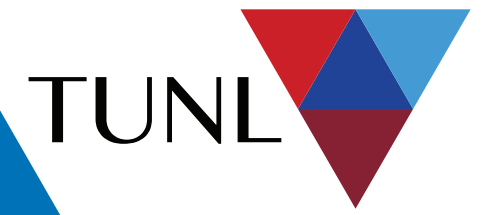


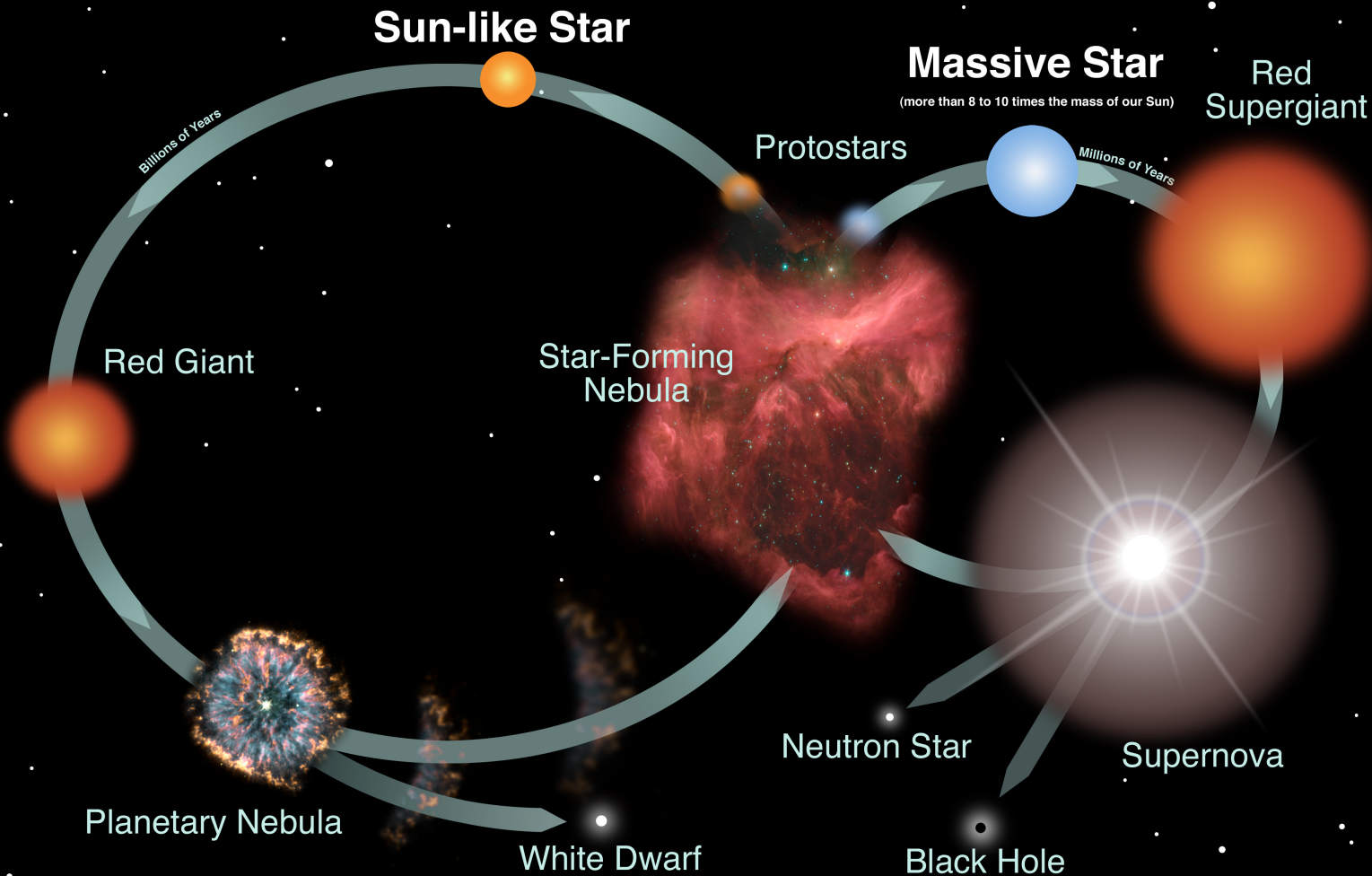
# NUCLEAR PHYSICS DATA FOR ASTROPHYSICS



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Triangle Universities Nuclear Laboratory**

