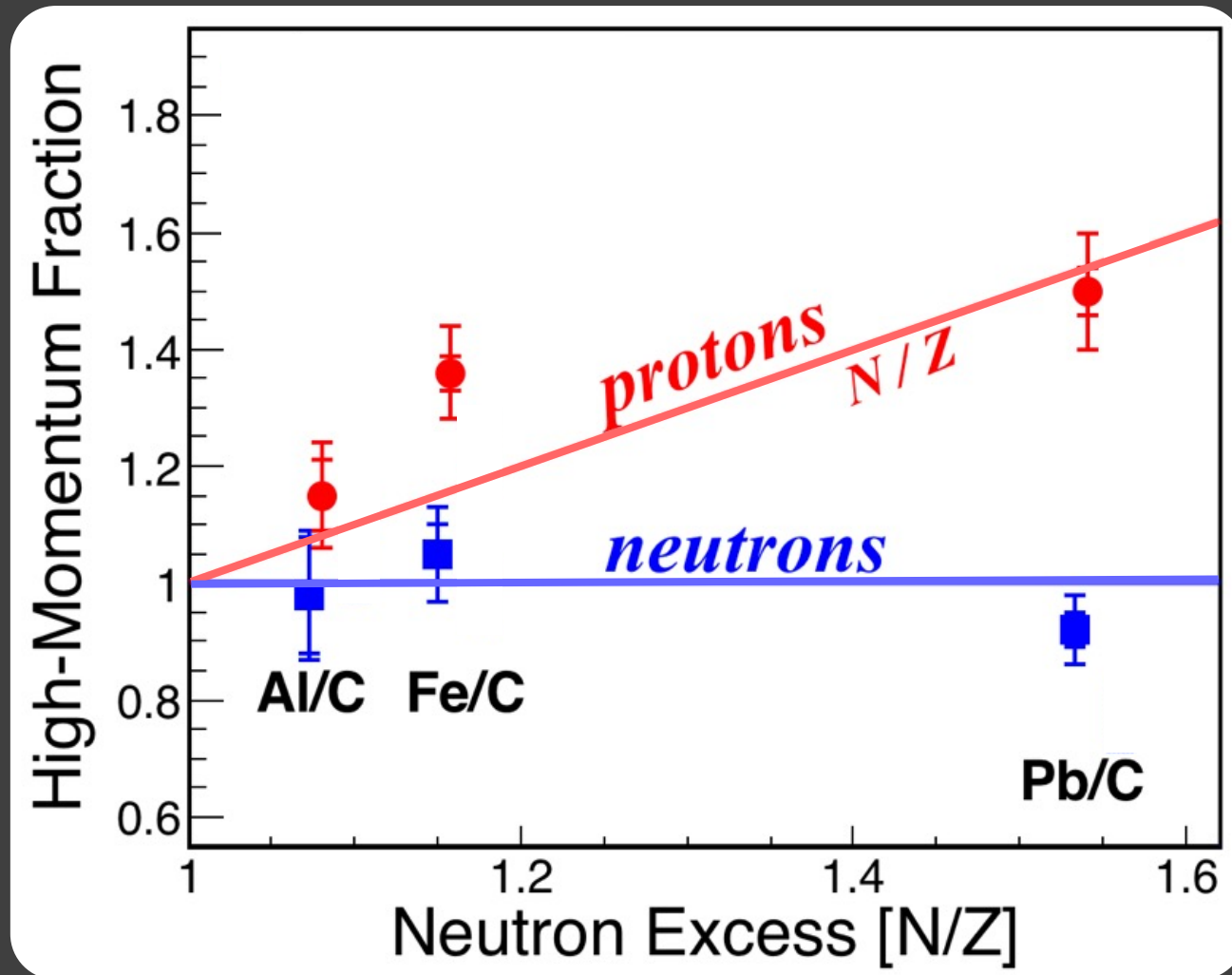


# Correlation Probability: Neutrons saturate

