Proton mass structure & physics of gravitational form factors

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SoLID Opportunities and Challenges of Nuclear Physics at the Luminosity Frontier

Argonne National Lab

Outline

- Proton mass structure & trace anomaly
- Mass & scalar/anomaly radii
- Longitudinal and transverse spin sum rules
- Momentum flux/current density tensor
 - "Mechanical" properties
 - Gravitational tensor-monopole moment
- Summary

Two recent talks:

Dec, 10, 2021, online, SoLID review June 13, 2022, in person, INT workshop

proton mass structure and trace anomaly

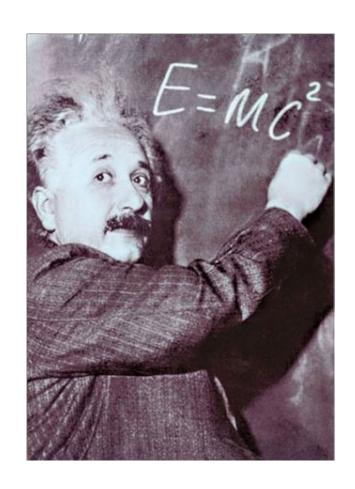
Mass in classical physics

- Newton's law, $\vec{F}=m_I \vec{a}$, m_I (inertia) as a proportionality constant
 - Momentum conservation $\vec{p} = m_I \vec{v}$ is a conserved quantity
 - Energy conservation kinetic energy: $T = \frac{1}{2} m_I \vec{v}^2$
- Gravitational force, $\vec{F}_g \sim m_g$ (gravity charge)
 - $m_I = m_g \rightarrow$ Einstein's general relativity
- Additive mass: "mass conservation law"

Mass in relativity

- Mass is not a fundamental concept
 Space-time translational symmetry
 → energy-momentum and their conservation.
 - Mass as a relativistic invariant $m = \sqrt{E^2 p^2} = E_0/c^2$
- Mass is the energy in the rest frame!
- Mass is NOT additive, but energy is!
 In particular, the proton mass is not a sum of other masses!

 $M_p = 3m_q + ...$ does not make sense!



Energy sources in QCD

QCD Hamiltonian:

$$H = \int T^{00} d^3 \vec{x}$$

Quark mass

$$\sum_q m_q \bar{q}q$$
 but not $3m_q$

Quark "kinetic energy" density

$$\sum \overline{q} (i \vec{\gamma} \cdot \vec{D}) q \sim \sum_q \sqrt{k_q^2 + m_q^2} - m_q$$

Gluon energy density

$$\frac{1}{2}(\overrightarrow{E_a}^2 + \overrightarrow{B_a}^2)$$

UV divergences & anomaly

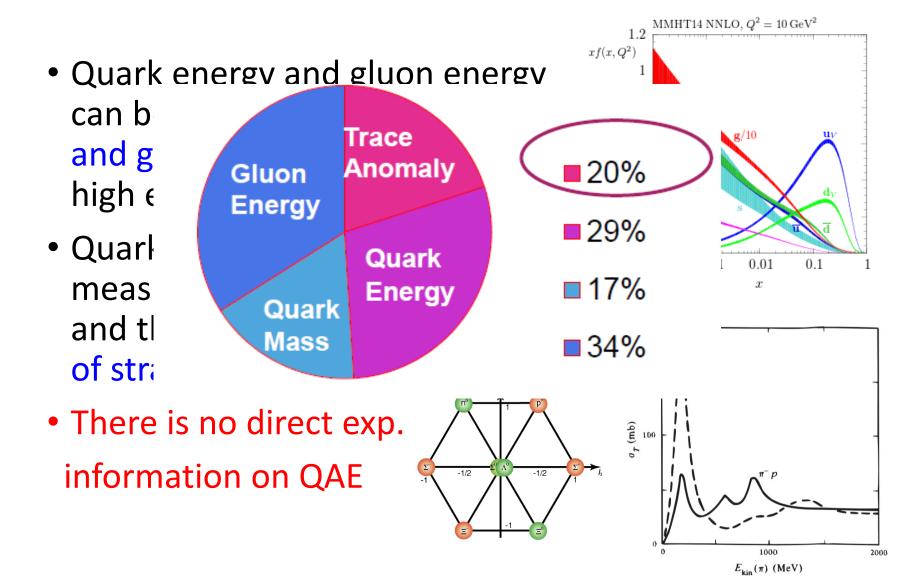
- In a naïve calculation, all contributions are infinite, but the sum is finite.
- Quarks and gluons are not objects that can be taken out the nucleon and studied individually.
- They have to be "defined" (renormalization) by theorists so that all calculations make sense.
 - Definition = "scheme and scale"
- However, there is a piece coming out of renormalization that is independent of any sensible scheme

$$H_a = \int d^3\vec{x} \; \frac{9\alpha_s}{16\pi} (\mathbf{E}^2 - \mathbf{B}^2).$$

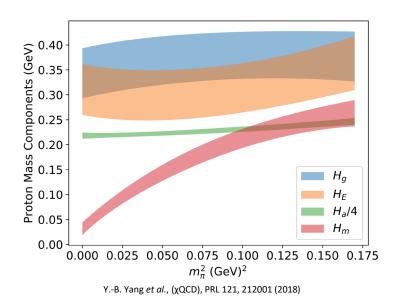
Trace anomaly & mass scale

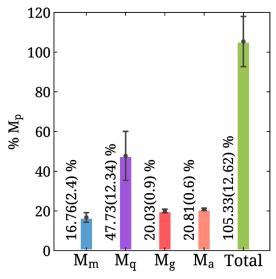
- Anomaly is a scalar.
- Anomaly sets the strong interaction scale.
 - In principle, the most fundamental scale in QCD is Λ_{QCD} , however, this appears only in QCD running.
- The trace anomaly sets QCD scale in a more fundamental way: it is the source of scale symmetry breaking in classical chromodynamics (CCD).
 - It is the bag constant in MIT bag model.
 - it is the gluon condensate in QCD sum rules.
 - It related to the instanton density in QCD vacuum (I. Zahed, 2021)
 - It is a yardstick for all mass scales in QCD, It is like the Higgs field in EW theory. (Ji, Liu, & Schaefer, 2021)

Proton mass structure from data (Ji,1995)



Proton mass on the lattice





C. Alexandrou *et al.*, (ETMC), PRL 119, 142002 (2017) C. Alexandrou *et al.*, (ETMC), PRL 116, 252001 (2016)

Trace anomaly only constrained through sum-rules not calculated directly.

Anomaly matrix element

- Direct calculation on lattice challenging (K. Liu, Y. B. Yang et al) dealing with gluon scalar operator
- Direct measurement in high-energy exp. twist-4 matrix element enhanced by $\frac{1}{\alpha_s} \sim lnQ^2$ (Y. Hatta et al.)

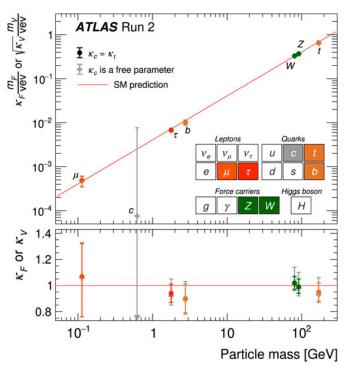
Test of the QCD "Higgs mechanism"

 The couplings of the scalars with the hadrons are proportional to the hadron masses.

$$g_{HHS} \sim m_H$$

this also works for pion and kaon.