**Town Meeting, November 2022, Argonne National Laboratory**



- all nuclear data are important for nuclear astrophysics

- in future: much more emphasis on uncertainties of nuclear data and derived quantities

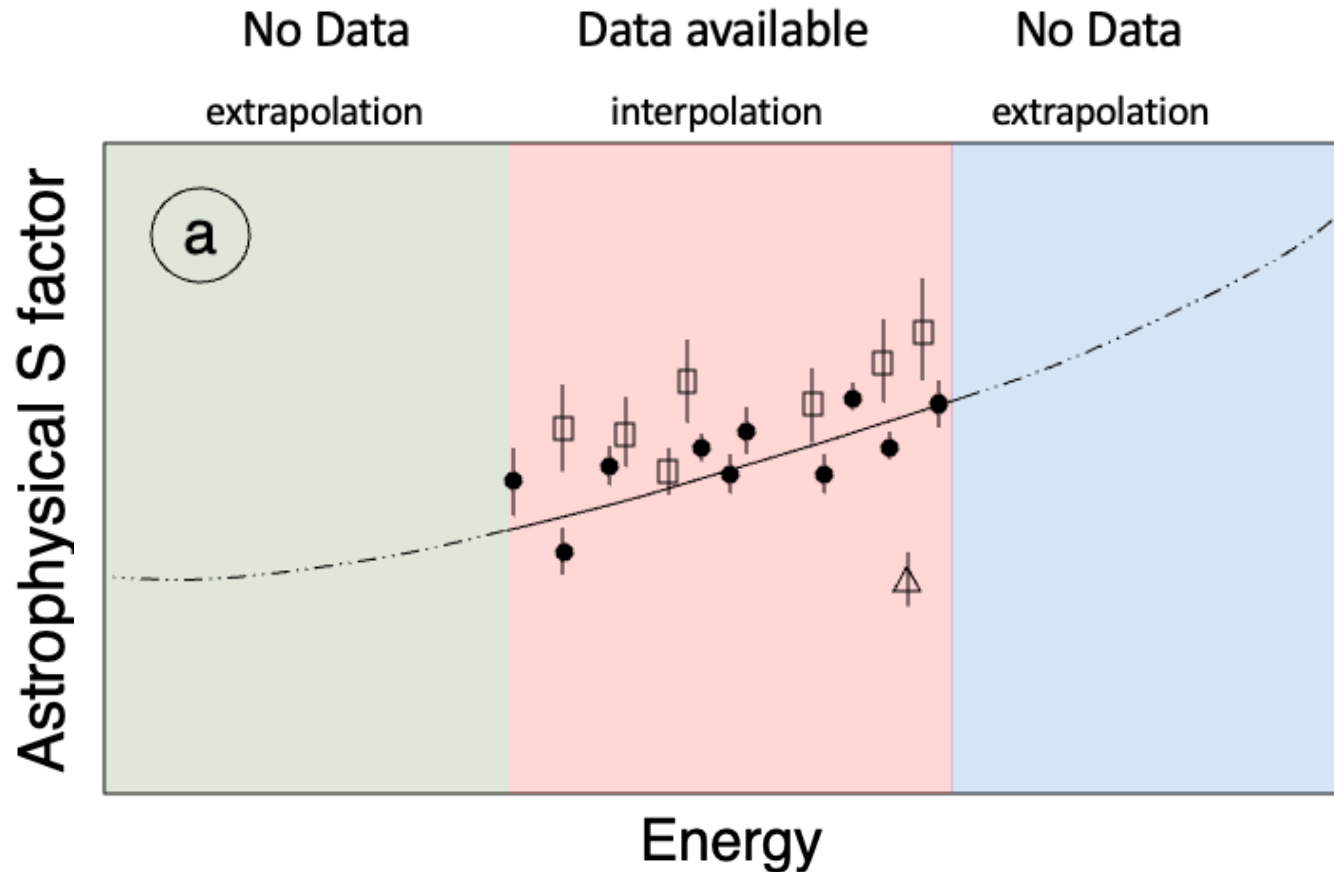
- astrophysical S factor:  

$$\sigma(E) = E^{-1} e^{-2\pi\eta} S(E)$$

- reaction rate:  

$$\langle\sigma v\rangle \sim T^{-3/2} \int E \sigma(E) e^{-E/kT} dE$$





needed:

- measured cross section
- nuclear theory
- statistical [e.g., Bayesian<sup>1,2</sup>] model for fitting

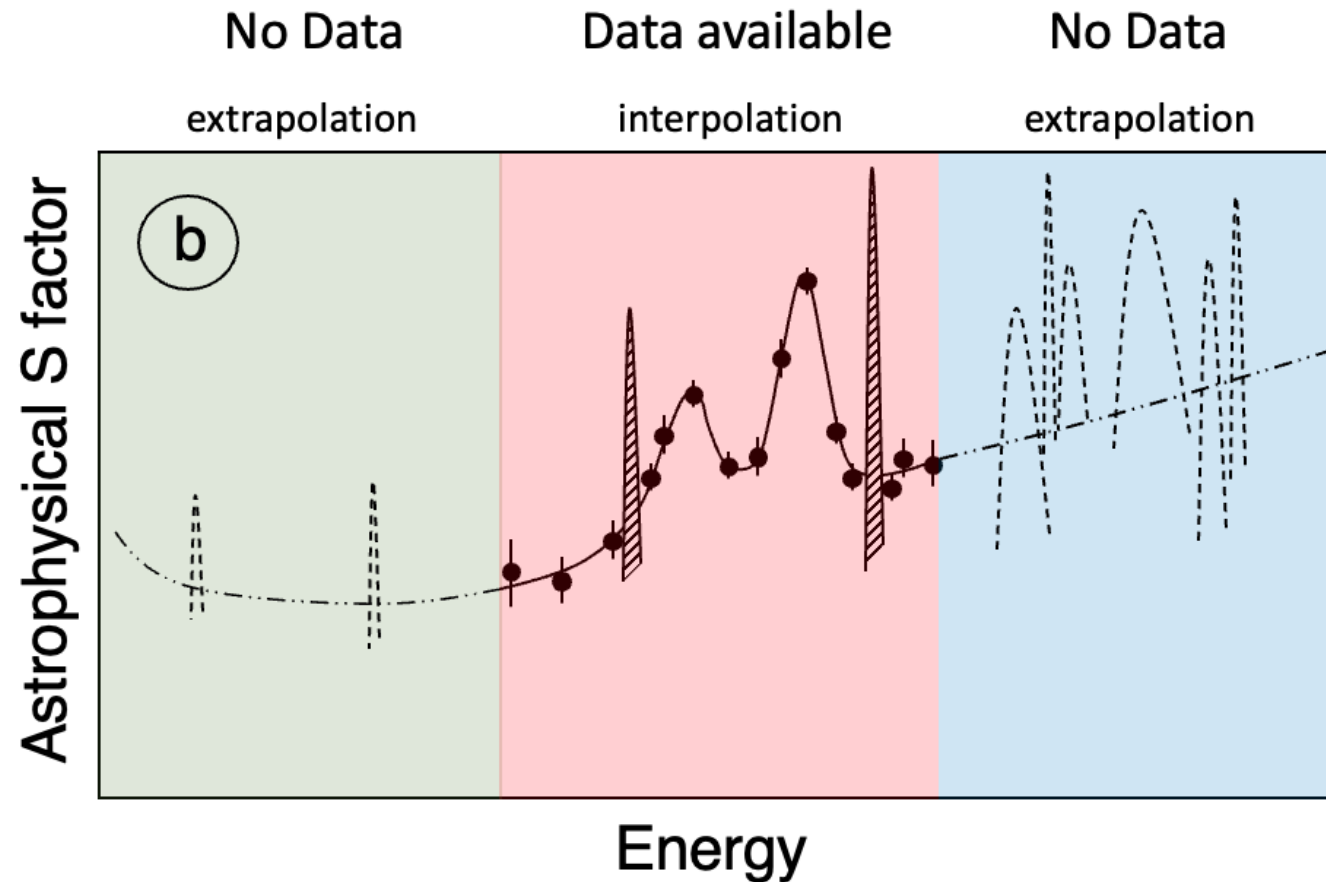
data bases:

- EXFOR
- original literature

examples:  $D(p,\gamma)^3\text{He}$ ,  $D(d,n)^3\text{He}$ ,  $D(d,p)^3\text{H}$ ,  $^{16}\text{O}(p,\gamma)^{17}\text{F}$ ,...