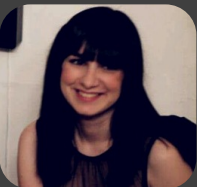
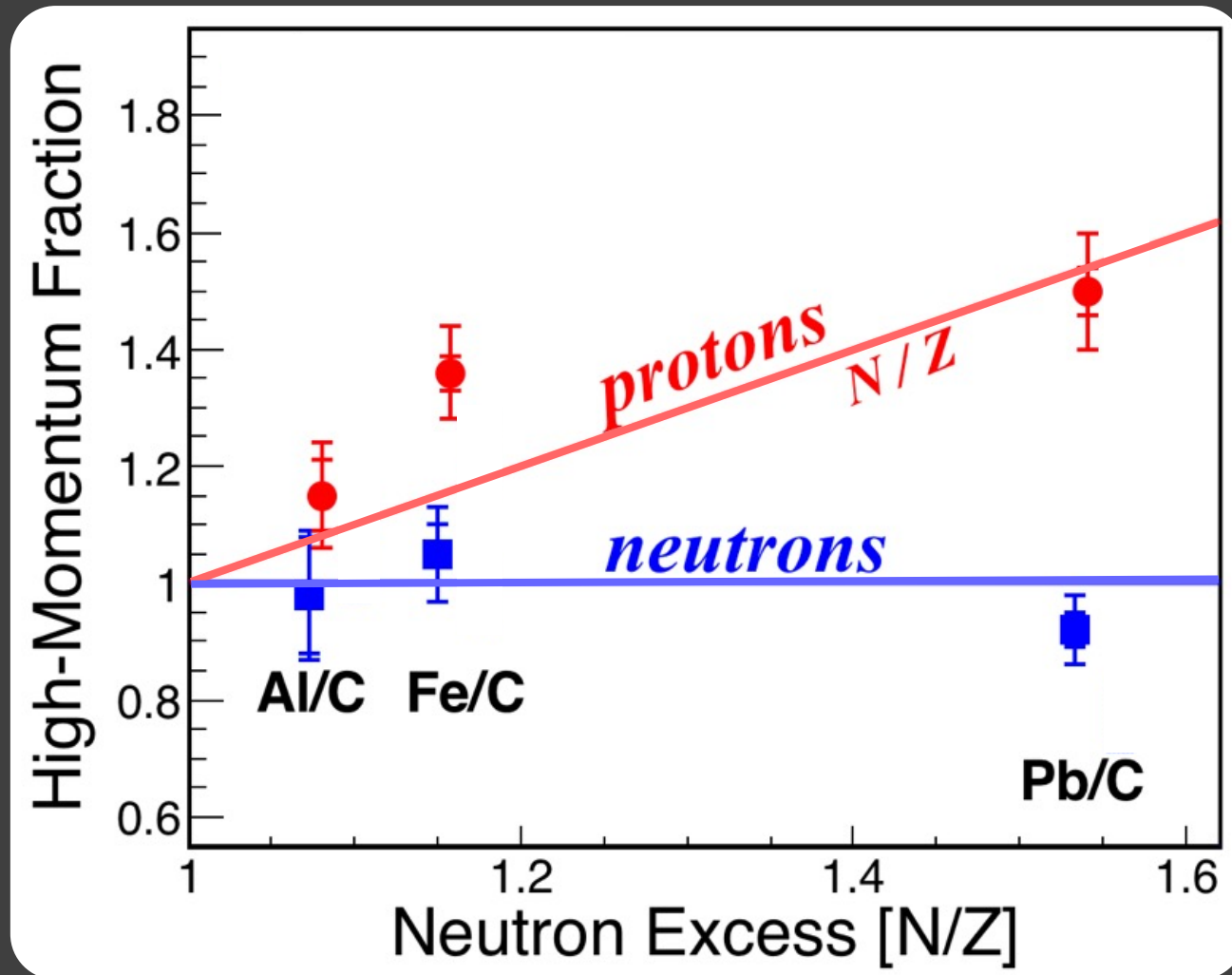