One can do the similar test as one does for Higgs particles at LHC but much more complicated



Mass and scalar/anomaly radii

Mass structure: gravitations form factors

Form factors of EMT for quarks and gluons (Ji,1996)

$$\langle P'|T_{q,g}^{\mu\nu}|P\rangle = \overline{U}(P') \left[A_{q,g}(\Delta^2) \gamma^{(\mu} \overline{P}^{\nu)} + B_{q,g}(\Delta^2) \overline{P}^{(\mu} i \sigma^{\nu)\alpha} \Delta_{\alpha} / 2M + C_{q,g}(\Delta^2) (\Delta^{\mu} \Delta^{\nu} - g^{\mu\nu} \Delta^2) / M + \overline{C}_{q,g}(\Delta^2) g^{\mu\nu} M \right] U(P) ,$$

• Form factors for the total EMT (Pagels, 1966)

$$\begin{split} \langle P' \left| T^{\mu\nu} \right| P \rangle &= \bar{u} \left(P' \right) \left[A \left(Q^2 \right) \gamma^{(\mu} \bar{P}^{\nu)} \right. \\ &+ B \left(Q^2 \right) \bar{P}^{(\mu} i \sigma^{\nu)\alpha} q_\alpha / 2 M \\ &+ C \left(Q^2 \right) \left(q^\mu q^\nu - g^{\mu\nu} q^2 \right) / M \right] u(P) \; , \\ A &= A_q \, + A_g \text{, } B \; \& \; C \; etc., \quad \bar{C}_q \, + \; \bar{C}_g \, = \, 0 \end{split}$$

Scalar/anomaly form factor and confinement

Form factor of the scalar density (Breit frame)

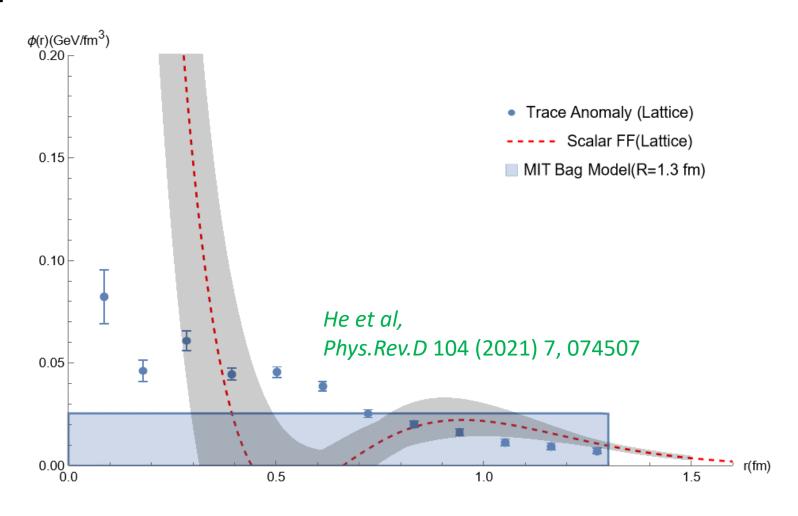
$$\left\langle P' \left| T_{\mu}^{\mu} \right| P \right\rangle = \bar{u} \left(P' \right) u(P) G_s(Q^2) ,$$
 where,

$$G_s(Q^2) = \left[MA(Q^2) - B(Q^2) \frac{Q^2}{4M} + C(Q^2) \frac{3Q^2}{M} \right]$$

result from conservation law

- Can also be calculated directly from the scalar
- Fourier transformation of Gs gives us the scalar field distribution in $M \rightarrow \text{large}$
- Dynamical MIT "bag constant".

Scalar field distribution inside the proton



Scalar/scale radius

The radius

$$\langle r^2 \rangle_s = -6 \frac{dA(Q^2)}{dQ^2} - 18 \frac{C(0)}{M^2}$$

MIT bag scalar radius

$$r_S^2 = \frac{3}{5}R^2$$
, $r_S = 1.3fm$

Mass form factor

How does energy distribute inside the nucleon?

Again choose the Breit frame (assuming M large)

$$\langle P' | T^{00} | P \rangle = \bar{u}(P') u(P) G_m(Q^2)$$
.

where

$$G_m(Q^2) = \left[MA(Q^2) - B(Q^2) \frac{Q^2}{4M} + C(Q^2) \frac{Q^2}{M} \right]$$

Scalar and mass radii

• Definition:

$$\langle r^2 \rangle_{s,m} = -6 \frac{dG_{s,m}(Q^2)}{dQ^2} ,$$

$$\langle r^2 \rangle_s = -6 \frac{dA(Q^2)}{dQ^2} - 18 \frac{C(0)}{M^2}$$

 $\langle r^2 \rangle_m = -6 \frac{dA(Q^2)}{dQ^2} - 6 \frac{C(0)}{M^2}$,

The difference

$$\langle r^2 \rangle_s - \langle r^2 \rangle_m = -12 \frac{C(0)}{M^2}$$

• Conjecture $\langle r^2 \rangle_{\scriptscriptstyle S} > \langle r^2 \rangle_{\scriptscriptstyle m}$ or C(0)<0

Lattice calculations

Radius from A-FF:

Hagler et al (2008)

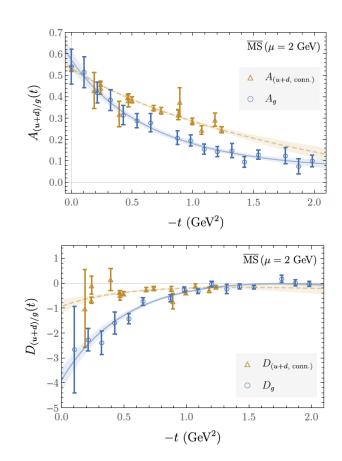
Shanahan et al (2018)

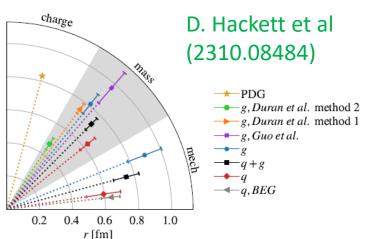
$$\langle r^2 \rangle_A = (0.5 fm)^2$$

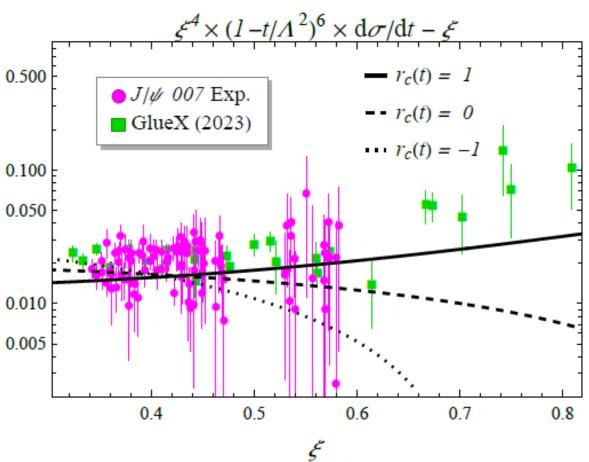
C-FF contribution

take
$$D = -5.0$$

$$\langle r^2 \rangle_S = (1.1 fm)^2$$
$$\langle r^2 \rangle_m = (0.75 fm)^2$$







• Guo, Ji, Yuan, 2023

Mass radius from separate quark and gluon contributions

- Method 1: without C-bar
 - Separate contributions to A and C from quarks and gluons (X. Ji, ...)
- Method 2: with C-bar, result depends on how to split the trace part of the energy-momentum tensor into q/g contributions
 - In the chiral limit, the scalar part is entirely from the gluon (K. Liu,...)
 - Partition from anomaly (Y. Hatta et al, JHEP 12 (2018) 008)

