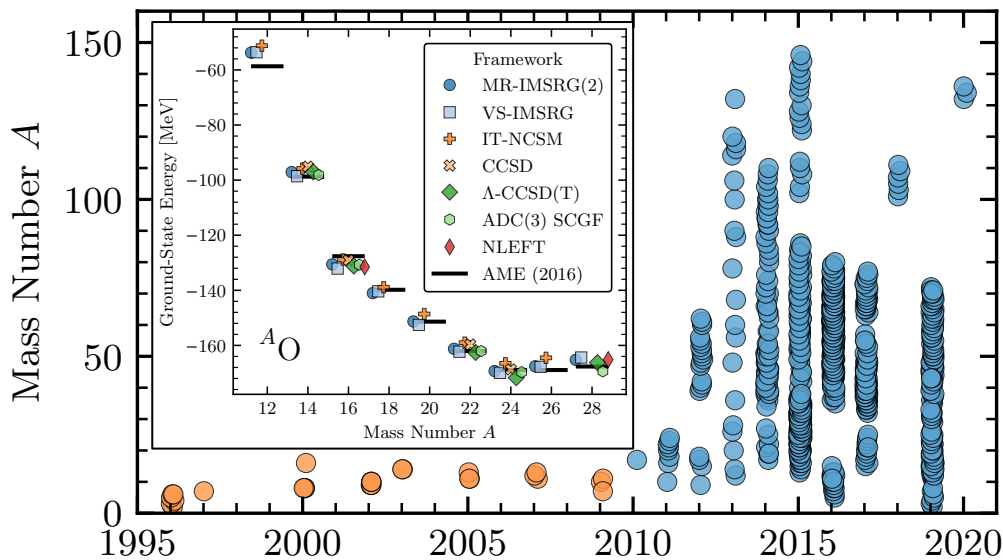


nuclear physics in the *precision era*:
 limitations due to NN+3N forces

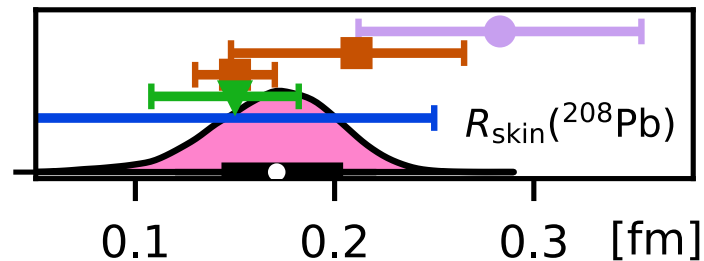


finite nuclei

Hergert, *Front. in Phys.* **8**, 379
 CD & Bogner, *Few Body Syst.* **62**, 109

first *ab initio* calculation of ^{208}Pb ,
 incl. neutron skin prediction with UQ

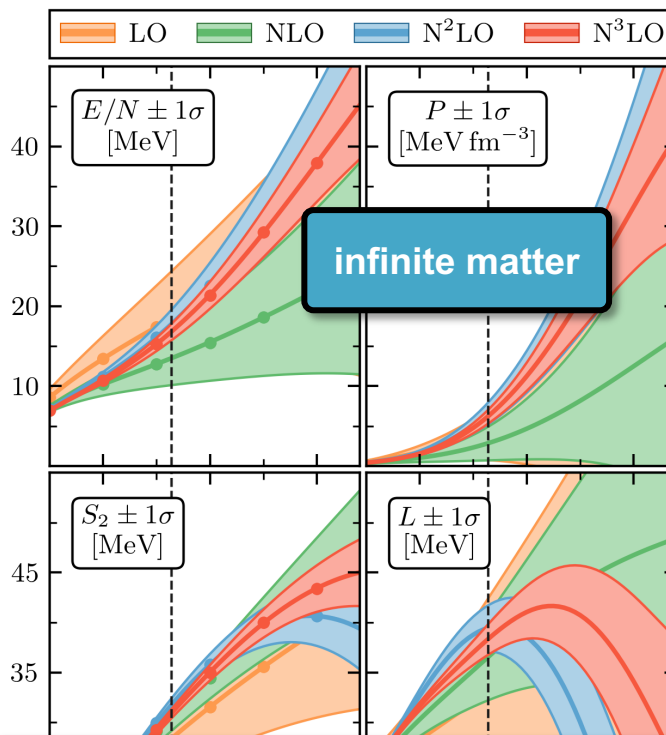
Hu, Jiang *et al.*,
Nature Phys. **18**, 1196



Constraints from:

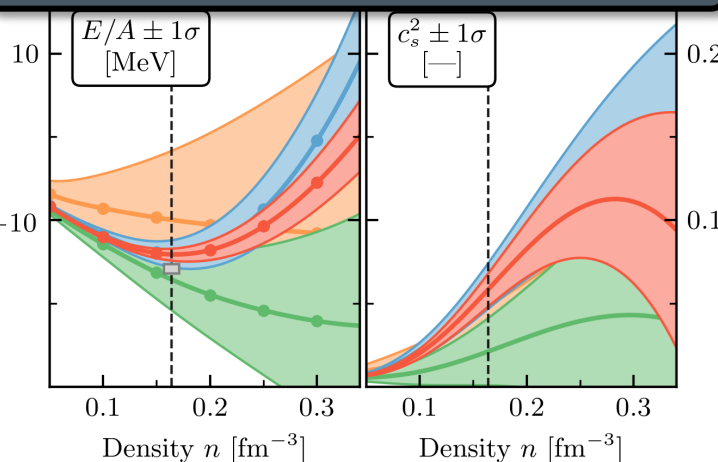
electroweak
 hadronic
 electromagnetic
 gravitational
 waves

CD, Furnstahl, Melendez, Phillips, *PRL* **125**, 202702
 CD, Hebeler, Schwenk, *PRL* **122**, 042501



infinite matter

statistical analysis with *correlated* EFT
 truncation errors & model checking



CD, Holt & Wellenhofer,
Ann. Rev. NPS **71**, 403



Tremendous progress
 in CEFT, many-body theory, UQ & HPC

CEFT enables *ab initio* calculations of
 finite nuclei & nuclear matter at $T \geq 0$
 and arbitrary proton fractions ($n \lesssim 2n_{\text{sat}}$)

Computational & algorithmic advances
 in all many-body frameworks

Here: many-body perturbation theory

- fully automated | GPU-accelerated
- full N³LO calculations with NN+3N contributions at high MBPT orders

Bayesian statistics allows for rigorous
 UQ & propagation in EFT-based
 calculations (use emulators!)