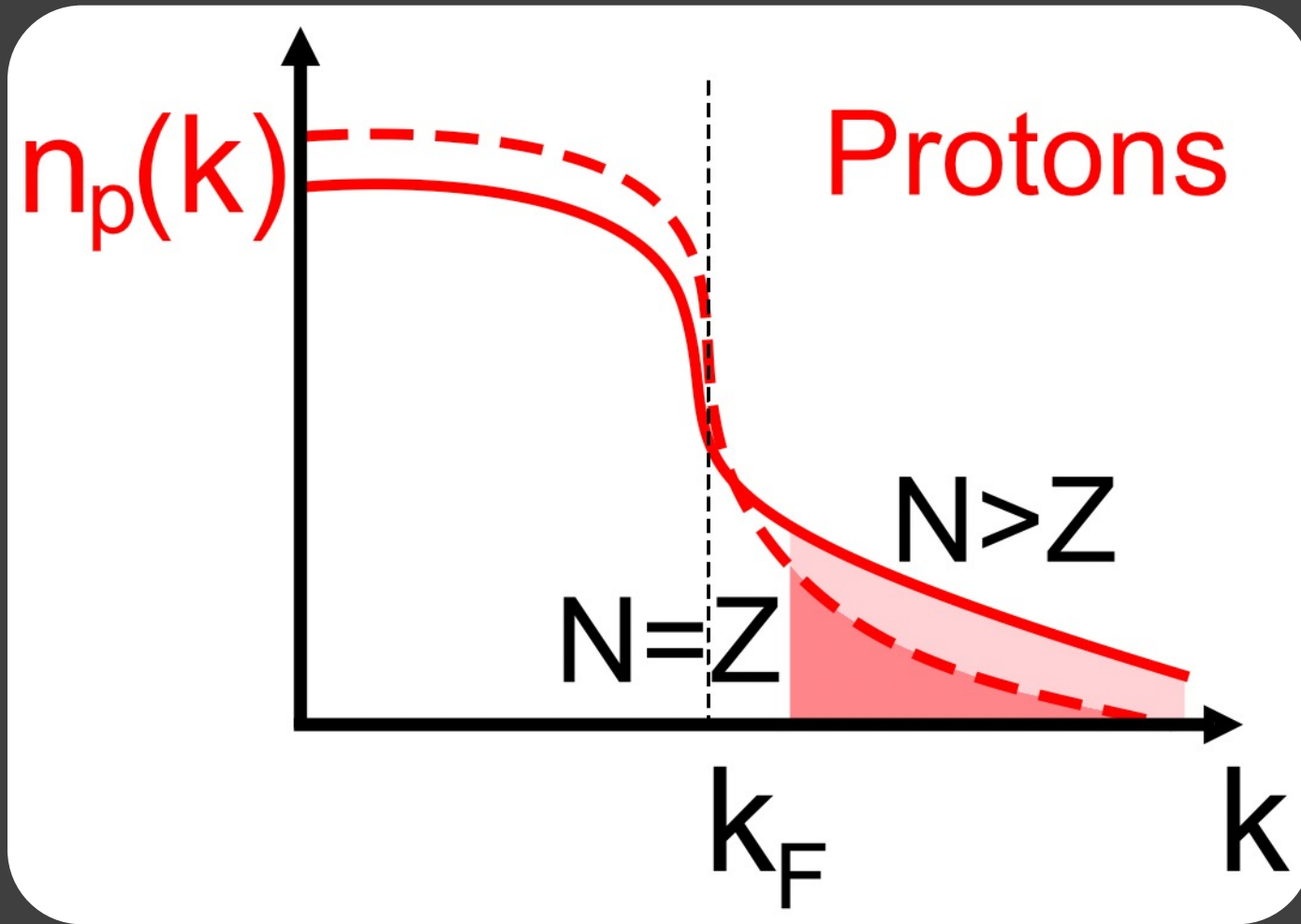
# Correlation Probability: Neutrons saturate **Protons grow**