Longitudinal and transverse spin structure

Spin sum rule

A spin sum rule was derived in 1996

$$\frac{1}{2} = J_q + J_g$$

 $J_q \& J_g$ are related to the EMT form factor

$$J_{q,g} = \frac{1}{2} [A_{q,g}(0) + B_{q,g}(0)]$$

Lattice calculations by Cyprus, Kentucky, MIT groups

- What does it mean physically?
 - It works for the nucleon spin in the rest frame.
 - This is a twist-2 transverse spin sum rule for a proton moving at the speed of light

Ji & Yuan Phys.Lett.B 810 (2020) 135786

Twist-3 spin sum rules

- Longitudinal polarization
 - Jaffe and Manohar (1990)

$$\frac{1}{2} = \frac{1}{2} \Delta\Sigma + \Delta G + \ell_q^z + \ell_g^z$$

OAM are twist-3 quantities requires twist-3 GPDs (Hatta, Ji, Yuan...)

- Transverse polarization
 - Rotated version of the Jaffe Manohar Longitudinal spin sum rule

X. Ji, Y. Guo & K. Shiells, Nuc. Phys. B 969, 115440 (2021) $\frac{1}{2}\Delta q_T + \Delta G_T + l_q^{x(3)} + l_g^{x(3)} = \frac{\hbar}{2} \qquad \text{Twist-3} \\ \text{Our 2021 result} \\ \Delta q_T \; , \; \Delta G_T \qquad \text{Involve measurable PDFs in DIS and correspond to spin} \\ l_q^{x(3)} \; , \; l_g^{x(3)} \qquad \text{Involve twist-3 GPDs and correspond to canonical OAM}$

Momentum flux/current density tensor (MFDT), and form factor C(D)

Space-components of EMT

• Consider the space-only components T^{ij} of the gravitational form factors

$$\langle P'|T_{q,g}^{\mu\nu}|P\rangle = \overline{U}(P')\left[A_{q,g}(\Delta^2)\gamma^{(\mu}\overline{P}^{\nu)} + B_{q,g}(\Delta^2)\overline{P}^{(\mu}i\sigma^{\nu)\alpha}\Delta_{\alpha}/2M + C_{q,g}(\Delta^2)(\Delta^{\mu}\Delta^{\nu} - g^{\mu\nu}\Delta^2)/M + \overline{C}_{q,g}(\Delta^2)g^{\mu\nu}M\right]U(P),$$

- In the Breit frame, only C and C-bar are relevant, and in the total T^{ij} , only C
- T^{ij} entering momentum density-conservation as a momentum flux/current density tensor (MFDT)

Momentum conservation

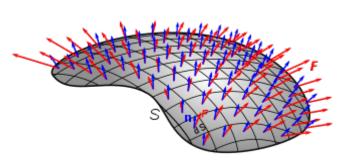
• The momentum density $p^i(\vec{r}) = T^{i0}(\vec{r})$ satisfies the conservation law

$$\frac{\partial p^{i}(\vec{r},t)}{\partial t} + \partial_{j}T^{ij} = 0$$

The flux of i-momentum following through a surface dS is the force per unit area.

$$F^i = \int T^{ij} dS_j$$

which can be + or -.



Nucleon mechanical properties?

• T^{ij} has been compared with a stress tensor of fluids and solids, and its form factor C has been attributed the meaning in fluids and solids

(M. Polyakov et al., Int.J.Mod.Phys.A 33 (2018) 26, 1830025)

Pressure interpretation

$$T^{ij}(\mathbf{r}) = \left(\frac{r^i r^j}{r^2} - \frac{1}{3} \delta^{ij}\right) s(r) + \delta^{ij} p(r).$$

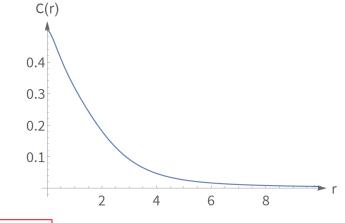
Stability

Conjecture: C(D)<0 to be stable

Momentum flux density in H atom

 We find that C(D) is positive for hydrogen atom

$$T_{\mathrm{QM}}^{ij}(\vec{r}) = (\delta^{ij}\nabla^2 - \nabla^i\nabla^j)\frac{C_{\mathrm{QM}}(r)}{M}$$



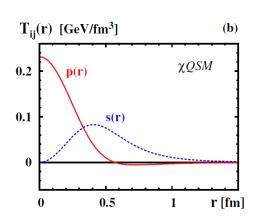
$$\frac{C_{\text{QM}}(r)}{M} = \frac{1}{2\nabla^2} T_{\text{QM}}^{ii} = \frac{e^{-2\alpha r} \alpha (2\alpha r + 1)}{16\pi r^2} - \frac{\alpha}{16\pi r^2}$$

$$D = 1$$

• It is not certain that sign of D tells us the stability of a system. All bound states are stable due to Quantum Mechanics.

Is T^{ij} related to pressure?

In fluid mechanics
 pressure is always positive (stability),
 resulting from random thermal motion.



- However, p(r) defined from T^{ij} is not.
- What is pressure?

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random motion vs. collective motion (laser beam)
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(pressure)

positive

(impact force/unit area)

depends on the impact dir.

motion be either way

Gravitational field from MFDT

Linearized Einstein equation

$$\Box \bar{h}^{\mu\nu} = \frac{16\pi G}{c^4} T^{\mu\nu}$$
 where $g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}$ and $\bar{h}^{\mu\nu} = h^{\mu\nu} - \frac{\eta^{\mu\nu}}{2} h^{\rho}_{\rho}$

The solution with C form factor is

$$h_C^{00}(\vec{r}) = -\frac{8\pi G}{c^4 M} C(r)$$

$$h_C^{ij}(\vec{r}) = \frac{8\pi G}{c^4 M} C(r) \delta^{ij}$$

• Given C(r) decays exponentially, so does the metric perturbation.