<sup>1</sup> Zhang et al., PLB 751, 535 (2015)

<sup>2</sup> Iliadis et al., ApJ 831, 107 (2016)



examples:  $^{14}\text{N}(p,\gamma)^{15}\text{O}$ ,  $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ ,  $^{17}\text{O}(p,\gamma)^{18}\text{F}$ ,  $^{23}\text{Na}(p,\gamma)^{24}\text{Mg}$ ...

<sup>1</sup> de Souza et al., PRC 99, 014619 (2019)    <sup>2</sup> Odell et al., PRC 105, 014625 (2022)

<sup>3</sup> Longland et al., NPA 841, 1 (2010)    <sup>4</sup> Endt, NPA 521, 1 (1990)

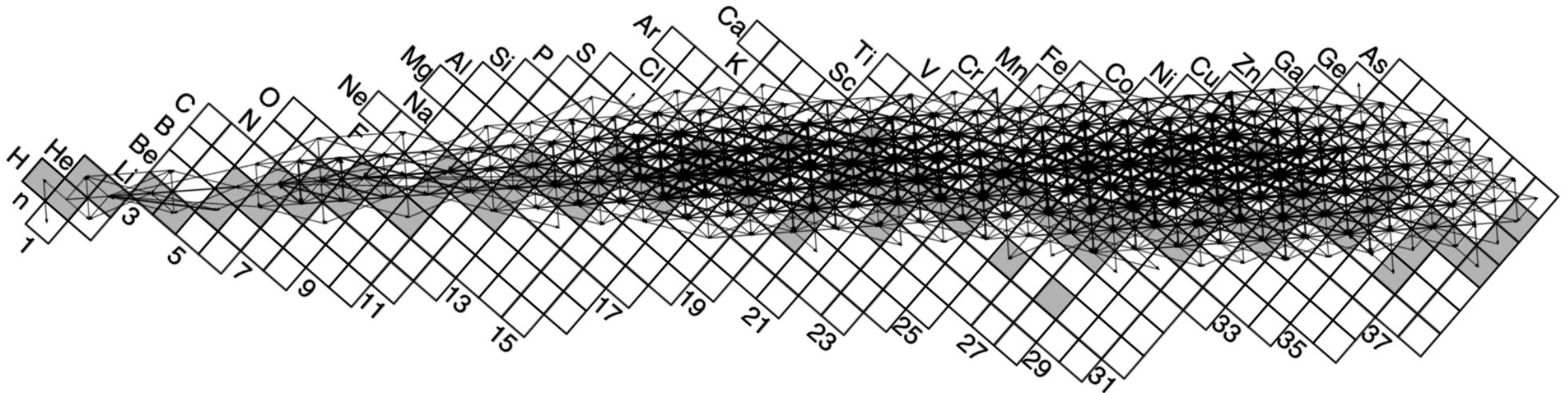
needed:

- measured  $E_r$ ,  $E_x$ ,  $Q$  [masses]
- measured  $\omega\gamma$
- measured cross sections
- measured  $J$ ,  $C^2S$  or ANCs,  $\tau$
- nuclear theory [R matrix]<sup>1,2</sup>
- reaction rate uncertainty<sup>3</sup>
- Hauser-Feshbach model
- level densities/OMP
- strength functions

data bases/codes:

- AME/NUBASE
- ENSDF
- Endt 1990<sup>4</sup>
- original literature
- TALYS

examples: NSE or QSE in Type Ia Supernovae  
 [NSE: Nuclear Statistical Equilibrium; QSE: Quasi-Statistical Equilibrium]