# Protons 'Speed-Up' In Neutron-Rich Nuclei

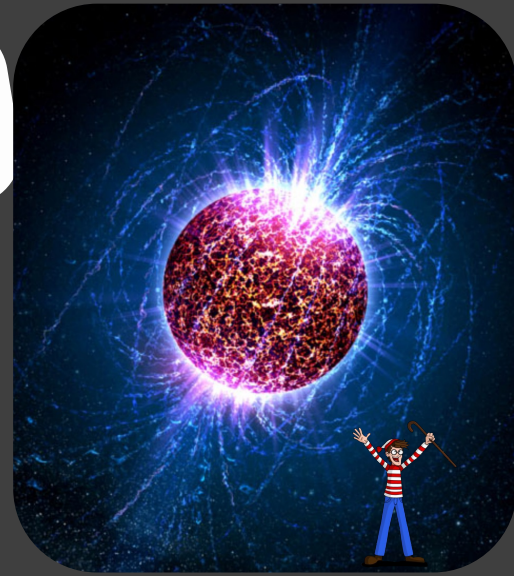


# Protons 'Speed-Up' In Neutron-Rich Nuclei



Daily Press

Jefferson Lab breaks new ground, from nucleons to neutron stars

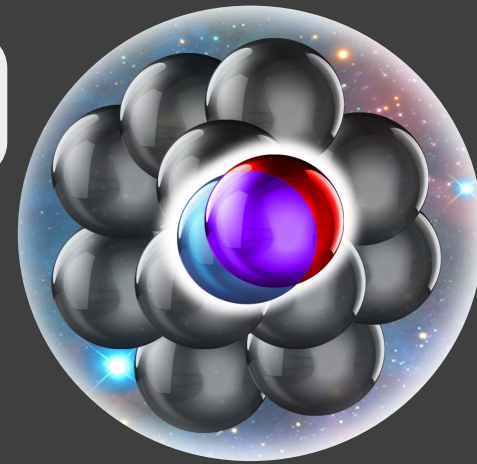


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Protons may have an outsized influence on the properties of neutron stars and other neutron-rich objects



Protons strongly influence the behaviour of neutron stars



 Astronomy Now

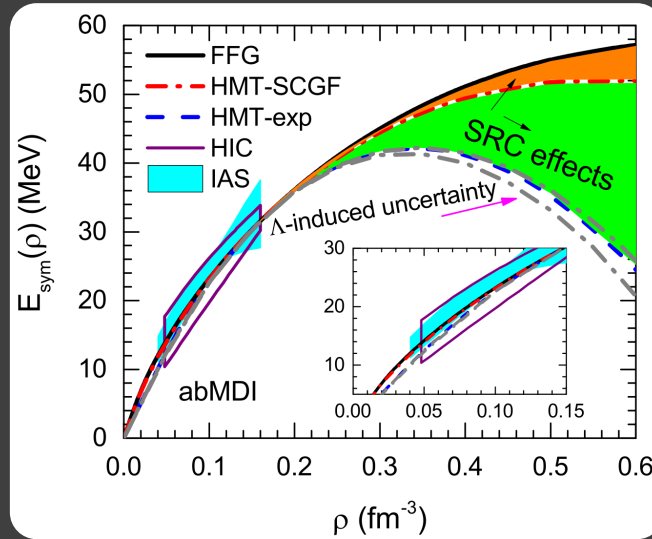
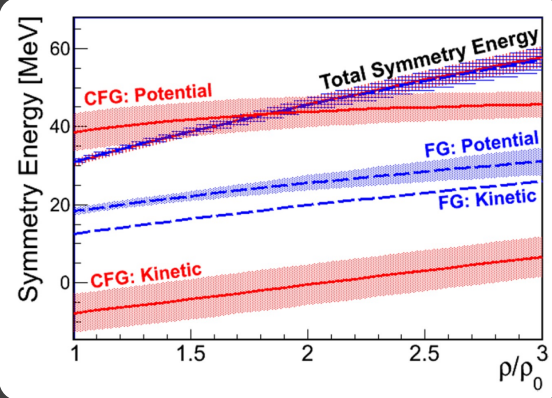
 physicsworld  
Volume 29 No 7 July 2016