Multipoles from momentum current

Scalar multipoles ("pressure" multipoles)

$$S(r)=T^{ii}(r)$$

$$S^{J}=\int d^{3}\vec{r}\;S(r)r_{r_{1}}\;...\;r_{i_{j}}$$
 (S⁽⁰⁾=0, however, scalar monopole density does not)

Tensor multipoles (natural parity, "shear pressure")

$$T^{J}$$
 from $\int d^{3}\vec{r} T_{ij}(\mathbf{r})\mathbf{r}_{i} \dots \mathbf{r}_{ij}$

Tensor multipoles (unnatural parity)

$$\tilde{T}^{J}$$
 from $\int d^{3}\vec{r} T_{i[j}(\mathbf{r})\mathbf{r}_{r_{1}]} \dots \mathbf{r}_{i_{1}}$

Tensor monopole moment

Tensor monopole T0

$$T^{(0)} = \frac{1}{5} \int d^3 \vec{r} T_{ij}(\vec{r}) \left(r_i r_j - \frac{\delta_{ij}}{3} r^2 \right)$$

normalization,
$$\tau = -T^{(0)}/2 = D/4M$$

• For H-atom, we find (Ji & Liu, 2022)

$$\tau = \hbar^2/4M (1+O(\alpha))$$

$$\tau_H/\tau_0 - 1 = \frac{4\alpha}{3\pi} \left(\ln \alpha^2 - 0.028 \right)$$

Summary

- Anomaly contribution to the nucleon mass is perhaps the most interesting,
 - Higgs mechanism
 - QCD confinement
 - Radius
- Mass/energy radius, a fundamental property to be determined by exp and theory
- Spin sum rule is the simplest at twsit-2 level.
- Momentum current density: gravitational tensor monople moment

Mass as internal energy

Internal mass as a store of energy

$$Mc^{2} = \langle N | \widehat{H}_{QCD} | N \rangle |_{\vec{P}=0}$$

This is how the lattice QCD calculate.

 For any relativistic system, the Hamiltonian can be separated into two terms (Ji, PRL,1995),

$$\widehat{H}_{QCD} = \widehat{H}_T + \widehat{H}_S$$

This separation is a fundamental property of special relativity and both parts are scale invariant

Lorentz symmetry

- Energy is related to $H = \int d^3\vec{r} \ T^{00}(\vec{r})$
- $T^{\mu\nu}$ has a mixed symmetry under Lorentz transformation (1,1) + (0,0), and the separate parts are scheme and scale independent.

$$T^{\mu\nu} = T_S^{\mu\nu} + T_T^{\mu\nu}$$
,
 $H = H_S + H_T$,
 $M = M_S + M_T$

Tensor and scalar energies

Tensor energy

$$E_T = \langle H_T \rangle$$

is related to the usual kinetic and potential energy sources.

Scalar energy

$$E_S = \langle H_S \rangle$$

is related to related to scale-breaking properties of the theory $(\partial^{\mu} j_{Du} \sim H_s)$, such as

- Quark mass m_q
- Trace anomaly (quantum breaking of scale symmetry).

Does the trace anomaly contribute to the mass?

• Of course! (can also be derived by through time-translation)

•
$$\mathbf{T}^{\mu}_{\mu} = (1 + \gamma_m)m\bar{\psi}\psi + \frac{\beta(g)}{2g}F^2$$

- In the massless QCD limit, all contribution to the scalar energy (mass) comes from the anomaly. (without trace anomaly, there is no mass!)
- The scalar field from trace anomaly plays the key role for mass generation, like a Higgs mechanism.

Relativistic "virial theorem"

As an important feature of relativity, one can show

$$E_T = 3E_S$$
 (virial theorem)

3 is the dimension of space.

- Scalar energy sets the scale of the tensor energy (kinetic and potential energies of the system).
- In non-relativistic limit of QED & gravity, it reduces

$$\langle V \rangle = -2 \langle T \rangle$$

kinetic energy sets the scale for potential energy!

Renormalization: tensor part

Standard!

$$T^{(\mu\nu)} = T_q^{(\mu\nu)} + T_g^{(\mu\nu)}$$

renormalization of $T_{q,g}^{(\mu\nu)}$ is know to four-loops

$$\begin{pmatrix} T_q^{\mu\nu} \\ T_g^{\mu\nu} \\ T_{gv}^{\mu\nu} \end{pmatrix} = \begin{pmatrix} Z_{qq} & Z_{qg} & Z_{qa} & Z_{qe} \\ Z_{gq} & Z_{gg} & Z_{ga} & Z_{ge} \\ 0 & 0 & Z_{aa} & Z_{ae} \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} T_q^{\mu\nu} \\ T_g^{\mu\nu} \\ T_{gv}^{\mu\nu} \\ E^{\mu\nu} \end{pmatrix}^R$$

$$\frac{d}{d \ln \mu_f^2} \begin{pmatrix} q_s \\ g \end{pmatrix} = \begin{pmatrix} P_{qq} & P_{qg} \\ P_{gq} & P_{gg} \end{pmatrix} \otimes \begin{pmatrix} q_s \\ g \end{pmatrix}$$



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Low moments of the four-loop splitting functions in QCD



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ABSTRACT

We have computed the four lowest even-N moments of all four splitting functions for the evolution of flavour-singlet parton densities of hadrons at the fourth order in the strong coupling constant α_s . The perturbative expansion of these moments, and hence of the splitting functions for momentum fractions $x \gtrsim 0.1$, is found to be well behaved with relative α_s -coefficients of order one and sub-percent effects on the scale derivatives of the quark and gluon distributions at $\alpha_s \lesssim 0.2$. More intricate computations, including other approaches such as the operator-product expansion, are required to cover the full x-range relevant to LHC analyses. Our results are presented analytically for a general gauge group for detailed checks and validations of such future calculations.

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$$\gamma_{ik}(N,\alpha_s) = \sum_{n=0}^{\infty} a_s^{n+1} \gamma_{ik}^{(n)}(x)$$
 with $a_s \equiv \frac{\alpha_s(\mu_f^2)}{4\pi}$.

$$\begin{split} \gamma_{\rm ps}^{(3)}(N=2) &= n_f\,C_F^3\left(\frac{227938}{2187} + \frac{1952}{81}\,\zeta_3 + \frac{256}{9}\,\zeta_4 - \frac{640}{3}\,\zeta_5\right) \\ &+ n_f\,C_AC_F^2\left(-\frac{162658}{6561} + \frac{8048}{27}\,\zeta_3 - \frac{1664}{9}\,\zeta_4 + \frac{320}{9}\,\zeta_5\right) \\ &+ n_f\,C_A^2\,C_F\left(-\frac{410299}{6561} - \frac{26896}{81}\,\zeta_3 + \frac{1408}{9}\,\zeta_4 + \frac{4480}{27}\,\zeta_5\right) \\ &+ n_f\,\frac{d_R^{abcd}}{n_c}\frac{d_R^{abcd}}{0}\left(\frac{1024}{9} + \frac{256}{9}\,\zeta_3 - \frac{2560}{9}\,\zeta_5\right) - n_f^2\,C_F^2\left(\frac{73772}{6561} + \frac{5248}{81}\,\zeta_3 - \frac{320}{9}\,\zeta_4\right) \\ &+ n_f^2\,C_A\,C_F\left(\frac{160648}{6561} + 48\,\zeta_3 - \frac{320}{9}\,\zeta_4\right) + n_f^3\,C_F\left(-\frac{1712}{729} + \frac{128}{27}\,\zeta_3\right)\,, \end{split}$$