- [reaction rates NOT needed for equilibrium]
- reaction rates needed for freezeout

needed:

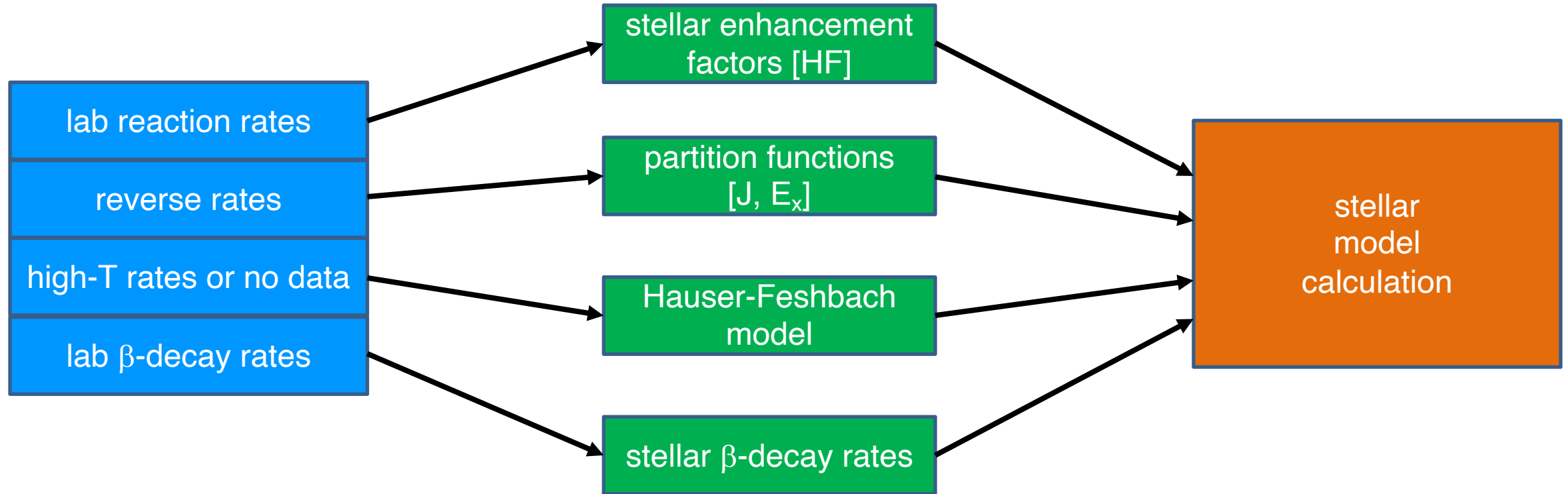
- measured masses [binding energies]
- measured spins, excitation energies
- $\beta$ -decay rates

data bases:

- AME
- NUBASE
- ENSDF
- ...