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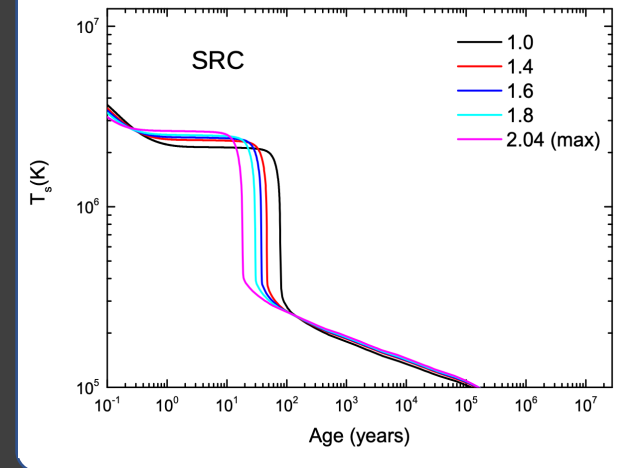
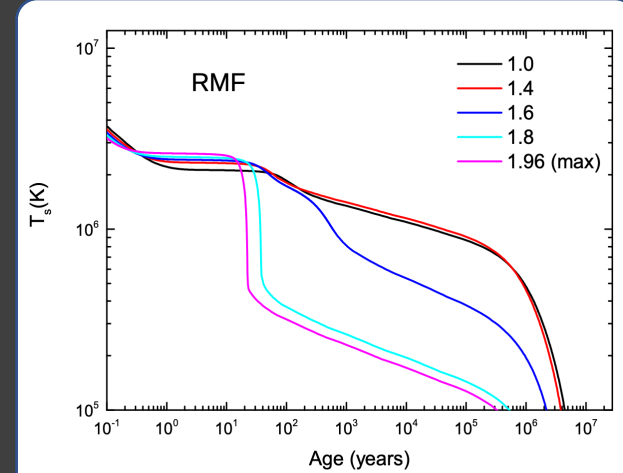
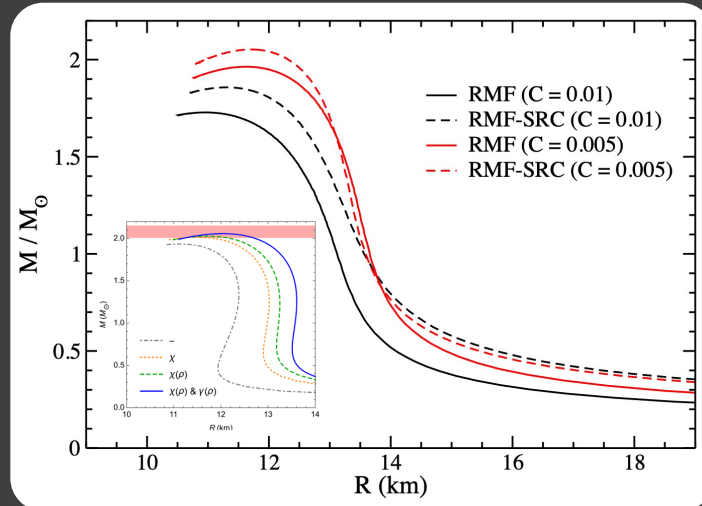
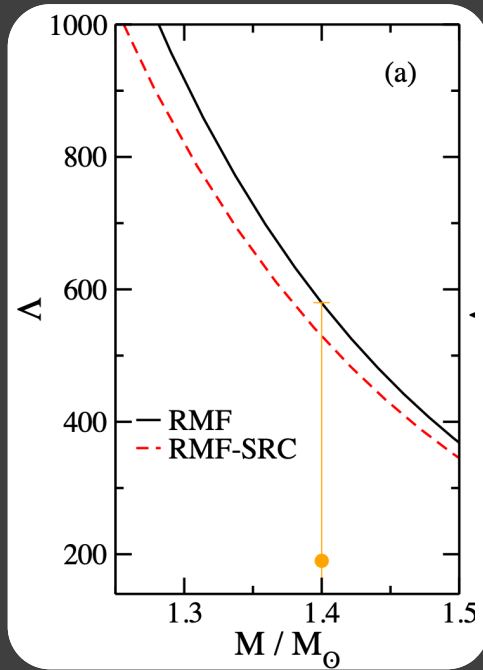
Surprising Accelerator Finding Could Change the Way We Think About Neutron Stars

# Affects neutron star calculations

Hen+, PRC '15 (>100 citations)