$$\begin{split} & \gamma_{\rm qg}^{(3)}(N=2) &= n_f \, C_F^3 \left(\frac{16489}{729} + \frac{736}{81} \, \xi_3 + \frac{256}{9} \, \xi_4 - \frac{320}{3} \, \xi_5 \right) \\ &+ n_f \, C_A^3 \left(-\frac{88769}{729} + \frac{31112}{81} \, \xi_3 - 132 \, \xi_4 - \frac{3560}{27} \, \xi_5 \right) - n_f \, C_A C_F^2 \left(\frac{1153727}{13122} - \frac{7108}{81} \, \xi_3 + \frac{1136}{9} \, \xi_4 - \frac{2000}{9} \, \xi_5 \right) + n_f \, C_A^2 \, C_F \left(\frac{763868}{6561} - \frac{12808}{27} \, \xi_3 + \frac{2068}{9} \, \xi_4 + \frac{40}{9} \, \xi_5 \right) \\ &+ n_f \, \frac{d_F^{abcd} \, d_A^{abcd}}{n_a} \left(\frac{368}{9} - \frac{992}{9} \, \xi_3 - \frac{2560}{9} \, \xi_5 \right) - n_f^2 \, C_F^2 \left(\frac{110714}{6561} + \frac{272}{9} \, \xi_3 - \frac{224}{9} \, \xi_4 \right) \\ &+ n_f^2 \, C_A \, C_F \left(\frac{249310}{6561} + \frac{5632}{81} \, \xi_3 - \frac{440}{9} \, \xi_4 \right) + n_f^2 \, C_A^2 \left(\frac{48625}{2187} - \frac{3572}{81} \, \xi_3 \right) \\ &+ 24 \, \xi_4 + \frac{160}{27} \, \xi_5 \right) + n_f^2 \, \frac{d_F^{abcd} \, d_F^{abcd}}{n_a} \left(-\frac{928}{9} - \frac{640}{9} \, \xi_3 + \frac{2560}{9} \, \xi_5 \right) \\ &+ n_f^3 \, C_F \left(-\frac{8744}{2187} + \frac{128}{27} \, \xi_3 \right) + n_f^3 \, C_A \left(\frac{3385}{2187} - \frac{176}{81} \, \xi_3 \right) \, , \end{split}$$

$$\gamma_{\rm gq}^{(3)}(N=2) = -\gamma_{\rm qq}^{(3)}(N=2)$$
,

$$\gamma_{\rm gg}^{(3)}(N\!=\!2) = -\gamma_{\rm qg}^{(3)}(N\!=\!2) ,$$

Renormalization: scalar part

Scale invariant!

$$(1+\gamma_m)m\bar{\psi}\psi + \frac{\beta(g)}{2g}F^2$$

 There is no mixing between scalar and tensor in any renormalization method (lattice, DR)!

Unless Lorentz symmetry is broken.

Other schemes breaking Lorentz symmetry is not interesting.

Mass separation

Independent of dynamics

$$M_S = \frac{1}{4}M$$
, $M_T = \frac{3}{4}M$
 $M_T = 3 M_S$,

• $H_S = \frac{1}{4} T_{\mu}^{\mu}$ is the source for dilatation symmetry breaking. Dilaton field.

QCD energies in the nucleon

Four different types

$$H_{\mathrm{QCD}} = H_q + H_m + H_g + H_a$$
.
$$H_q = \int d^3 \vec{x} \; \bar{\psi} (-i \mathbf{D} \cdot \alpha) \psi, \qquad \text{Quark energy}$$

$$H_m = \int d^3 \vec{x} \; \bar{\psi} m \psi, \qquad \qquad \text{Quark mass}$$

$$H_g = \int d^3 \vec{x} \; \frac{1}{2} (\mathbf{E}^2 + \mathbf{B}^2), \qquad \qquad \text{Gluon energy}$$

$$H_a = \int d^3 \vec{x} \; \frac{9\alpha_s}{16\pi} (\mathbf{E}^2 - \mathbf{B}^2). \qquad \qquad \text{Quantum Anomalous}$$
 Energy (QAE)

Dynamical origin: scalar field & Higgs mechanism

Dynamical origin

- In the nucleon models, Λ_{QCD} is replaced by some scale related to dynamics.
 - Chiral symmetry breaking
 Scale set by chiral condensate or constituent quark masses
 - Color confinement
 MIT bag constant: energy density of false vacuum
 - Instanton liquid
 Scale related to instanton size and density
 - AdS/CFT...
- In lattice QCD, the scale is related to the lattice spacing, a.

Scalar energy, quantum anomalous energy (QAE)

Trace relation for mass

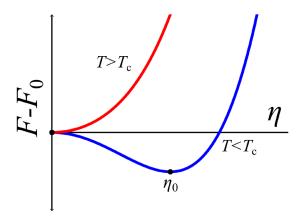
$$2M^{2} = \left\langle P \left| (1 + \gamma_{m}) m \bar{\psi} \psi + \frac{\beta(g)}{2g} F^{2} \right| P \right\rangle$$

- Not a "total" mass sum rule.
- It is a sum rule for scalar part of the mass

$$E_{S} = \frac{M}{4} = \alpha \langle p | m \bar{\psi} \psi | p \rangle + \beta \langle p | F^{2} | p \rangle$$

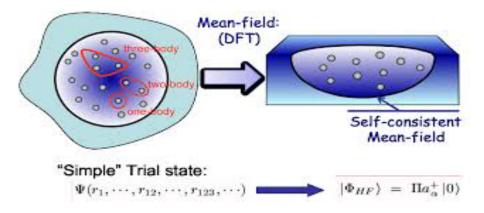
Energy of a scalar field

• Scalar field is special because it can have a vacuum condensate: non-vanishing expectation value in the physical vacuum. $\langle \eta \rangle \neq 0$, which stores the internal energy (latent heat)



Mean field theory for nuclear structure

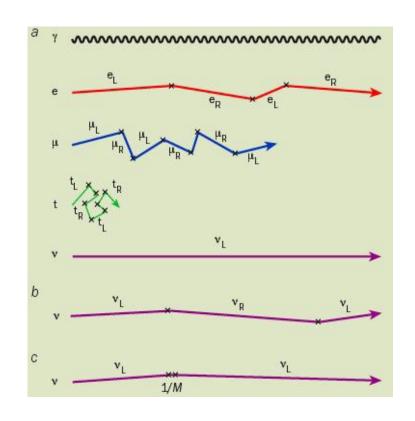
 Traditional theory for nuclear structure: mean field theory



- Two important features:
 - a) Mean field depth is about ~40 MeV
 - b) Large spin-orbit splitting for nuclear shells.

Masses of electrons and leptons: Higgs mechanism

• There is a scalar field H, which interacts with the fermions $g\bar{\psi}\psi H$. H acquires an expectation value in the vacuum after SSB, $\langle H \rangle = v = 246$ GeV, hence the fermion mass,

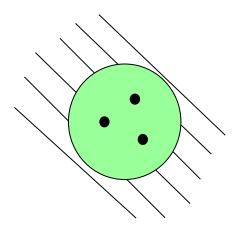


Quantum anomalous energy (QAE) contribution to the proton mass:

- The scalar field has a VEV: $\langle 0|F^2|0\rangle$
- QAE comes from the scalar response to the presence of the quarks.

$$\phi = F^2 - \langle 0|F^2|0\rangle$$

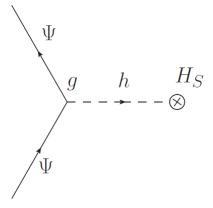
 The contribution is similar to the Higgs mechanism in electroweak theory, with gluon scalar as a dynamical Higgs field.