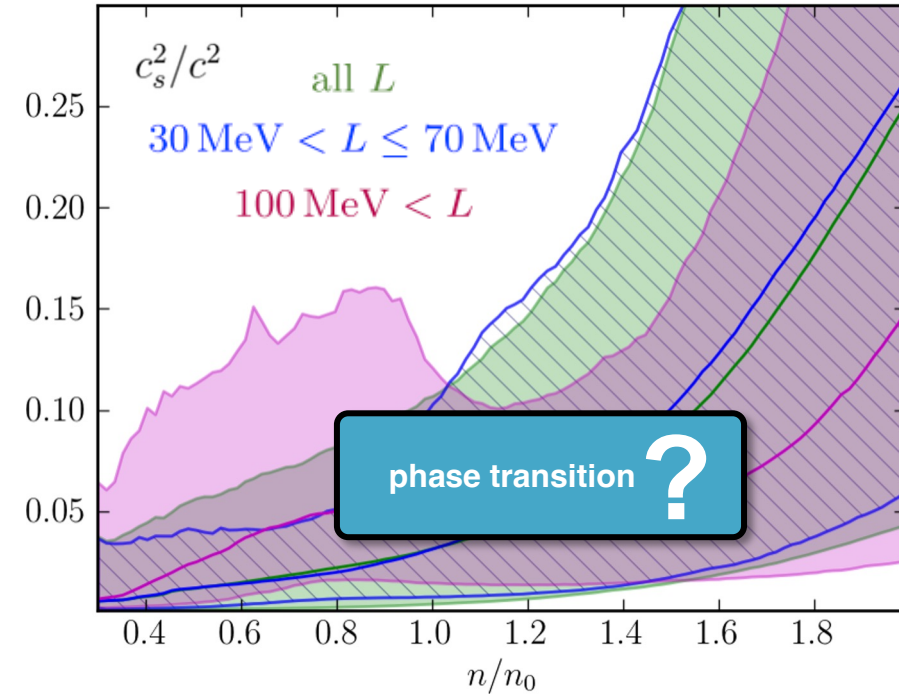
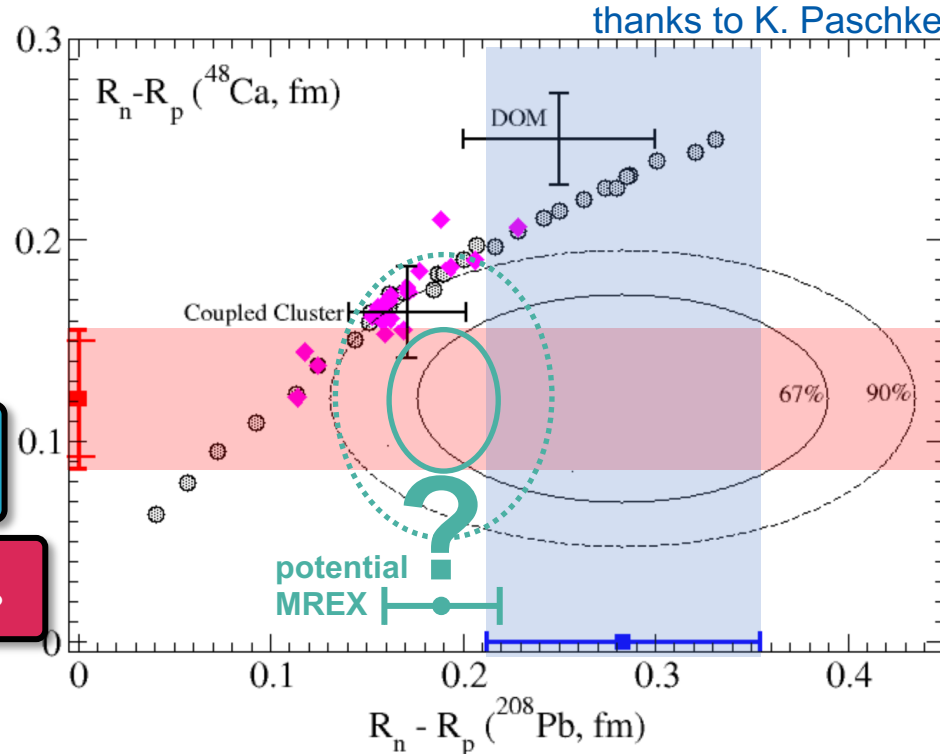
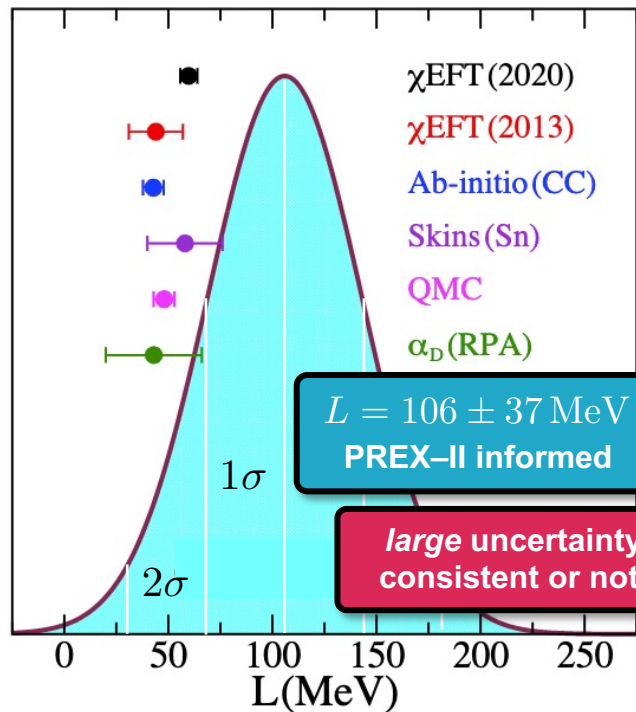
Correlated EFT truncation errors from
 order-by-order calculations and GP
 truncation error model (physics-informed)

At what density does CEFT
 break down and how ?

Need a more precise PREX-II measurement

Tension (at $\approx 1\sigma$) between PREX-II-informed and other determinations of R_{skin} or L and α_D (RCNP)

Adhikari *et al.* (CREX), PRL **129**, 042501
 Reinhard *et al.*, PRL **127**, 232501
 Essick *et al.*, PRL **127**, 192701
 Reed *et al.*, PRL **126**, 172503



MREX at the MESA facility (Mainz) will double the precision of PREX-II (≈ 2030): $R_{\text{skin}}(^{208}\text{Pb})$ to ± 0.03 fm uncertainty

stringent constraint on the neutron-rich matter EOS at n_{sat}

MREX will also provide invaluable data for **calibrating experiments with hadronic probes** at FRIB/FRIB400 and other (future) beam facilities

for MREX, see Thiel, Sfienti *et al.*, J. Phys. G: Nucl. Part. Phys. **46**, 093003

Golden era for neutron-rich matter studies, with great potential for discovery:

- neutron star observation (tidal deformabilities, masses, radii, ...)
- nuclear experiment (skins, polarizabilities, ...)
- nuclear theory (*ab initio* predictions, ...)

Careful: PREX/CREX measured A_{PV} and not the somewhat model-dependent R_{skin} or L

needs joint efforts and funding