Prog. Part. Nucl. Phys. 99, 29 (2018)  
Eur. Phys. J. A 55, 117 (2015)



arXiv 2004.10309

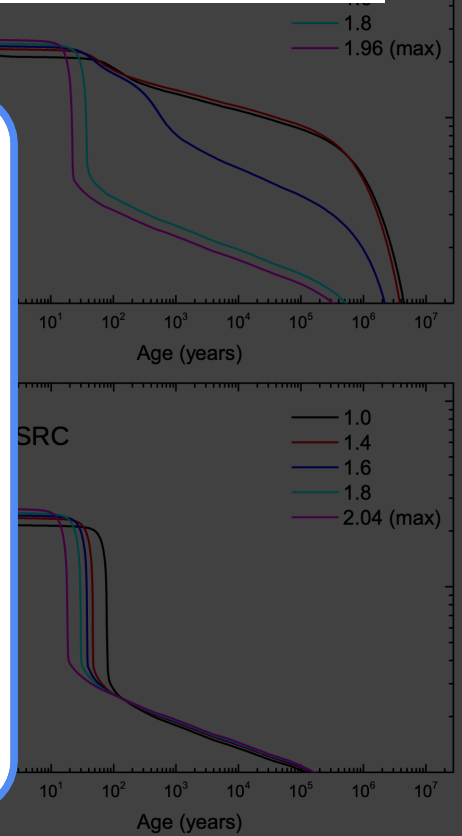
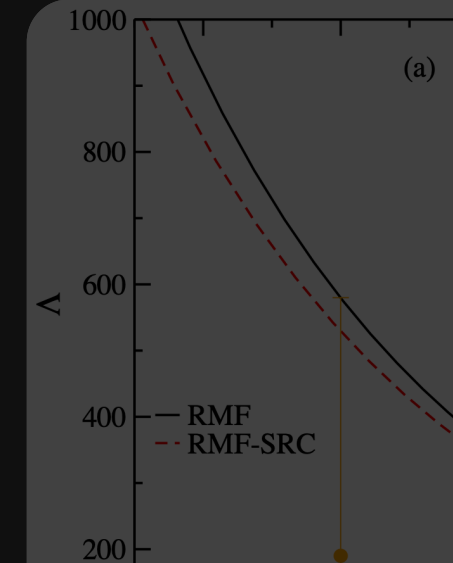
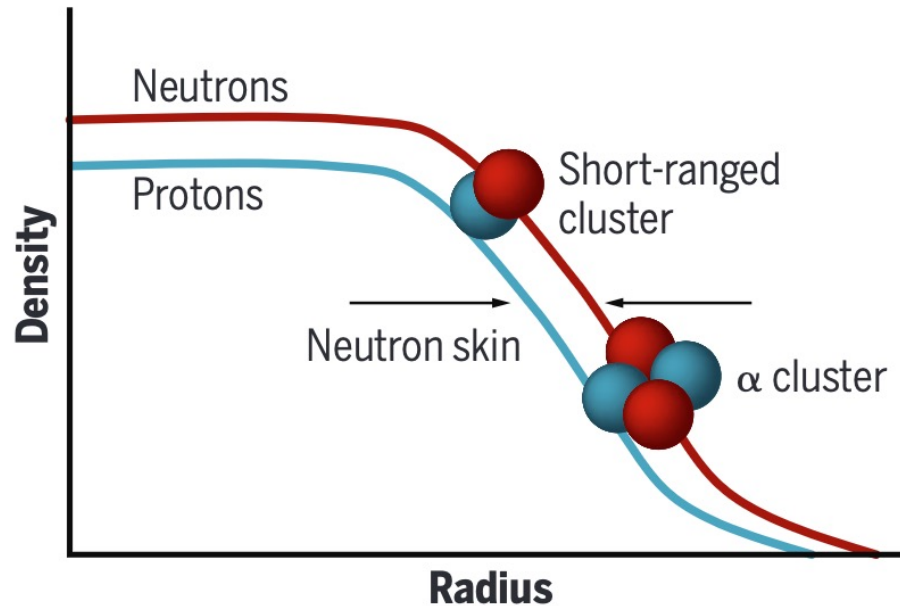
PRC 101, 065202 (2020)

NPA 990, 118 (2019)

NUCLEAR PHYSICS

# From nuclear clusters to neutron stars

## Nucleon density in neutron-rich nuclei



Hen, Science 371, 232 (2021)  
 Tanaka et al., Science 371, 260 (2021)  
 Miller et al., Phys. Lett. B 793, 360 (2019)

arXiv 2004.10309