Dynamical picture of the fermion mass in Higgs mechanism

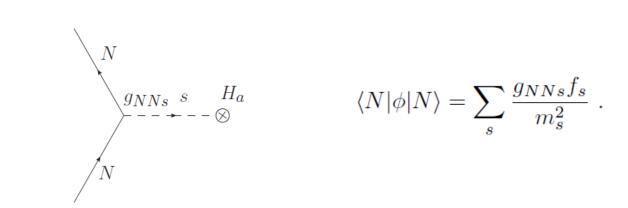
 Part of the fermion mass comes from the dynamical excitation of the higgs field in the presence of fermion

$$m_f \sim \langle f|H_S|f\rangle \sim \langle f|h|f\rangle$$



QAE as a dynamical response

- $E_a \sim \langle N|F^2|N\rangle$
- This matrix element can also be calculated through dynamical scalar excitations



Test of the QCD "Higgs mechanism"

 The couplings of the scalars with the hadrons are proportional to the hadron masses.

$$g_{HHS} \sim m_H$$

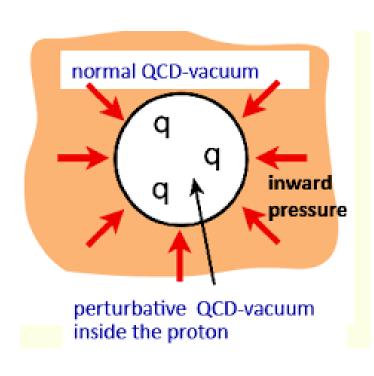
this also works for pion and kaon.

One can do the similar test as one does for Higgs particles at LHC but much more complicated

Scalar spectrum



Physics of QAE in the MIT bag model



M.I.T. Bag Model

 The boundary condition generates discrete energy eigenvalues.

$$\varepsilon_n = \frac{x_n}{R}$$

R - radius of the Bag

$$x_1 = 2.04$$

$$E_{kin}(R) = N_q \frac{x_n}{R}$$

N_q = # of quarks inside the bag

$$E_{pot}(R) = \frac{4}{3}\pi R^3 B$$

B – bag constant that reflects the bag pressure

26

Mass = quark kinetic energy + B(scalar-field condensate)

Gravitational form factors

Scalar form factor

Form factor of the scalar density

$$\left\langle P' \left| T^{\mu}_{\mu} \right| P \right\rangle = \bar{u} \left(P' \right) u(P) G_s(Q^2) ,$$
 where,
$$G_s(Q^2) = \left[MA(Q^2) - B(Q^2) \frac{Q^2}{4M} + C(Q^2) \frac{3Q^2}{M} \right]$$

- Fourier transformation of Gs gives us the scalar field distribution inside the Nucleon
- Dynamical MIT "bag constant".

Scalar radius

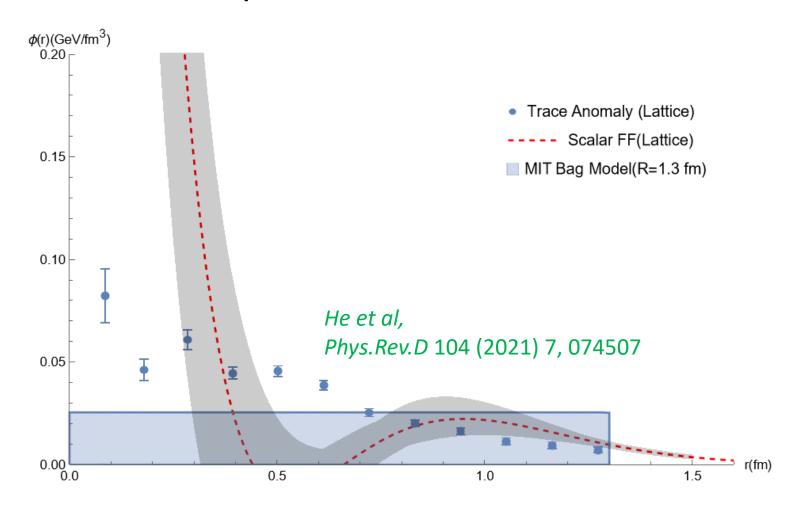
- Scalar field radius might be similar to confinement radius
- The radius

$$\langle r^2 \rangle_s = -6 \frac{dA(Q^2)}{dQ^2} - 18 \frac{C(0)}{M^2}$$

MIT bag scalar radius

$$r_s^2 = \frac{3}{5}R^2$$
, $r_s = 1.3fm$

Scalar field (QAE) distribution inside the proton



Mass form factor

$$\langle P' | T^{00} | P \rangle = \bar{u} (P') u(P) G_m(Q^2)$$
.

where

$$G_m(Q^2) = \left[MA(Q^2) - B(Q^2) \frac{Q^2}{4M} + C(Q^2) \frac{Q^2}{M} \right]$$

Scalar and mass radii

• Definition:

$$\langle r^2 \rangle_{s,m} = -6 \frac{dG_{s,m}(Q^2)}{dQ^2} ,$$

$$\langle r^2 \rangle_s = -6 \frac{dA(Q^2)}{dQ^2} - 18 \frac{C(0)}{M^2}$$

 $\langle r^2 \rangle_m = -6 \frac{dA(Q^2)}{dQ^2} - 6 \frac{C(0)}{M^2}$,

The difference

$$\langle r^2 \rangle_s - \langle r^2 \rangle_m = -12 \frac{C(0)}{M^2}$$

• Conjecture $\langle r^2 \rangle_S > \langle r^2 \rangle_m$ or C(0)<0

Lattice calculations

Radius from A-FF:

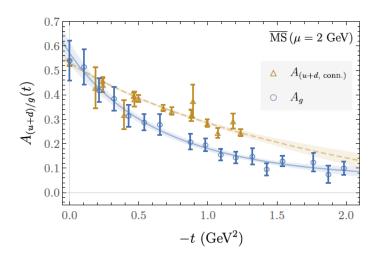
Hagler et al (2008)
Shanahan et al (2018)
$$\langle r^2 \rangle_A = (0.5 fm)^2$$

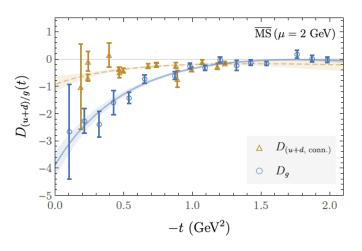
C-FF contribution

$$D = -5.0$$

$$\langle r^2 \rangle_s = (1.1 fm)^2$$

$$\langle r^2 \rangle_m = (0.75 fm)^2$$





Physics of gravitational form factor C(D)

Momentum current & pressure

Gravitational form factors

$$\langle P'|T_{q,g}^{\mu\nu}|P\rangle = \overline{U}(P') \left[A_{q,g}(\Delta^2) \gamma^{(\mu} \overline{P}^{\nu)} + B_{q,g}(\Delta^2) \overline{P}^{(\mu} i \sigma^{\nu)\alpha} \Delta_{\alpha} / 2M + C_{q,g}(\Delta^2) (\Delta^{\mu} \Delta^{\nu} - g^{\mu\nu} \Delta^2) / M + \overline{C}_{q,g}(\Delta^2) g^{\mu\nu} M \right] U(P) ,$$

- In the Breit frame, C and C-bar are related to the form factor of T^{ij}
- T^{ij} has been originally introduced as a stress tensor of fluids and solids, and is related pressure etc
- However, in relativistic theory, its definition is momentum (density) current

Momentum density current

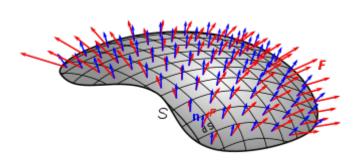
- T^{ij} is actually momentum density current (MC). It describes vector flow (j) of momentum component (i) or vice versa.
- The momentum density $p^i(\vec{r}) = T^{i0}(\vec{r})$ satisfies the conservation law

$$\frac{\partial p^{i}(\vec{r},t)}{\partial t} + \partial_{j} T^{ij} = 0$$

The flux of i-momentum following through a surface dS is just

$$F^i = \int T^{ij} dS_j$$

which can be + or -.