figure courtesy of Matt Mumpower

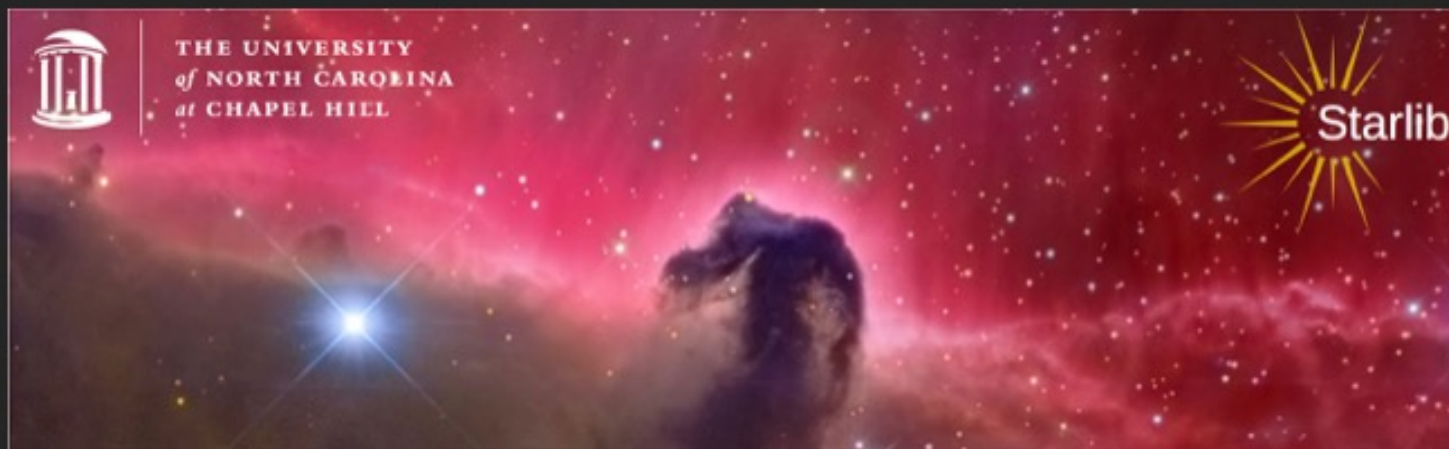


- nuclear data needed: **excitation energies, spins, level densities, branching ratios, Gamow-Teller strengths, ...**
- **what are the uncertainties in the “corrected” quantities [from model predictions]?**

# Rate-Library

Project maintained by **Starlib**

Hosted on GitHub Pages — Theme by **mattgraham**



## STARLIB: Thermonuclear Rate Library

<https://starlib.github.io/Rate-Library/>

Sallaska, Iliadis, Champagne, et al., ApJS 207, 18 (2013)

other rate libraries:

- REACLIB [Thielemann <2000]
- REACLIB [Rauscher/Thielemann]
- JINA REACLIB
- BRUSLIB [2012]
- KADoNiS [2014]
- several others...

STARLIB contains:

- reaction rates
- rate uncertainties
- rate probability densities
- used in MC sensitivity studies





updates of:

REVIEW OF MODERN PHYSICS, VOLUME 83, JANUARY–MARCH 2011

## Solar fusion cross sections. II. The $pp$ chain and CNO cycles

E. G. Adelberger, A. García, R. G. Hamish Robertson, and K. A. Snover

*Department of Physics and Center for Experimental Nuclear Physics and Astrophysics,  
University of Washington, Seattle, Washington 98195, USA*

+ 44 additional co-authors



Available online at [www.sciencedirect.com](http://www.sciencedirect.com)



Nuclear Physics A 841 (2010) 1–30



[www.elsevier.com/locate/nucphysa](http://www.elsevier.com/locate/nucphysa)

## Charged-particle thermonuclear reaction rates: I. Monte Carlo method and statistical distributions

R. Longland<sup>a,b</sup>, C. Iliadis<sup>a,b,\*</sup>, A.E. Champagne<sup>a,b</sup>, J.R. Newton<sup>a,b</sup>,  
C. Ugalde<sup>a,b</sup>, A. Coc<sup>c,d</sup>, R. Fitzgerald<sup>e</sup>

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Berkeley SF III workshop [July 2022; organizers:  
Serenelli, Guglielmetti, Bemmerer, Haxton]

- $\approx$  60 participants [co-authors]
- focus on S factors
- goal to publish by July 2023

- 
- mainly: Richard Longland & CI
  - focus on reaction rates
  - goal to publish by early 2024
  - concurrent major new STARLIB release

- 1. Ensure that evaluated nuclear data are **up-to-date**:**
  - measured ground-state masses, spin-parities, half-lives [AME, NUBASE]
  - cross sections, S factors [EXFOR,...]
  - energies, spin-parities, lifetimes of excited levels [ENSDF]
  - resonance parameters [strengths, partial widths,...]
- 2. Support evaluation of derived nuclear quantities:**
  - laboratory reaction rates [MC method to derive uncertainties]
  - Hauser-Feshbach predictions: level densities, optical potentials, strength functions [**uncertainties?**]
  - stellar weak interaction rates [**uncertainties?**]
- 3. Provide **modern tools** for statistical analysis of nuclear data and derived nuclear quantities:**
  - Bayesian models, machine learning, genetic algorithms,...

**“Ultimately, the future of nuclear physics in this endeavor is a coherent combination of data and theory, which advances our understanding of astrophysics”**

[quote by Matt Mumpower]

**QUESTIONS?**