Questions about pressure interpretations

- 1. Is the pressure the same as we understand in a gas or liquid?
- 2. What does the negative pressure mean?
- 3. Does a proton need negative and positive pressure region to maintain its stability?
- 4. What the pressure is acting on?

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Stress tensor in H atom

By non-relativistic reduction of the Dirac equation, one can construct the following EMT $T_{\rm QM}^{ij}$ which consists of a kinetic term

$$T_K^{ij} = -\frac{1}{4m} \left(\phi^{\dagger} \partial^i \partial^j \phi - \partial^i \phi^{\dagger} \partial^j \phi + \text{c.c.} \right) , \qquad (94)$$

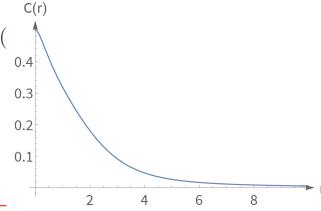
plus a potential term made of interacting electric fields of the proton and electron,

$$T_V^{ij} = \delta^{ij} \nabla V_p \cdot \nabla V_e - \partial^i V_e \partial^j V_p - \partial^i V_p \partial^j V_e \ .$$

The trace of $T_{\text{QM}}^{ij} = T_V^{ij} + T_K^{ij}$ can be calculated as

$$T_{\mathrm{QM}}^{ij}(\vec{r}) = (\delta^{ij}\nabla^2 - \nabla^i\nabla^j)\frac{C_{\mathrm{QM}}(r)}{M}$$

$$\frac{C_{\rm QM}(r)}{M} = \frac{1}{2\nabla^2} T_{\rm QM}^{ii} = \frac{e^{-2\alpha r} \alpha (2\alpha r + 1)}{16\pi r^2} - \frac{\alpha}{16\pi r^2}$$



Review on multipoles of electric current \vec{j}

• For static system, one has the current conservation

$$\partial_i J^i = 0$$

In momentum space (two D.O.F)

$$q_i J^i = 0$$

Thus, one has many vanishing moments,

$$\int d^3 \vec{r} \ r_{(i_1} ... r_{i_l} j_{i_l}(\vec{r}) = 0$$

Two independent series

Magnetic multipoles

$$\tilde{V}_{ii_{1}...i_{l}}^{(l)} \sim \int d^{3}\vec{r} \ m_{i}(\vec{r})r_{(i_{1}}...r_{i_{l-1})} \ ,$$

$$\vec{m}(\vec{r}) = \vec{r} \times \vec{j}(\vec{r}) \ ,$$

Most important: dipole moment or magnetic moment

Longitudinal multipoles

moments of $\vec{r} \cdot \vec{j}$

which does not contribute to static E&M multipoles

Multipoles of momentum current

Current conservation (only 3 DOF)

$$\partial_j T^{ij} = 0$$

One general identity

$$\frac{1}{k!} \sum_{P} \int d^3 \vec{r} T_{ii_{P(1)}} r_{i_{P(2)}} \dots r_{i_{P(k)}} = 0$$

or two vanishing series of moments

$$U_{iji_{1}...i_{l}}^{(l+2)} \equiv T_{(ij,i_{1}...i_{l})} ,$$

$$\tilde{U}_{iji_{1}...i_{l}}^{(l+1)} \equiv \frac{2l}{l+2} T_{i[j,i_{1}]...i_{l}} ,$$

Three non-vanishing moment series:

Scalar multipoles ("pressure" multipoles)

$$S(r)=T^{ii}(r)$$

$$S^{J}=\int d^{3}\vec{r} \, S(r) r_{r_{1}} \dots r_{i_{j}}$$
 (S⁽⁰⁾=0, however, scalar monopole density does not)

Tensor multipoles (natural parity, "shear pressure")

$$T^{J}$$
 from $\int d^{3}\vec{r} T_{ij}(\mathbf{r})\mathbf{r}_{i} \dots \mathbf{r}_{ij}$

Tensor multipoles (unnatural parity)

$$\tilde{T}^{J}$$
 from $\int d^{3}\vec{r} T_{i[j}(\mathbf{r})\mathbf{r}_{r_{1}]} \dots \mathbf{r}_{i_{1}}$

Form factors & multipoles: Scalar particle

Form factor

$$\begin{split} \langle P' | T^{\mu\nu} | P \rangle \\ &= 2 P^{\mu} P^{\nu} A(q^2) + 2 (q^{\mu} q^{\nu} - g^{\mu\nu} q^2) C(q^2) \ , \end{split}$$

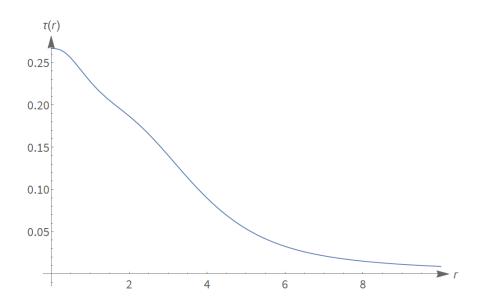
Tensor monopole T0 ("shear flow")

$$T^{(0)} = \frac{1}{5} \int d^3 \vec{r} T_{ij}(\vec{r}) \left(r_i r_j - \frac{\delta_{ij}}{3} r^2 \right)$$

normalization, $\tau = -T^{(0)}/2 = D/4M$ (the D term)

Tensor monopole density

$$\tau(r) = -\frac{2\pi}{5}r^2\left(r^ir^j - \frac{1}{3}r^2\delta^{ij}\right)T_{ij}(r)$$



when integrated over, one gets, $\frac{\hbar^2}{4m} \ (1 + O(\alpha))$

Tensor monopole moment τ

For a free boson

$$\tau_{\rm boson} = -\frac{\hbar^2}{4M}$$

- It has been argued that the stable system must have τ negative.
- However, for H-atom, we find

$$\tau = \hbar^2/4M (1+O(\alpha))$$

Stability of a system shall not depend on momentum current properties. It is quantum mechanics!

Gravitational field from form factor C

Linearized Einstein equation

$$\Box \bar{h}^{\mu\nu} = \frac{16\pi G}{c^4} T^{\mu\nu}$$
 where $g^{\mu\nu} = \eta^{\mu\nu} - h^{\mu\nu}$ and $\bar{h}^{\mu\nu} = h^{\mu\nu} - \frac{\eta^{\mu\nu}}{2} h^{\rho}_{\rho}$

The solution with C form factor is

$$h_C^{00}(\vec{r}) = -\frac{8\pi G}{c^4 M} C(r)$$

$$h_C^{ij}(\vec{r}) = \frac{8\pi G}{c^4 M} C(r) \delta^{ij}$$

• Given C(r) decays exponentially, so does the metric perturbation.

Spin-1/2 particle

Form factors

$$\langle P' | T^{\mu\nu} | P \rangle = \bar{u} \left(P' S' \right) \left[A(q^2) \gamma^{(\mu} \bar{P}^{\nu)} + B(q^2) \frac{\bar{P}^{(\mu} i \sigma^{\nu)\alpha} q_{\alpha}}{2M} + C(q^2) \frac{q^{\mu} q^{\nu} - g^{\mu\nu} q^2}{M} \right] u(PS) ,$$

Apart from the angular momentum multipoles,
 one has the tensor monopole

$$\tau = \frac{C(0)}{M} \ .$$

C(0) shall be negative for the nucleon from a different reason.

Spin-1 particle

Six form factors

$$\begin{split} &\langle P', \epsilon_f | T^{\mu\nu}(0) | P, \epsilon_i \rangle \\ &= -2 \bar{P}^{\mu} \bar{P}^{\nu} \left[(\epsilon_f^{\star} \cdot \epsilon_i) A(q^2) + E^{\alpha\beta} q_{\alpha} q_{\beta} \frac{\tilde{A}(q^2)}{M^2} \right] \\ &+ J(q^2) \frac{i \bar{P}^{(\mu} S^{\nu)\alpha} q_{\alpha}}{M} \\ &- 2 (q_{\mu} q_{\nu} - g_{\mu\nu} q^2) \left[(\epsilon_f^{\star} \cdot \epsilon_i) C(q^2) + E^{\alpha\beta} q_{\alpha} q_{\beta} \frac{\tilde{C}(q^2)}{M^2} \right] \\ &- \left[(E^{\mu\nu} q^2 - E^{\mu\alpha} q^{\nu} q_{\alpha} - E^{\alpha\nu} q^{\mu} q_{\alpha} + g^{\mu\nu} E^{\alpha\beta} q_{\alpha} q_{\beta} \right] D(q^2) \end{split}$$

 Mass quadrupole, tensor quadrupole, and scalar quadrupole:

$$T_{ij}^{(2)} = -\frac{\tilde{C}(0)}{48M^2}\hat{E}_{ij}$$
 $\sigma_{ij} = \frac{D(q^2 = 0)}{M}\hat{E}_{ij}$

Conclusions

- The mass sum rule with physical definition and symmetry is unique.
- The anomaly or scalar contribution to the mass of the proton acts like Higgs mechanism. It is important to measure mass and scalar radius
- The form factor C can best be characterized by gravtional tensor monopole density which generates specific type of gravity. (pressure and stability are not natural concepts in this context).

Questions

- 1. Is the pressure the same as we understand in a gas or liquid?
- 2. What does the negative pressure mean?
- 3. Does a proton need negative and positive pressure region to maintain its stability?
- 4. What the pressure is acting on?

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What is the (usual) pressure?

Pressures for non-interacting systems

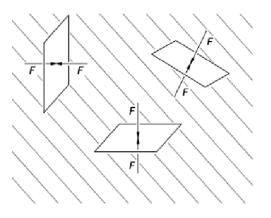
Average kinetic energy either in thermal or quantum state

Pressure in ideal gas

$$P = nk_BT = 2/3 K.E.$$

Pressure in a quantum fermi gas
 P = 2/3 K.E.

- Non-directional (locally in equilibrium)
- Pressure is always non-negative unless in unstable phase.



Pressure in a system with interactions : $P = -\frac{\partial F}{\partial V}$

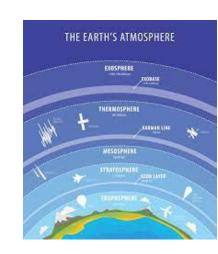


- Consider a small test volume δV in thermal equilibrium $(M.F.P. \ll (\delta V)^{\frac{1}{3}})$, the interaction range λ
- Short-range interaction, $\lambda \ll (\delta V)^{1/3}$

repulsive interaction: pt

attractive interaction: p.

- Long-range interaction, $\lambda \gg (\delta V)^{1/3}$
 - The interaction is not part of the pressure.
 - For a confining system (atmosphere on earth) the pressure goes to zero!



Energy-momentum tensor and pressure

 What has been calculated or measured is related to the energy-momentum tensor in space

$$T^{\mu\nu}(\vec{r}) \quad (\mu, \nu = 0,1,2,3)$$

It is well-known that in the ideal gas/fluid model

where p is the normal pressure in the gas. What has been measured is the analogue:

$$p = \frac{1}{3} (T^{11} + T^{22} + T^{33}).$$

This is not the usual pressure

- This pressure does not follow from the usual definition, but from an analogy that the nucleon is some sort of fluid of quarks and gluons.
- There is nothing wrong with a definition, but one needs to be careful (cannot be literal) when interpreting it.
- This pressure takes into account long-range confinement interactions, unlike the example of atmosphere on the earth: It is not just the kinetic energy of quarks and gluons.

The meaning of p(r)

• p =
$$\frac{1}{3}$$
 (T¹¹+T²²+T³³)

is the average of

the momentum flow (or current) in the x-direction for the momentum component-x, which can be positive or negative

with

the momentum flow (or current) in the y-direction for the momentum component y, which can be positive or negative

and

the momentum flow (or current) in the z-direction for the momentum component z, which can be positive or negative

Answer for Q2: negative pressure?

- Since the momentum flow pattern in a system can be rather arbitrary (no particular constraint), there is no reason that the "pressure" defined as such be positive definite.
- It is just a definition ©
- Note, however, that in real fluid $T^{ii} > 0$ everywhere.

Answer for Q3: stability

Laue condition for stability

$$\int d^3\vec{r} \, p(r) = 0$$

p(r) must be positive and negative.

• However, this follows directly from current conservation $\partial_i T^{ij} = 0$:

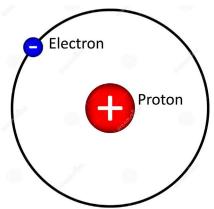
$$\int d^3\vec{r} T^{ij} = 0$$

$$\int d^3\vec{r} r^{(k}r^l \dots r^m T^{ij)} = 0$$

 Stability of a quantum system is guaranteed by quantum mechanics (no classical equivalent)

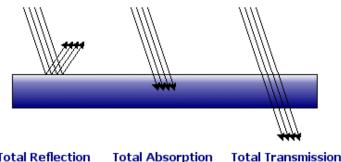
Q4: What is the pressure acting on?

- Consider H-atom, with one electron there is a momentum current flow, the pressure is certainly not on the electron.
- T^{ij} is about the motion pattern, not acting on other parts of the system
- One imagines a fictious surface which intersects the momentum current.
 - Normal definition of the pressure assumes it bounces
 - Here one assumes the current gets absorbed (Q=1)



Radiation Pressure Coefficient

© Blaze Labs Research



Total Reflection Q_{DD}=2

 $Q_{pp} = 1$

otal Transmissio Q_{pp}=0

My current understanding

- The pressure or shear pressure are introduced to characterize a momentum flow pattern.
- One shall not over-interpret them literally (mechanical stability etc.)

Why nucleon mass?

Only the nucleon can salvage a part of the energy from the Big Bang (the energy capsule!) and to be harnessed later to lit up the dark universe.

